

COLLAGE 2023

Lecture 09: Ionosphere Observation with Radio Waves

Yang Wang

Smead Department of Aerospace Engineering Sciences

University of Colorado Boulder

Email: yang.wang-2@colorado.edu

Source materials partially from Jade Morton and Robert Marshall



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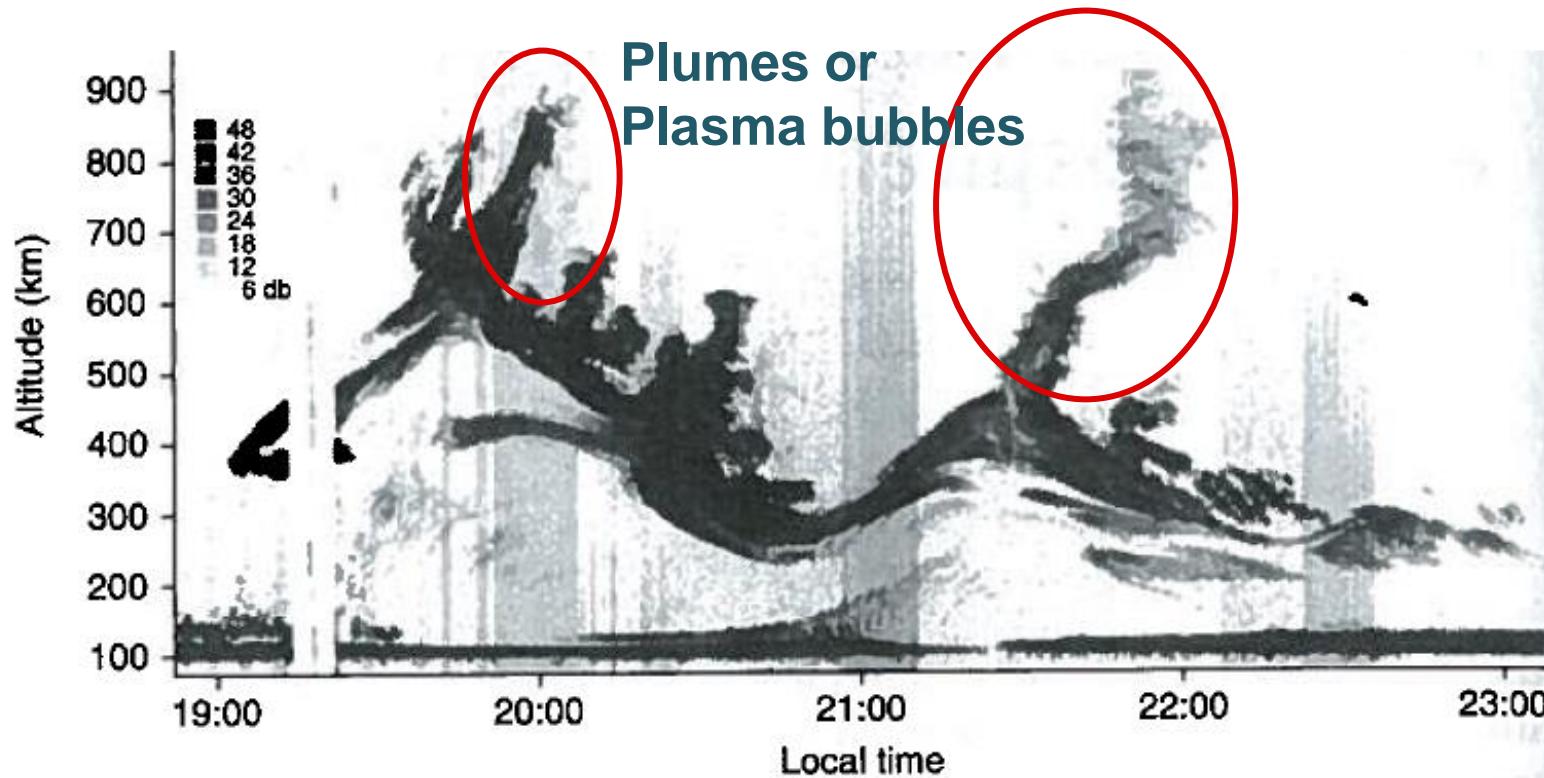
COLLAGE 2023 / ASEN-5519: Space Weather Overview

Outline

- **Overview of ionosphere observation approaches**
- **GNSS network-based ionosphere observation and monitoring**
- **GNSS space-based ionosphere observations**



Ionosphere Irregularities: Temporal Evolution



Jicamarca, Peru, vertical backscatter at 3m
3/21/1979

Kelley, M. C., Larsen, M. F., LaHoz, C.A., and McClure, J. P., "Gravity wave of equatorial spread F: A case study," J. Geophys. Res., 86, p9087, 1981.

- Electron Density
- Ion and Electron Temperature
- Ion Drift
- Ionospheric Composition

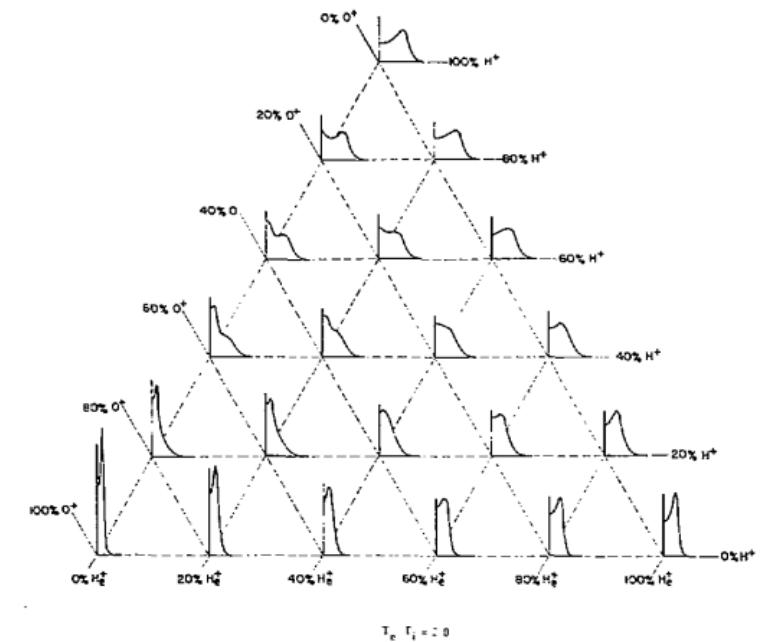


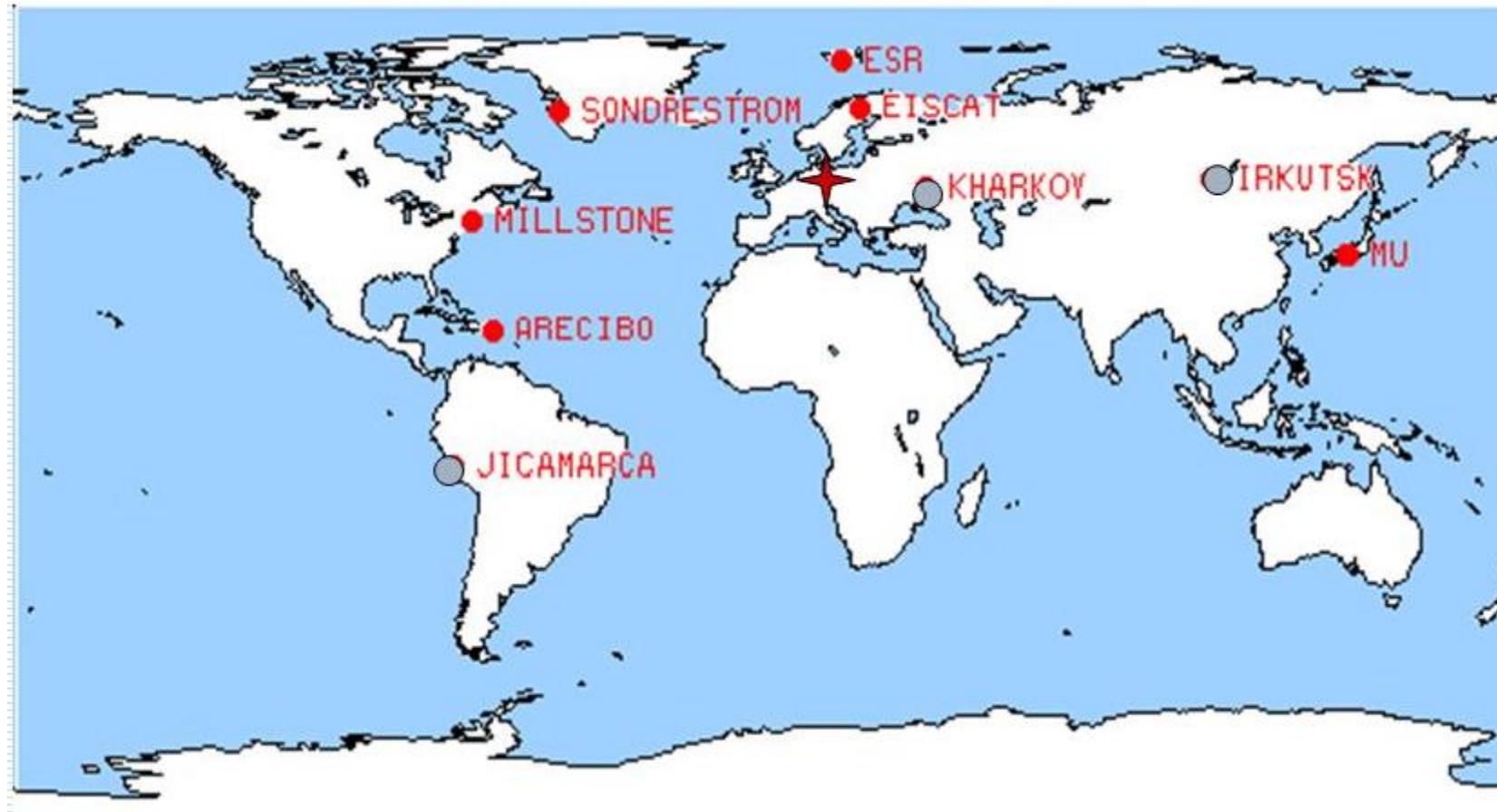
Fig. 2—Spectra for mixtures of O^+ , H_3^+ , and H^+ .

Gordon, 1964



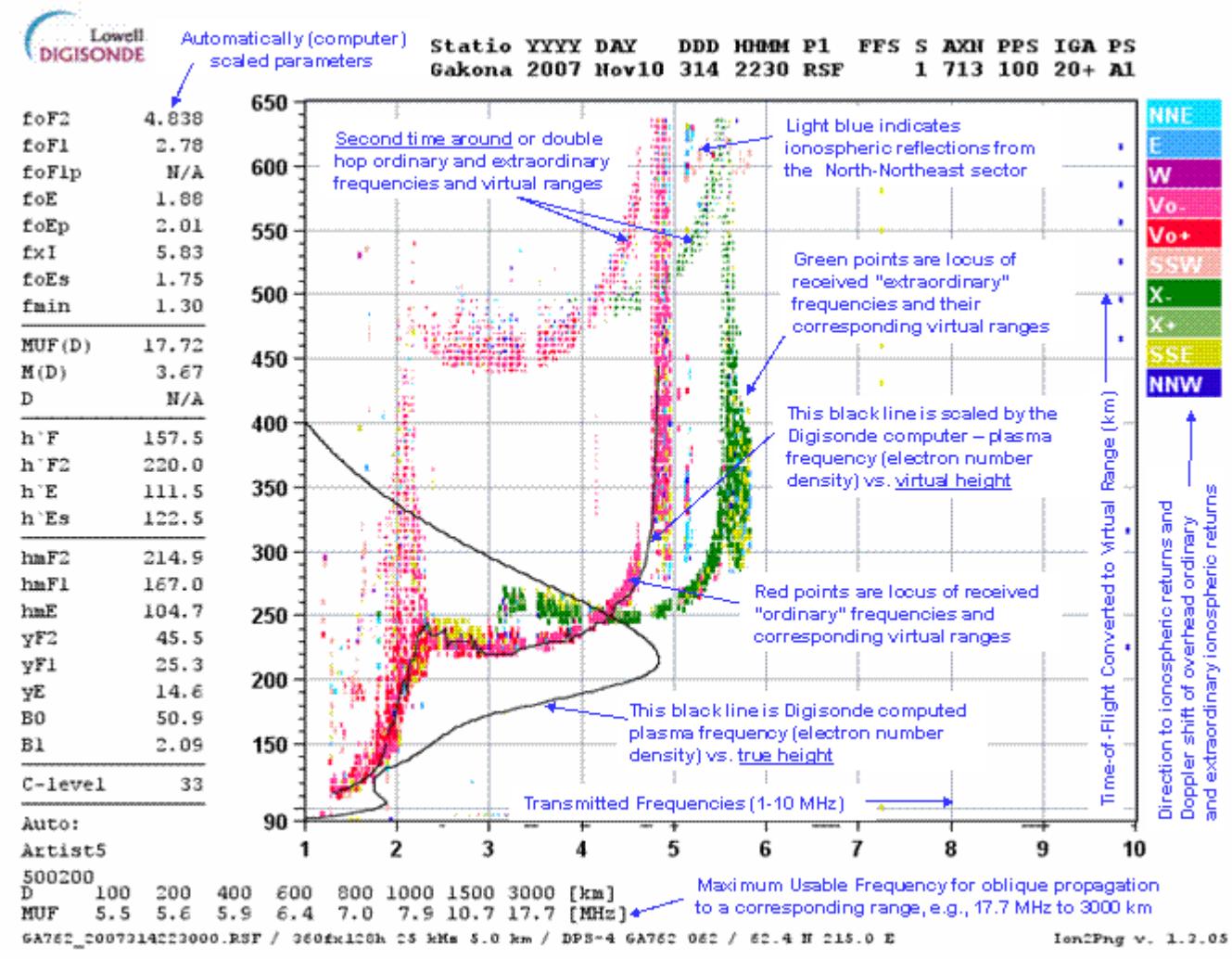
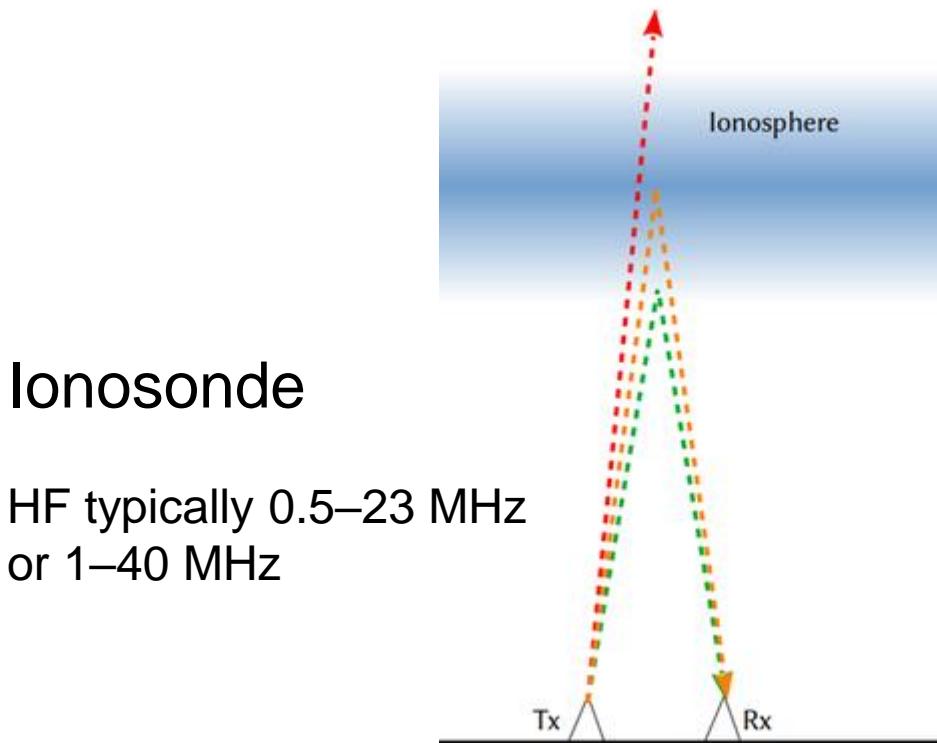
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World incoherent scatter radars



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Ionosphere Irregularity: Vertical Structure

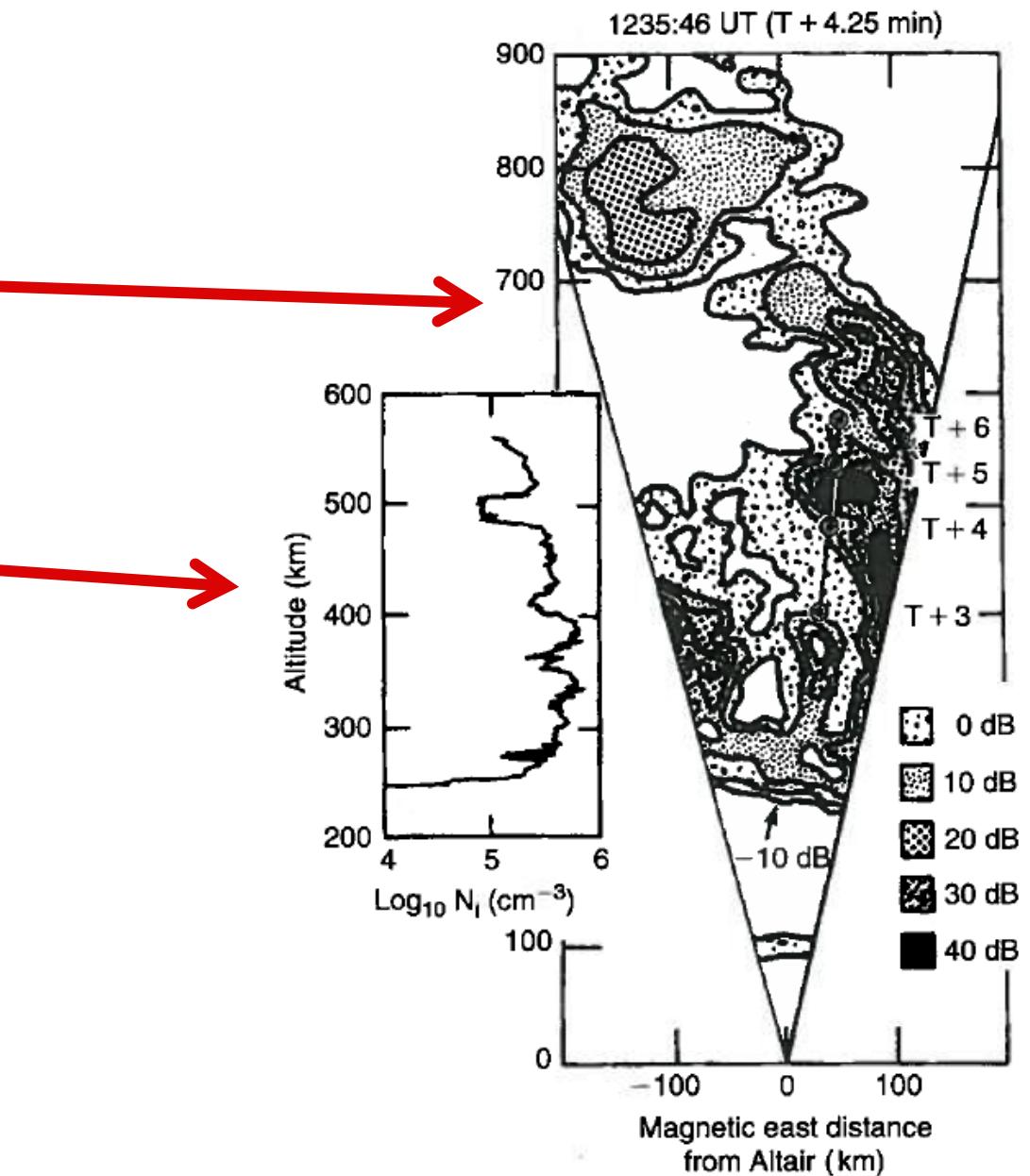


Ionosphere Irregularity: Vertical Structure

Altair radar on
Kwajalein island
0.96m wavelength

PLUMEX I rocket

Rino, Tsunoda, Petriceks, Livingston, Kelley,
Baker, "Simultaneous rocket-borne beacon
and in situ measurements of equatorial spread
F – intermediate wavelength results," J.
Geophys. Res., 86(4), p2411, 1981.

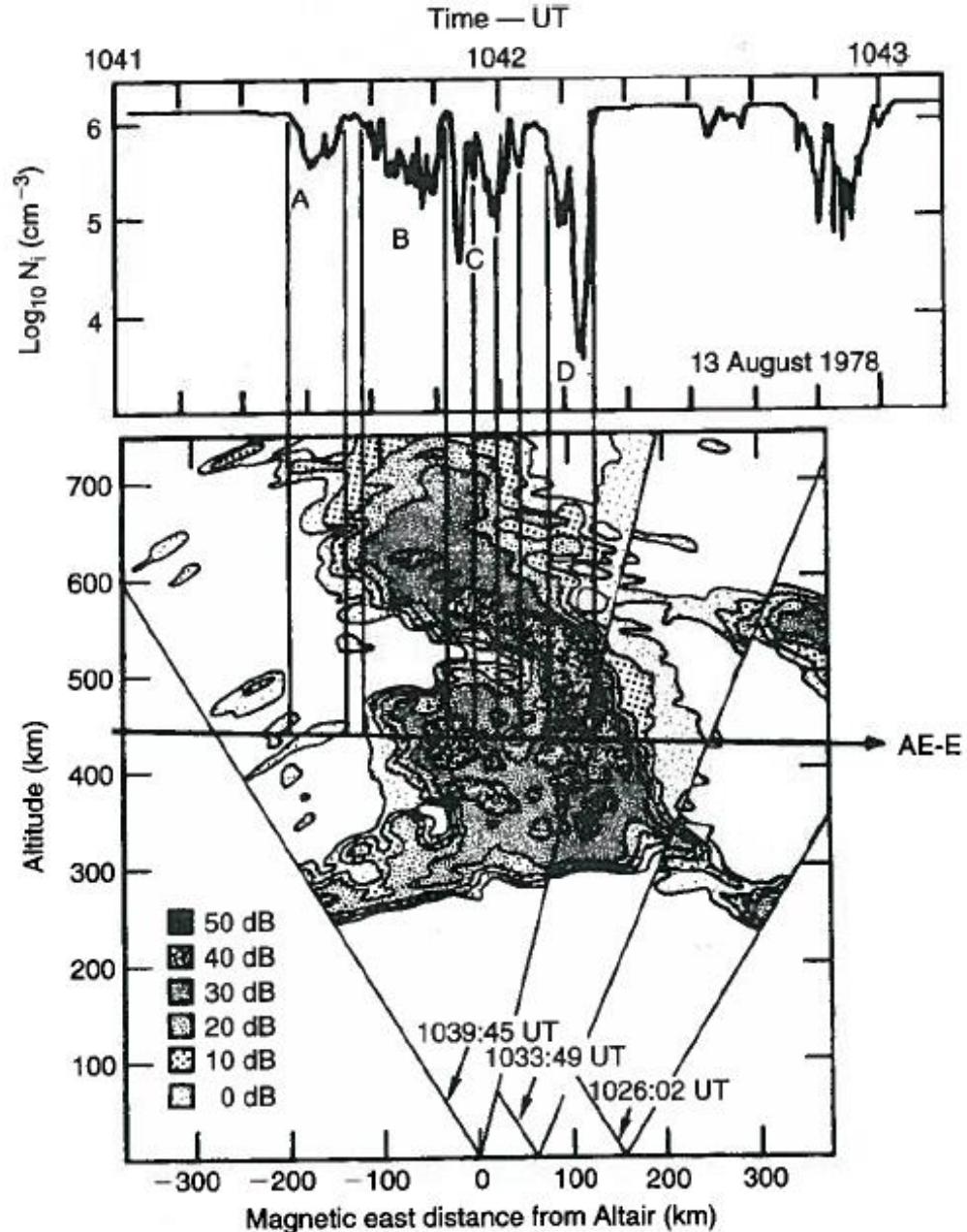


Ionosphere Irregularity: Horizontal Structure

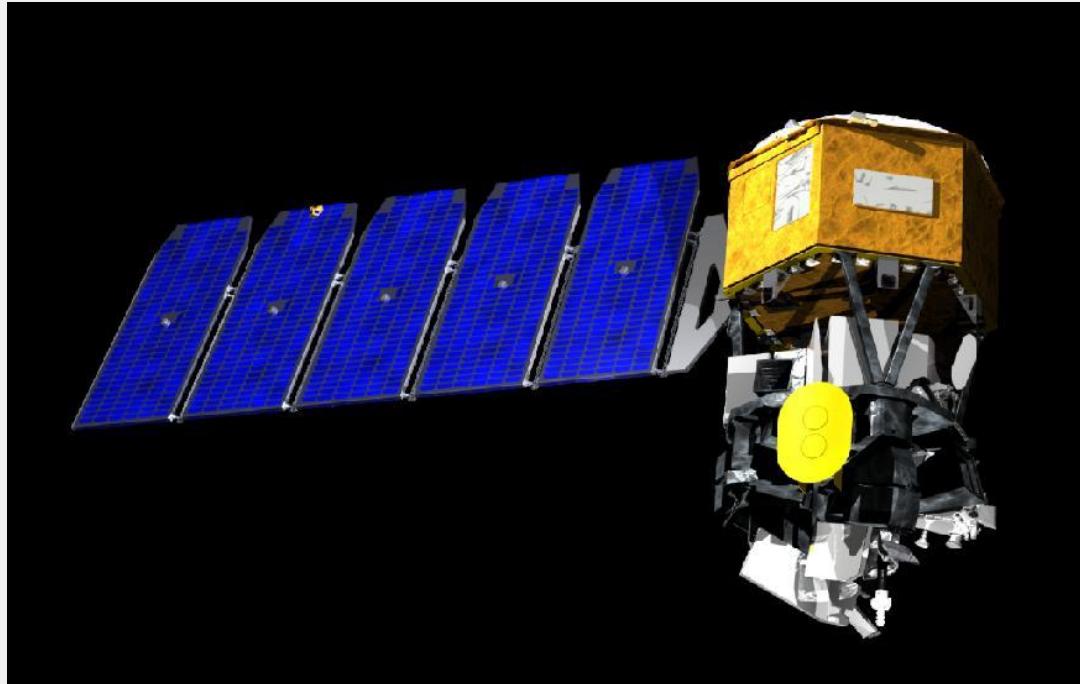
AE-E satellite 

Altair radar 

Tsunoda, Livingtons, McClure, Hanson,
"Equatorial plasma bubbles: vertically
elongated wedges from the bottomside F
layer," J. Geophys. Res., 87, p9171, 1982.



NASA's Ionospheric Connection Explorer (ICON)



- ❖ Understand drivers of ionospheric variability
- ❖ Explain how energy / momentum from lower atmosphere reach the space environment
- ❖ Explain how drivers create extreme conditions observed during solar-driven geomagnetic storms

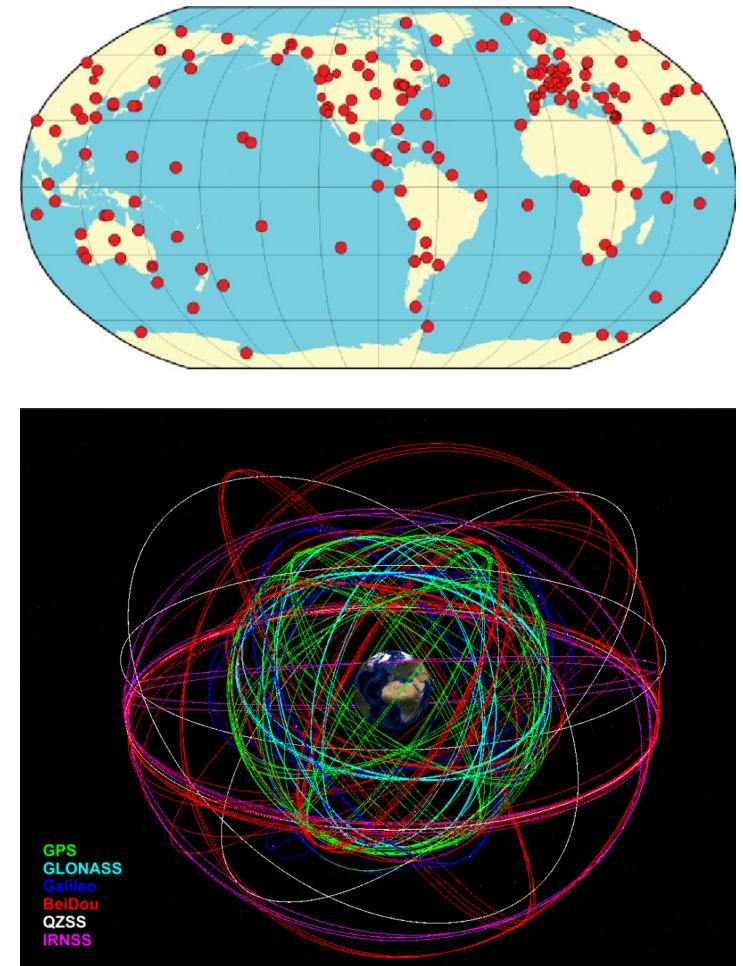
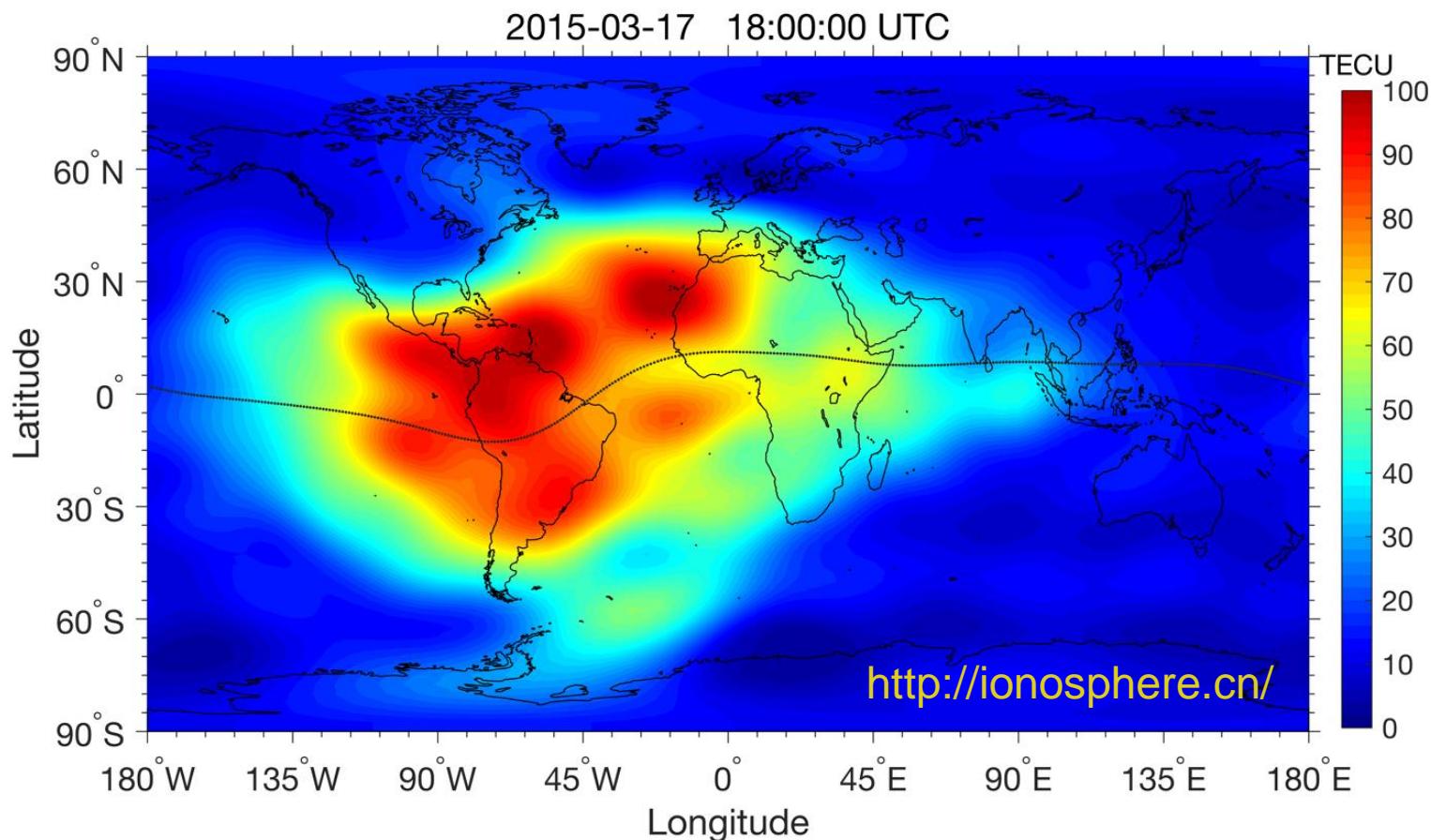
- ❖ Main instruments:
 - ❖ **MIGHTI** is a Michelson Interferometer to measure winds and temperatures
 - ❖ **FUV** is an FUV imager; observes UV emissions of N₂ and O to determine O/N₂ ratio
 - ❖ **EUV** images 83.4 nm emission from O; resonantly scattered by O+: gives ion density
 - ❖ **IVM** is the ion velocity meter; uses a Retarded Potential Analyzer (RPA) to measure relative velocity of ions, therefore winds, as well as temperature and density



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GNSS Networks

IGS Global Ionosphere Map (GIM)



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Outline

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Simplified Appleton-Hartree Equation

$$f_g \approx 1 \text{ MHz} \quad f_p \leq 10 \text{ MHz}$$

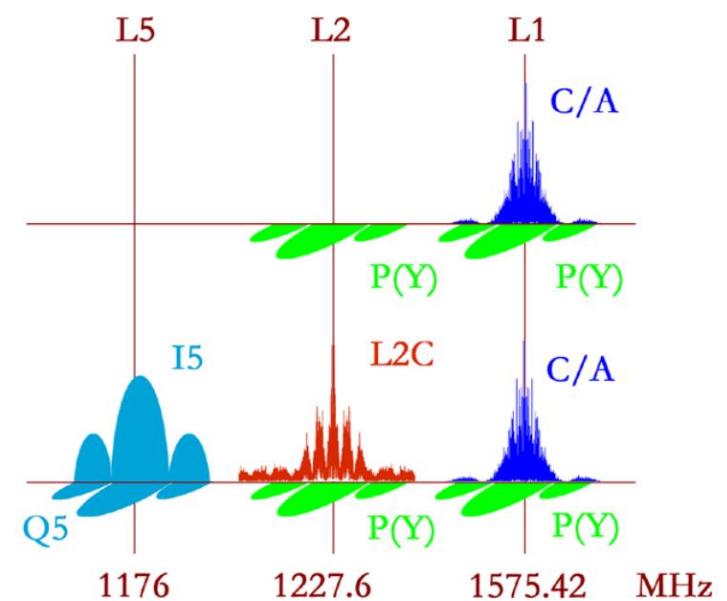
At GPS frequency, $f \sim \text{GHz}$: $X = \left(\frac{f_p}{f} \right)^2 \ll 1$ $Y = \frac{f_g}{f} \ll 1$

$$n_\phi = 1 - \frac{X}{1 - \frac{Y^2 \sin^2 \theta_B}{2(1-X)} \pm \sqrt{\frac{Y^4 \sin^4 \theta_B}{4(1-X)^2} + Y^2 \cos^2 \theta_B}}$$

$$n_\phi \approx 1 - \frac{X}{2} \pm XY|\cos \theta_B| - \frac{1}{4} X \left(\frac{X}{2} + Y^2(1 + \cos^2 \theta_B) \right)$$

Vacuum 1st order 2nd order 3rd order

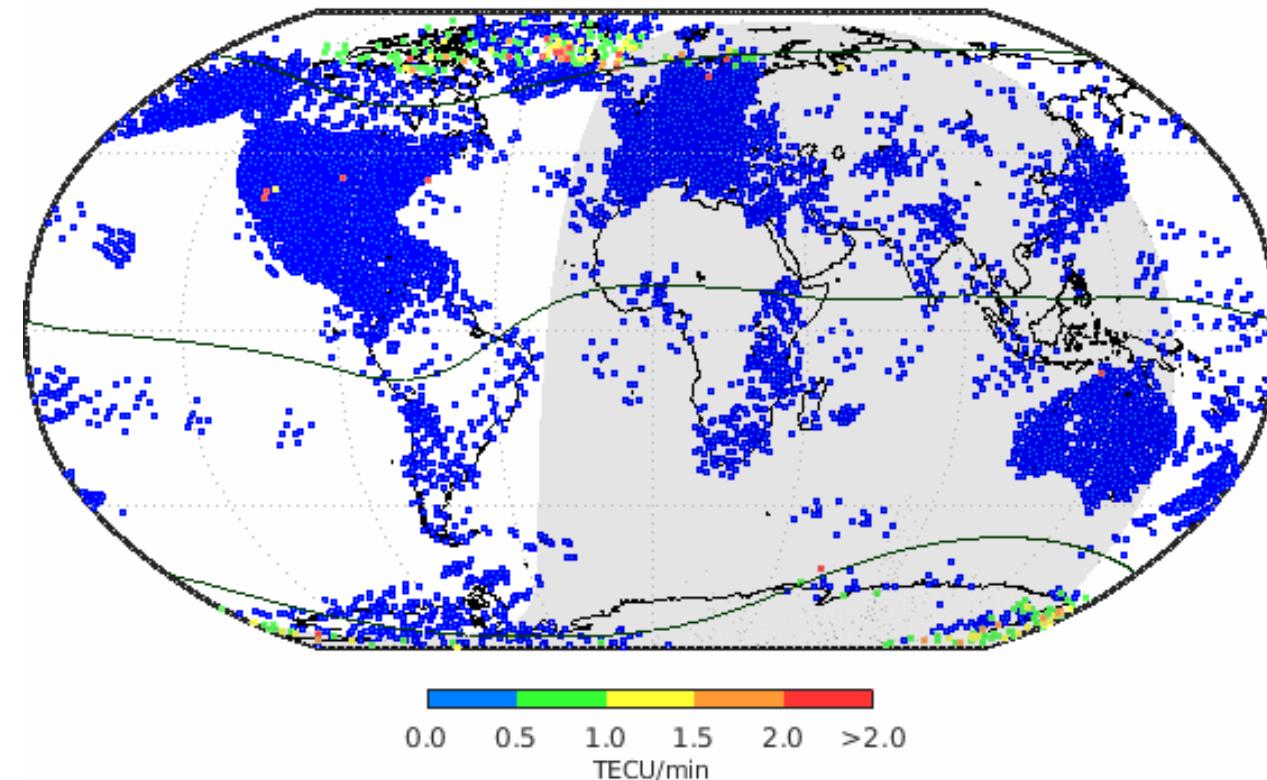
$$n_j \gg 1 - \frac{X}{2} \gg 1 - \frac{e^2}{2me_0 f^2} N_e = 1 - 40.3 \frac{N_e}{f^2}$$



Ionosphere Disturbance Impact on Mid-latitudes

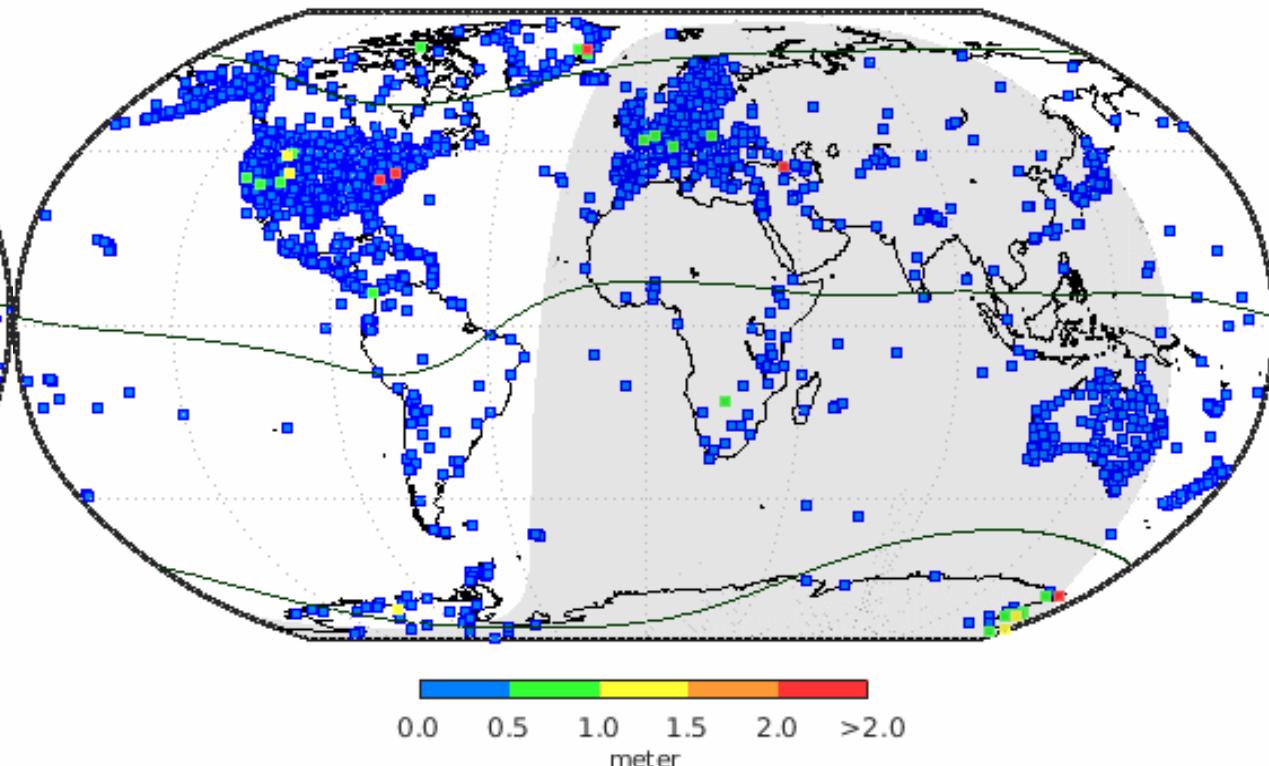
ROTI: Ionosphere Plasma Disturbance

09/07/2017 20:00 UT



Position errors

09/07/2017 20:00 UT



Yang, Z., S. Mrak, Y. Morton, "Geomagnetic storm induced mid-latitude ionospheric plasma irregularities and their implications for GPS positioning over North America: a case study," *Proc. IEEE/ION PLANS meeting*, 2020.



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Challenges in Measuring Ionospheric Irregularities

1. Availability

*Receivers cease to function if GNSS signal traverse irregularities
→ Data are not available when needed most!*

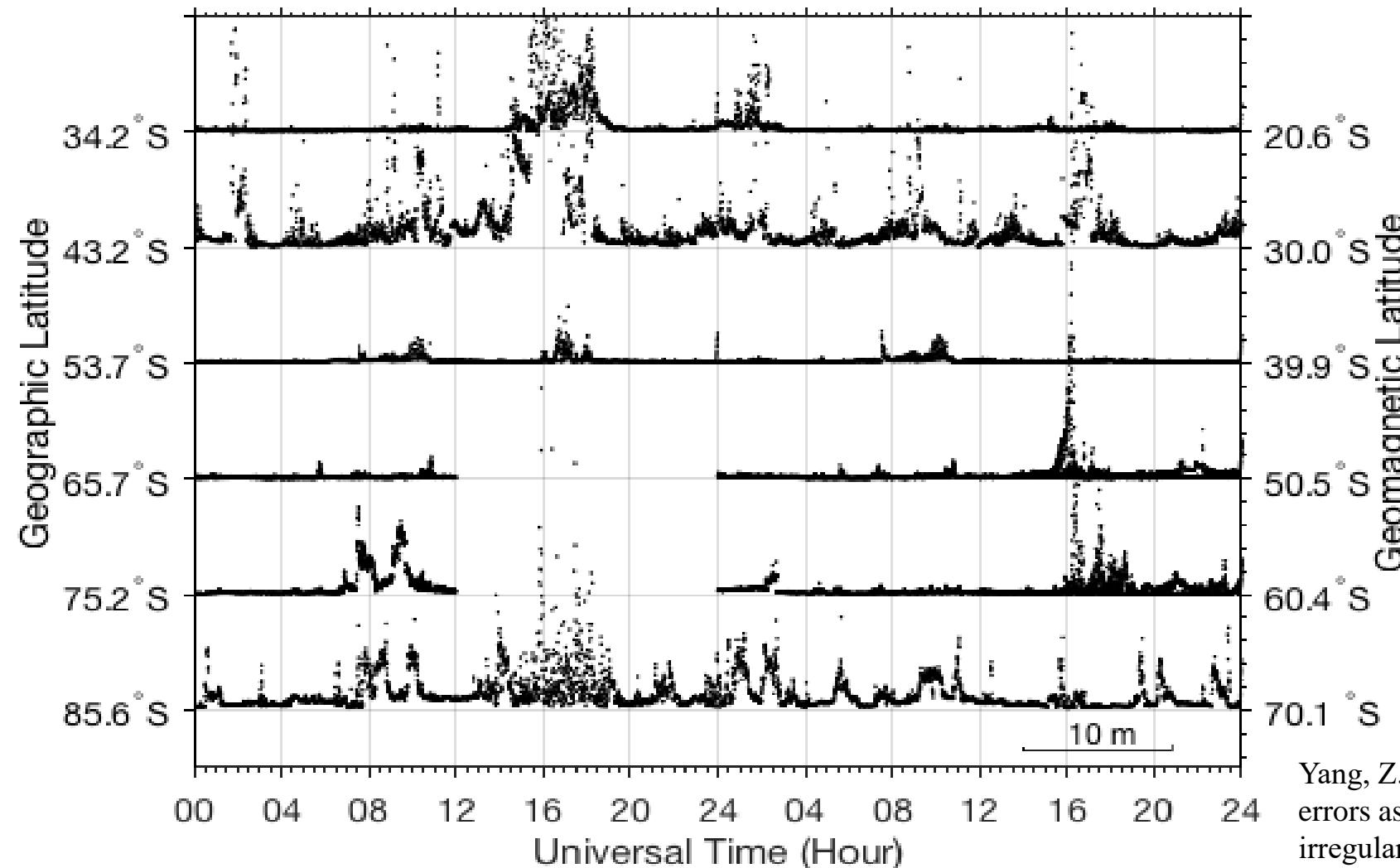
2. Accuracy

$(\text{Iono} + \text{other}) \times h(t) = \text{Observed Effects}$

Iono effects \neq Observed Effects



Availability Issue: March 17-18, 2015 St. Patrick's Day storm



Yang, Z., Y. Morton, "Kinematic PPP errors associated with ionospheric plasma irregularities during the 2015 St. Patrick's day storm," *Proc. ION GNSS+*, 2019.



Accuracy: Scintillation Indices

Phase scintillation index: $\sigma_\phi = \text{std}(\text{detrend}(\phi_{s,k}))$

Amplitude scintillation index: $S_4 = \sqrt{\frac{\langle SI^2 \rangle - \langle SI \rangle^2}{\langle SI \rangle^2}}$

Signal intensity (power): $SI = \frac{SI_{raw}}{SI_{trend}}$ $SI_{raw} = NBP - WBP$

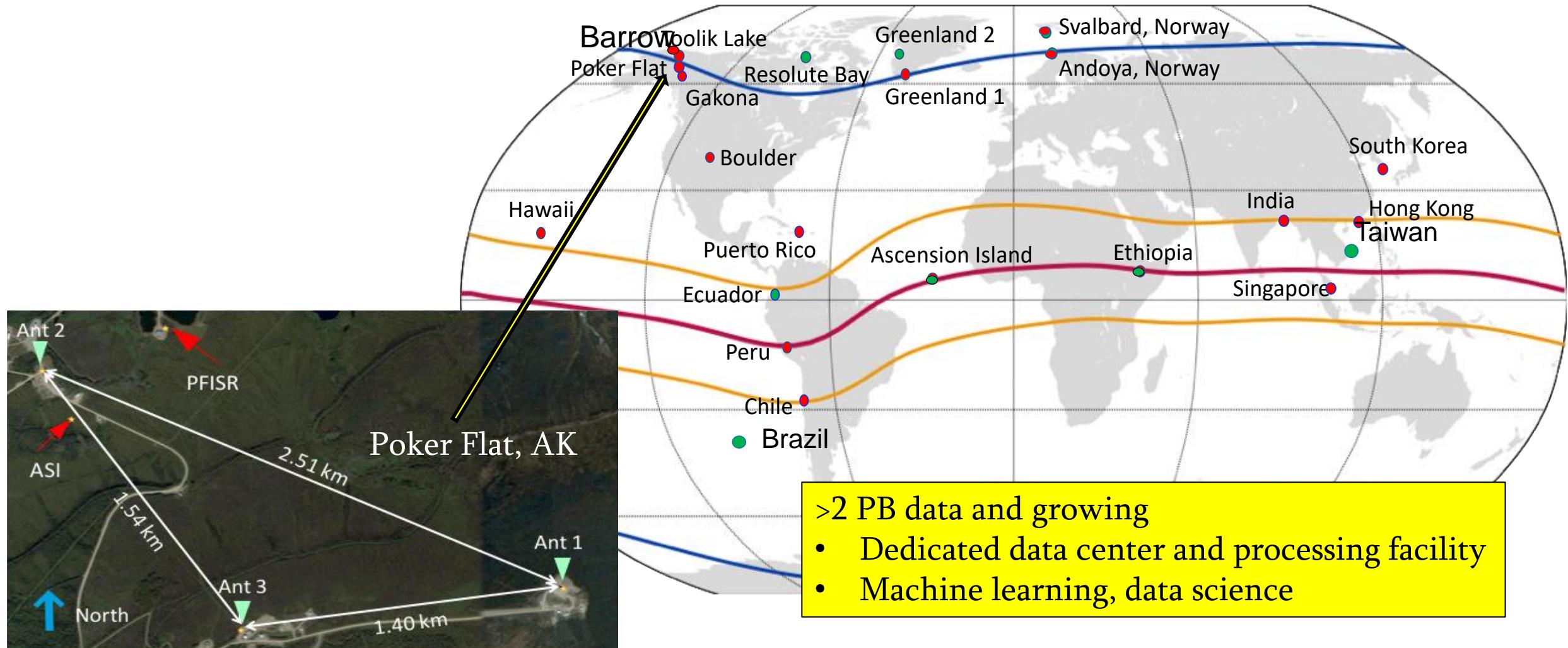
Narrowband power: $NBP = \left(\sum_{i=1}^M I_i \right)^2 + \left(\sum_{i=1}^M Q_i \right)^2$ M: number of correlation blocks over a selected period

Wideband power: $WBP = \sum_{i=1}^M I_i^2 + \sum_{i=1}^M Q_i^2$ Typical setting: $T_I = 1\text{ms} \rightarrow M=20; T_I = 10\text{ms} \rightarrow M=2$

Rate of TEC Index (ROTI): $ROTI = \sqrt{E \left[\frac{|TEC(t + \delta t) - TEC(t)|^2}{\delta t^2} \right]}$



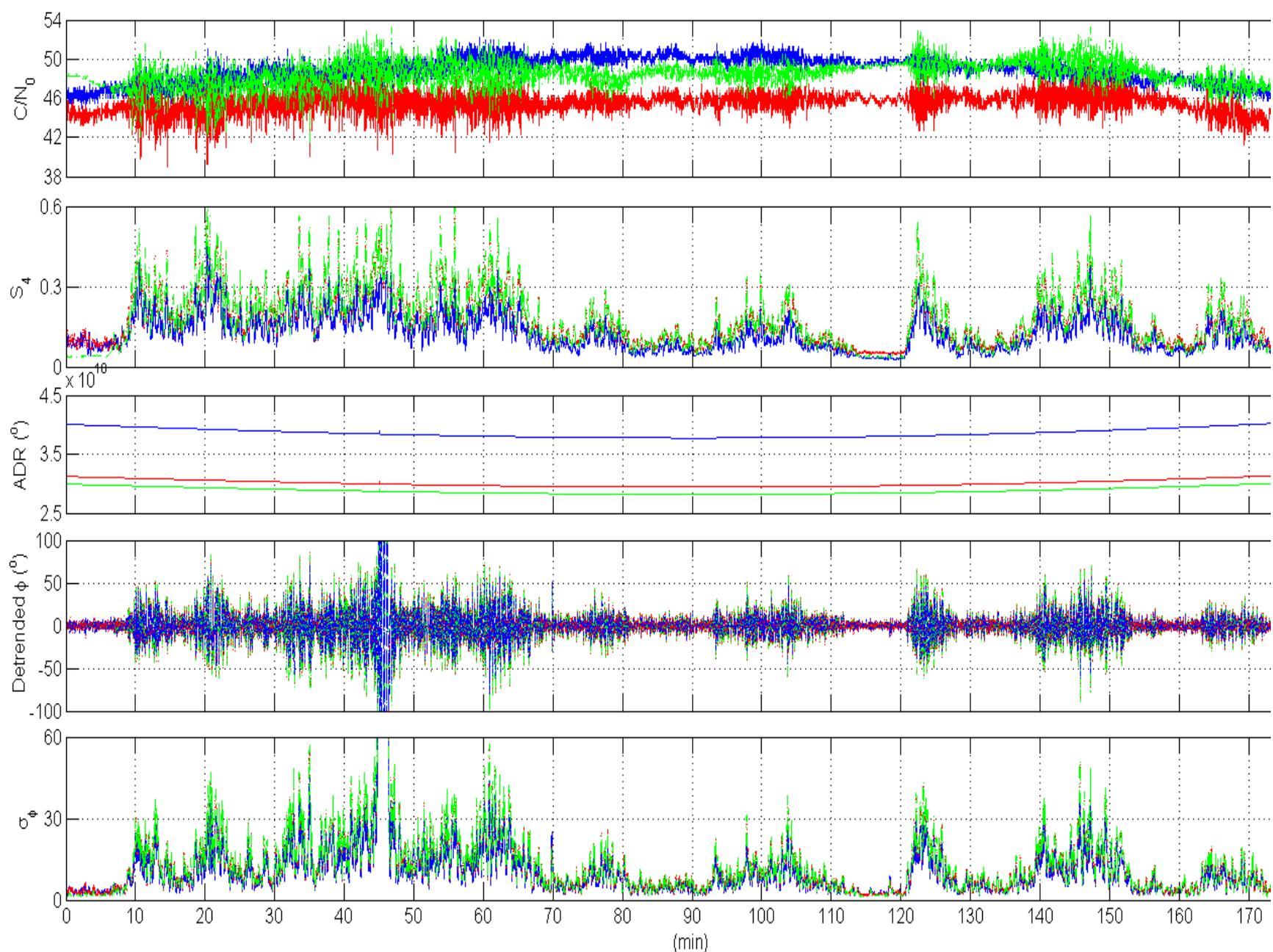
Global SDR Data Collection Network



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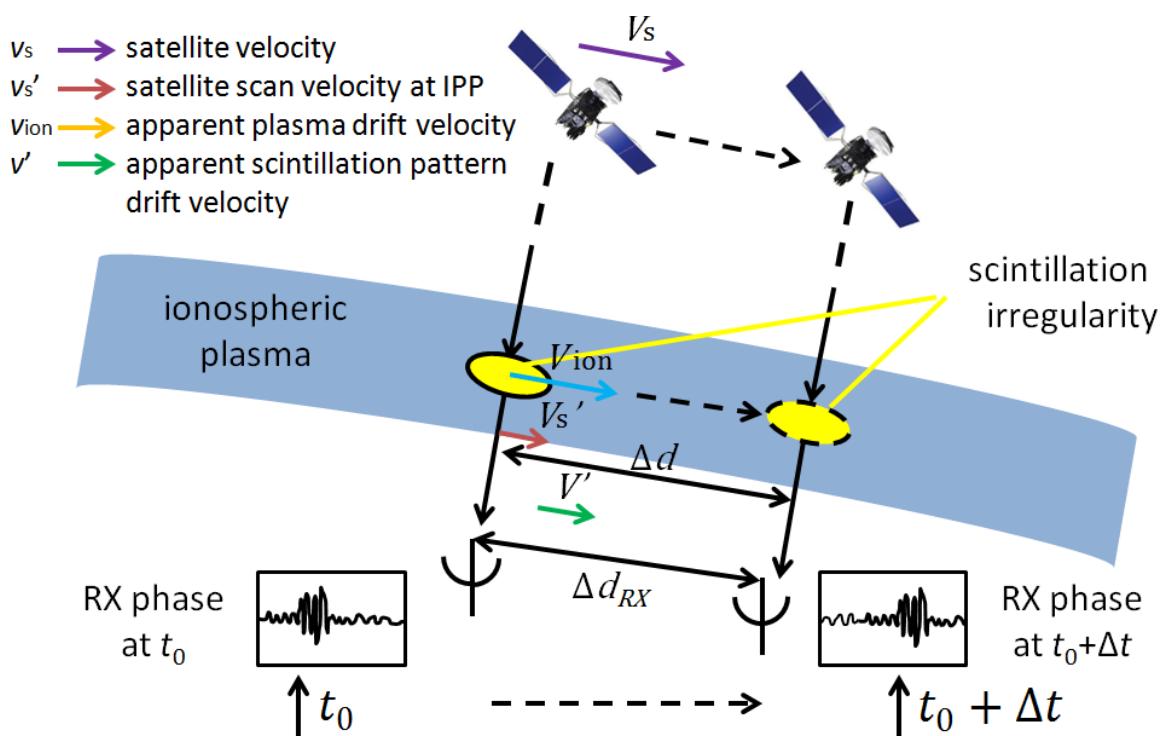
Low Latitude Scintillation Example

Peru
3/11/2013
13:30UTC



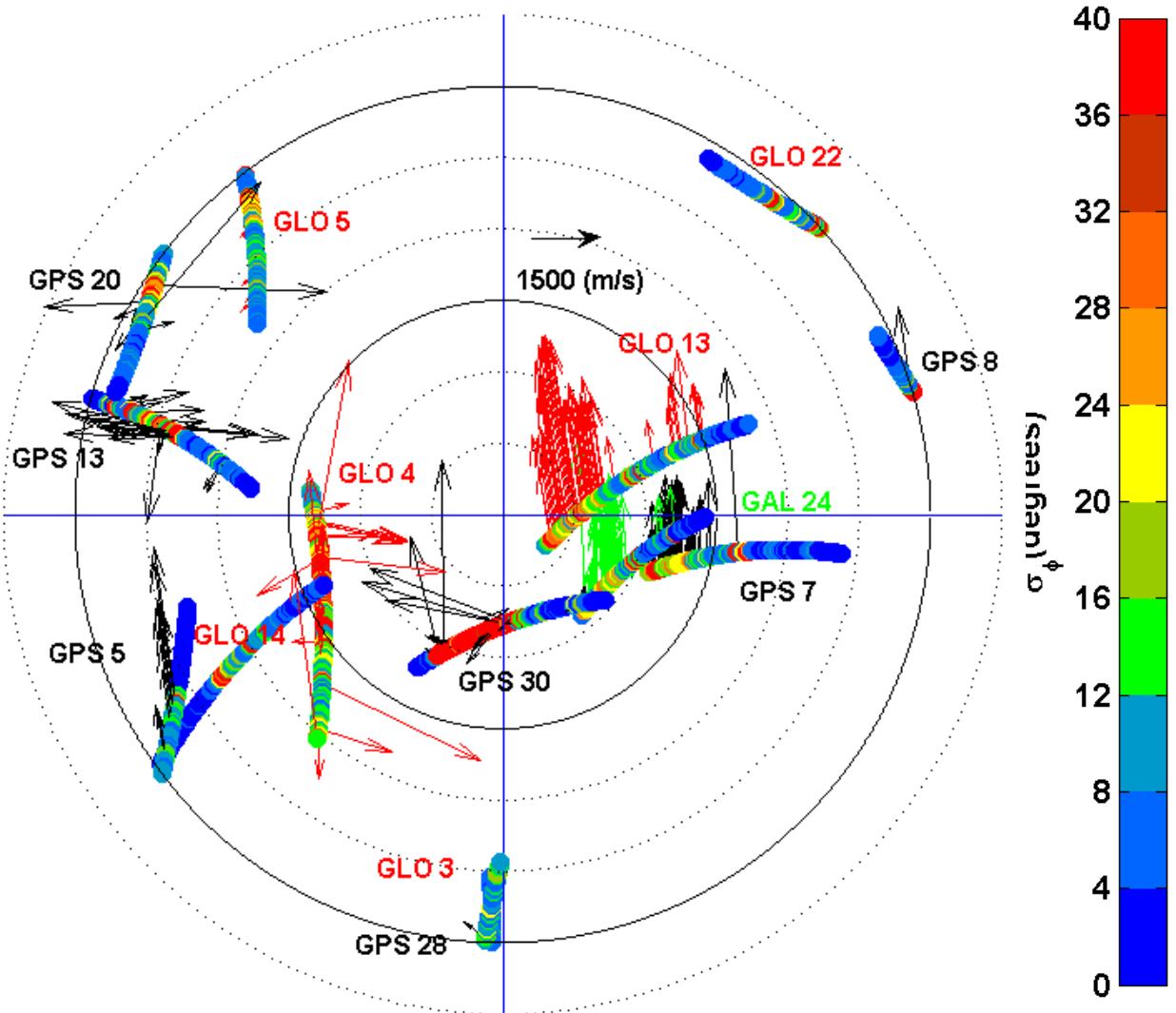
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Plasma Velocity Estimation



Skyplot of GNSS satellite tracks

2015/12/20 04:40:00



Wang, J., Y. Morton, "Ionospheric irregularity drift velocity estimation using multi-GNSS spaced-receiver array during high latitude phase scintillation," *Radio Sci.*, DOI: 10.1002/2017RS006470, 2018.



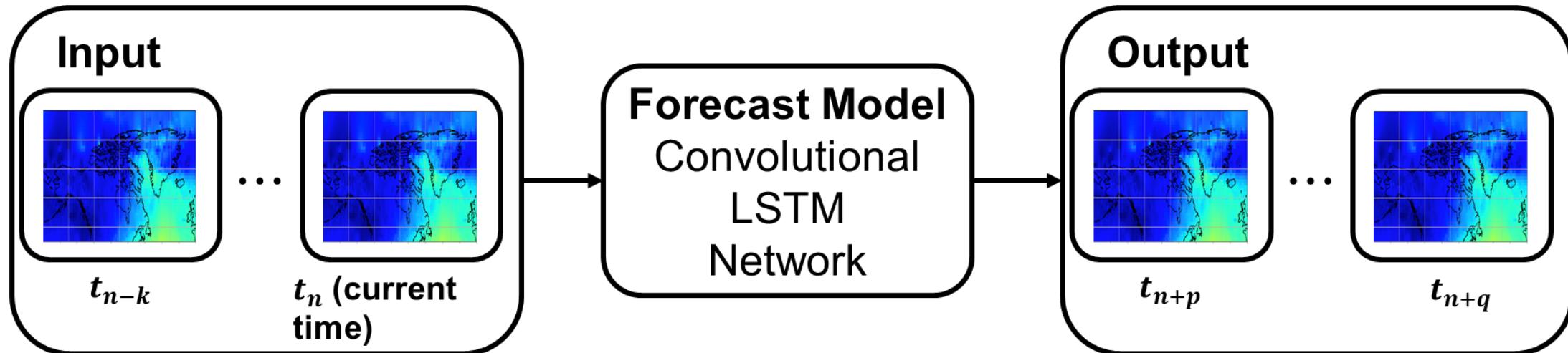
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Ionosphere TEC and Disturbance Forecasting

(Shi et al., 2015)

Machine Learning Forecast Framework Using ConvLSTM:

(Convolutional Long Short-Term Memory)



Input/Output

TEC Map: Background Ionosphere

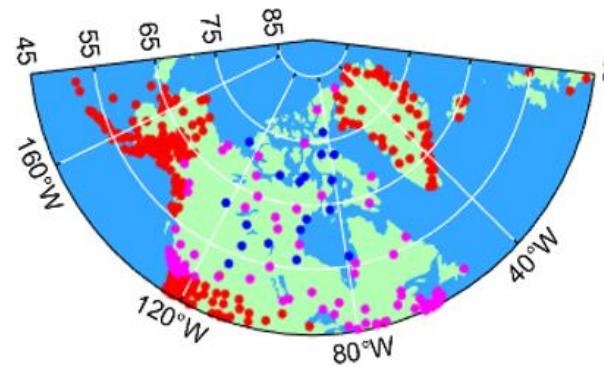
ROTI Map: Ionosphere Disturbances



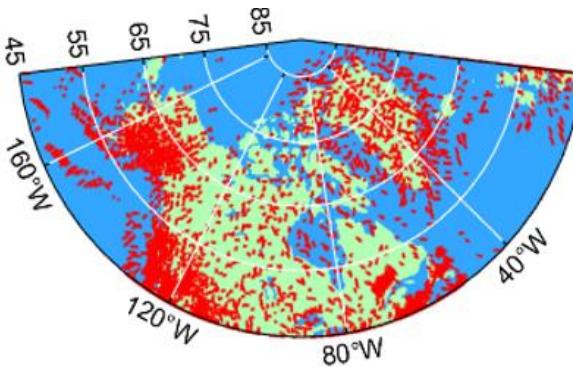
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Ionosphere Disturbance Forecasting with Ground GNSS Networks

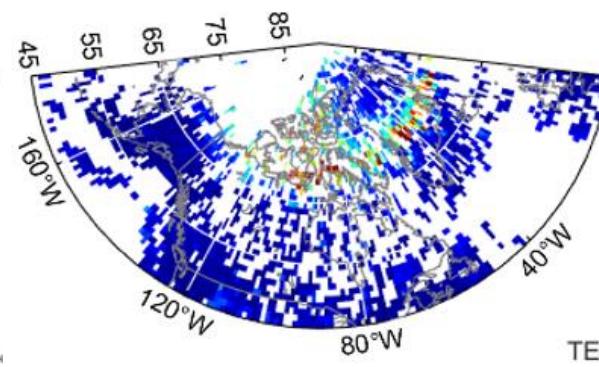
GNSS Receivers



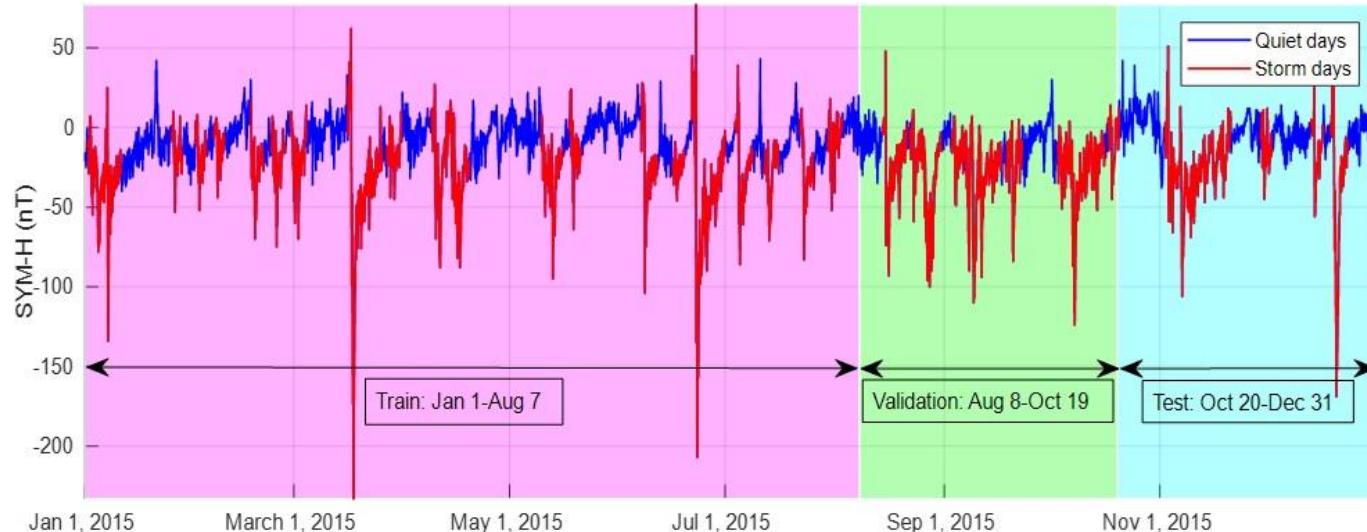
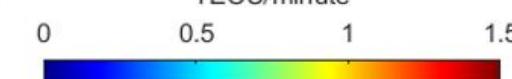
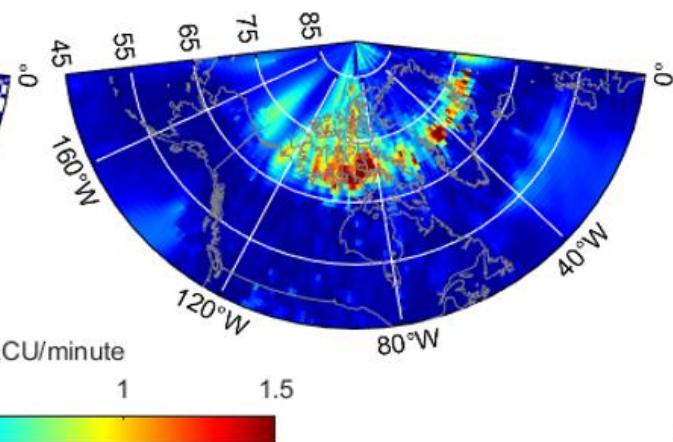
IPP



Raw ROTI



Interpreted ROTI



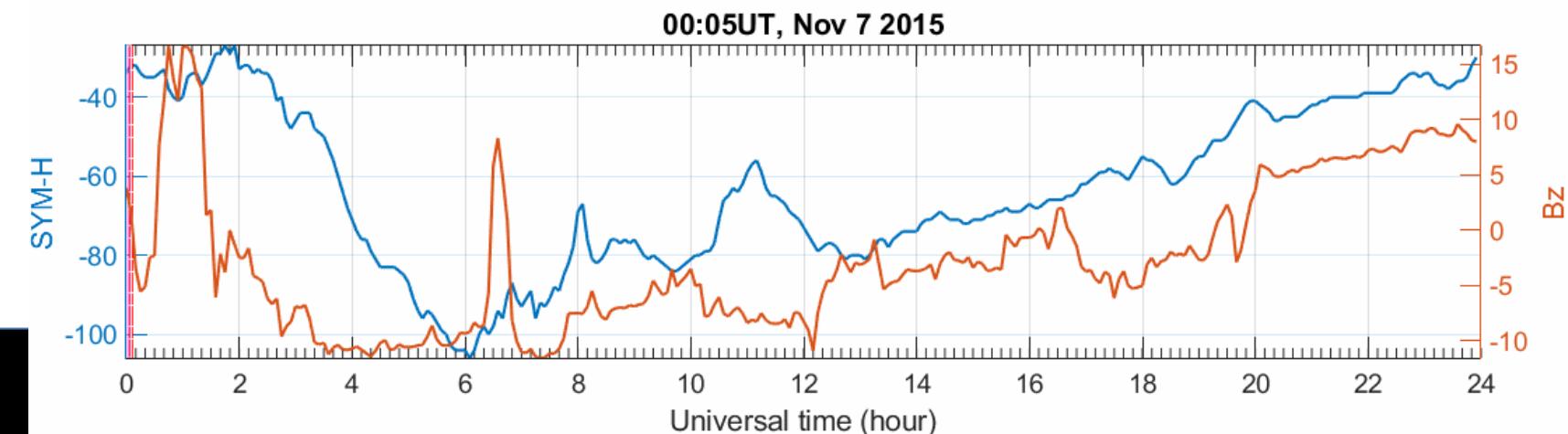
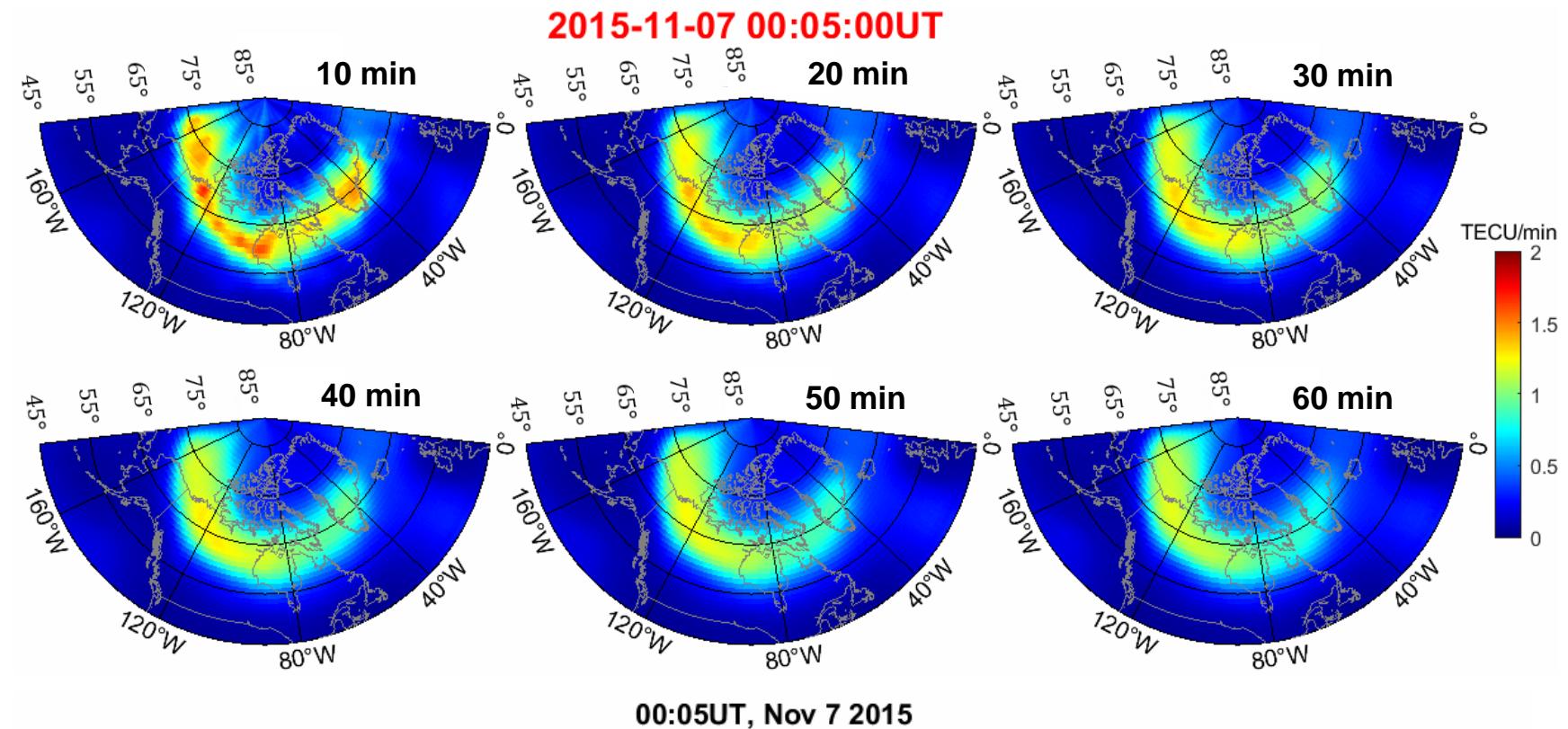
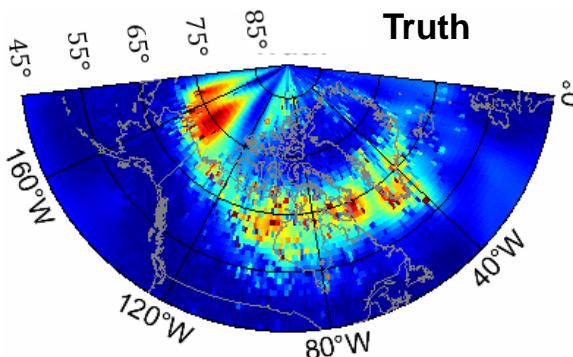
- Lead time steps: 10 minutes
- Resolution: 1° Latitude by 1° Longitude
- Select data of storm days with SYM-H index < -40 nT



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Ionosphere Disturbance Forecasting Results

Ionosphere Disturbance Forecast Results:



(Liu et al., 2021)



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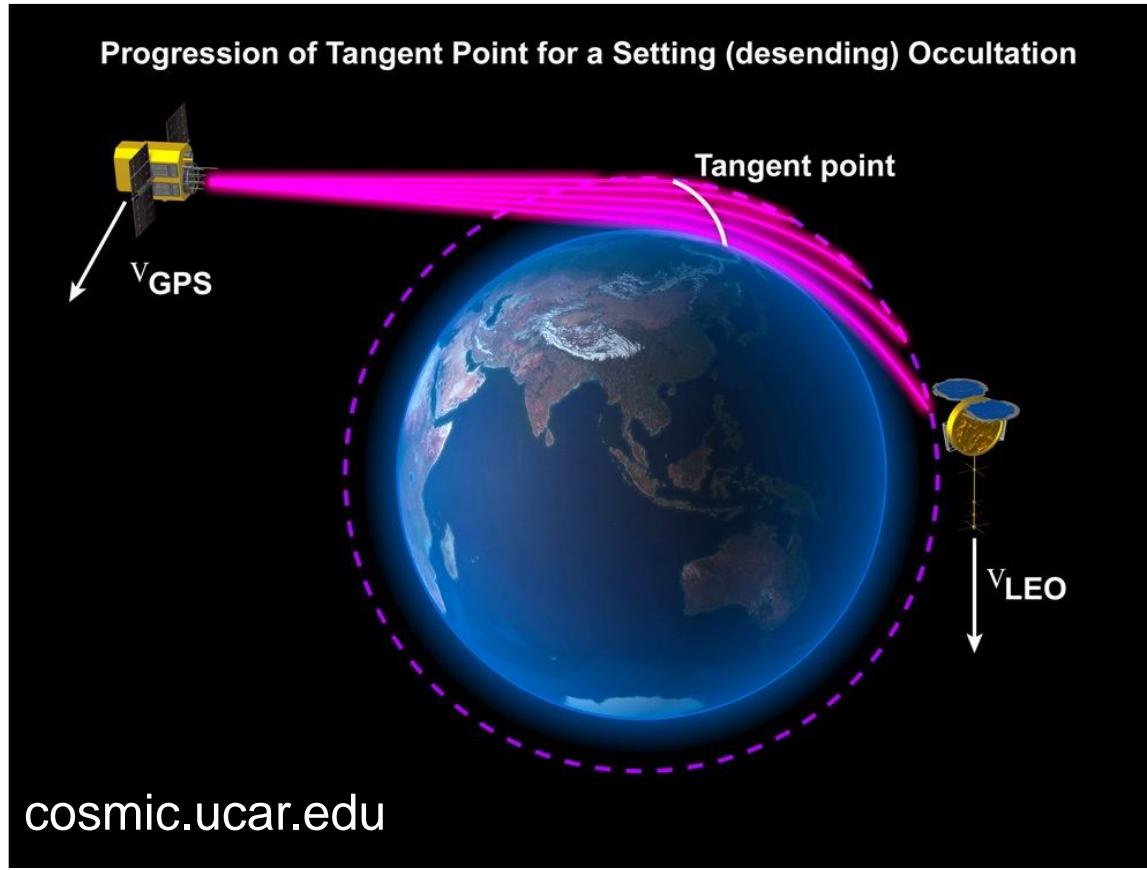
Outline

- Overview of ionosphere observation approaches
- GNSS network-based ionosphere observation and monitoring
- **GNSS space-based ionosphere observations**



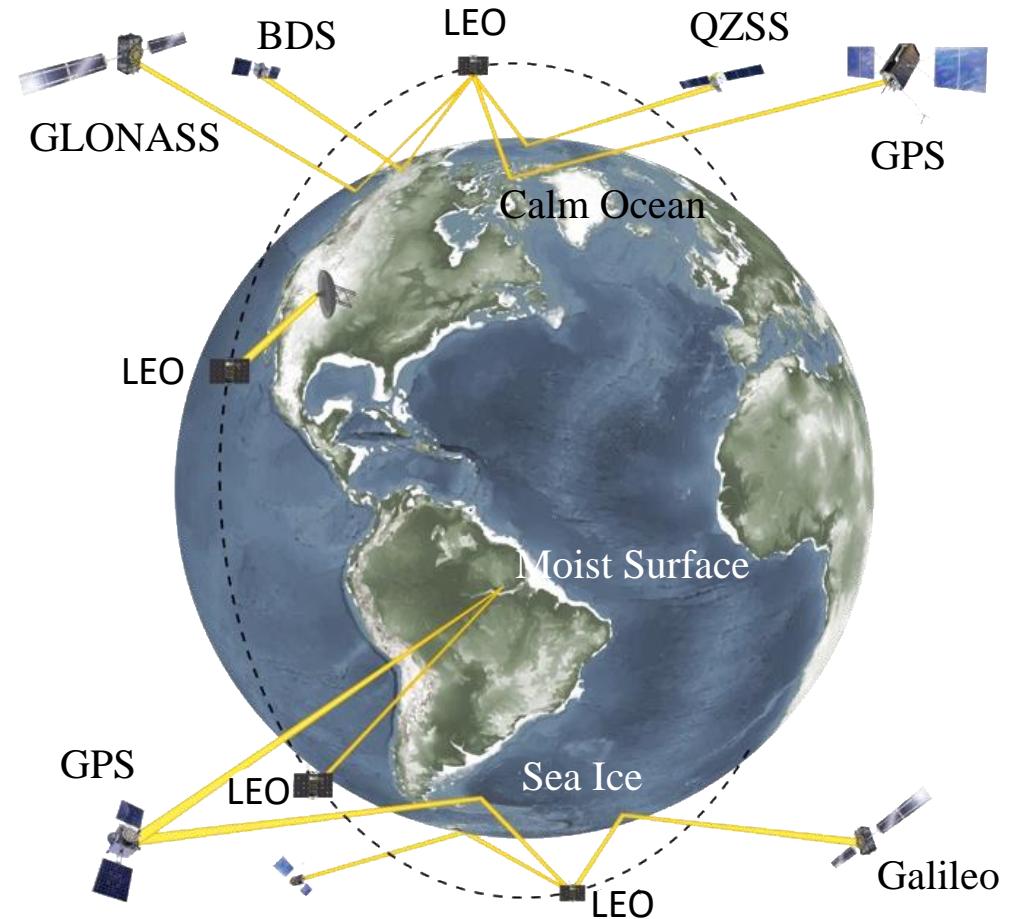
Filling the Data Gap: LEO Satellite-Based Observations

GNSS Radio Occultation (GNSS-RO)



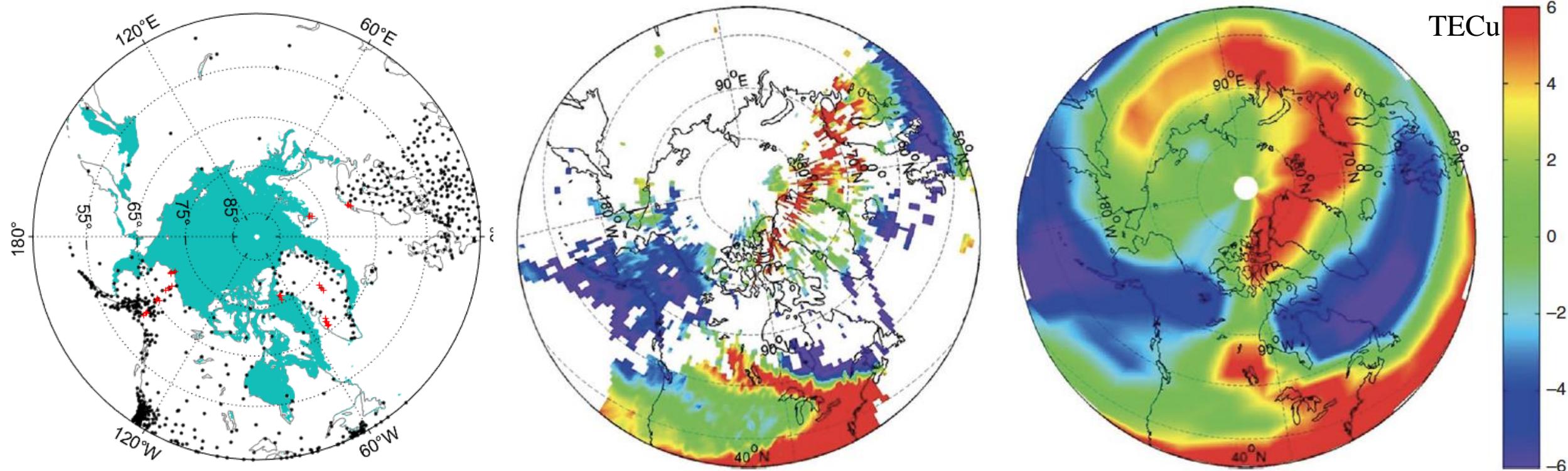
cosmic.ucar.edu

GNSS Reflectometry (GNSS-R)



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GNSS-RO

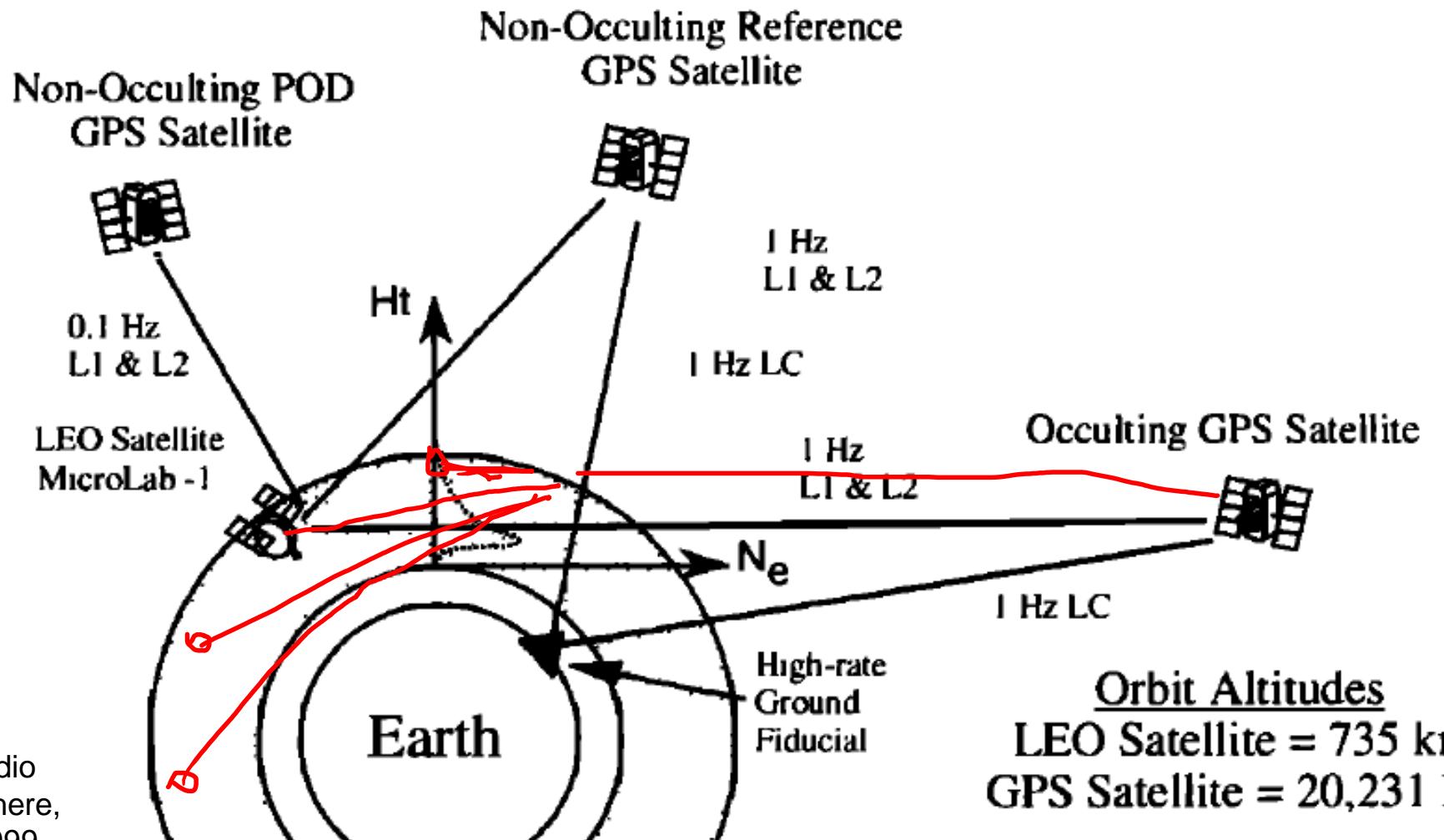


Yue, X., Wan, W., Liu, L., Liu, J., Zhang, S., Schreiner, W. S., ... & Hu, L. (2016). Mapping the conjugate and corotating storm-enhanced density during 17 March 2013 storm through data assimilation. *Journal of Geophysical Research: Space Physics*, 121(12), 12-202.



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GNSS RO Ionosphere Retrieval



Schreiner et al., GPS/MET Radio
Occultation data in the ionosphere,
Radio Scie., 34(4):949-966, 1999.



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GNSS-RO TEC Retrieval

$$\underline{\text{TEC}} \approx \frac{1}{\beta} \Delta\phi_{12} + \underline{\Delta B_{12}}$$

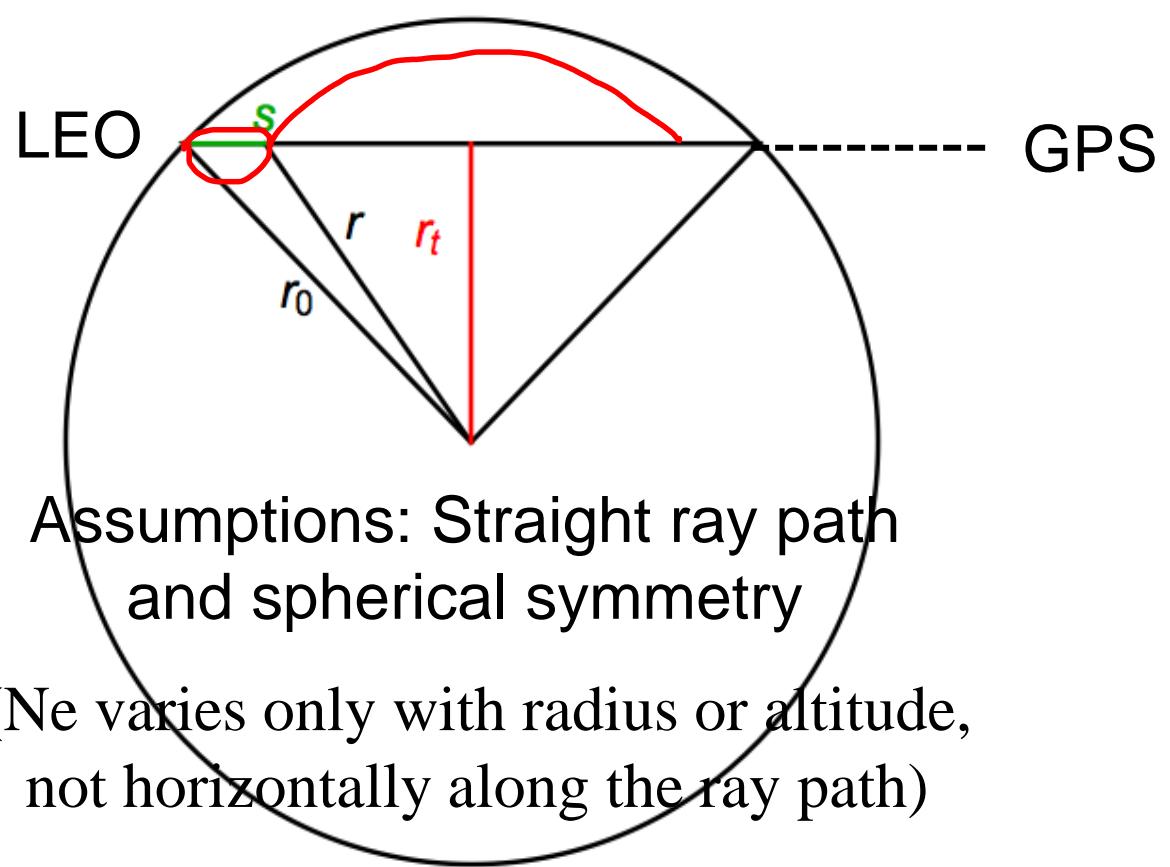
$$\frac{1}{\beta} = \frac{1}{40.3} \frac{f_1^2 f_2^2}{f_2^2 - f_1^2}$$

How to calibrate/estimate bias?

- GNSS satellite bias: use ground receiver network estimations, IGS products
- LEO satellite receiver bias:
 - Find geometries that tend to result in minimum TEC along a raypath and use climatological models of ionosphere to estimate the small Ne and TEC in the region. Example: at high latitudes where the ray path traverses regions of open magnetic fields near the poles.
 - Set TEC to 0 along minimum TEC ray path
 - Rely on receiver built-in calibration mechanism



Ionosphere Ne Profile Retrieval



Mannucci et al., Chapter 31 GNSS Radio Occultation, PNT21, 2020

$$TEC = \int_{raypath} N_e(s) ds$$

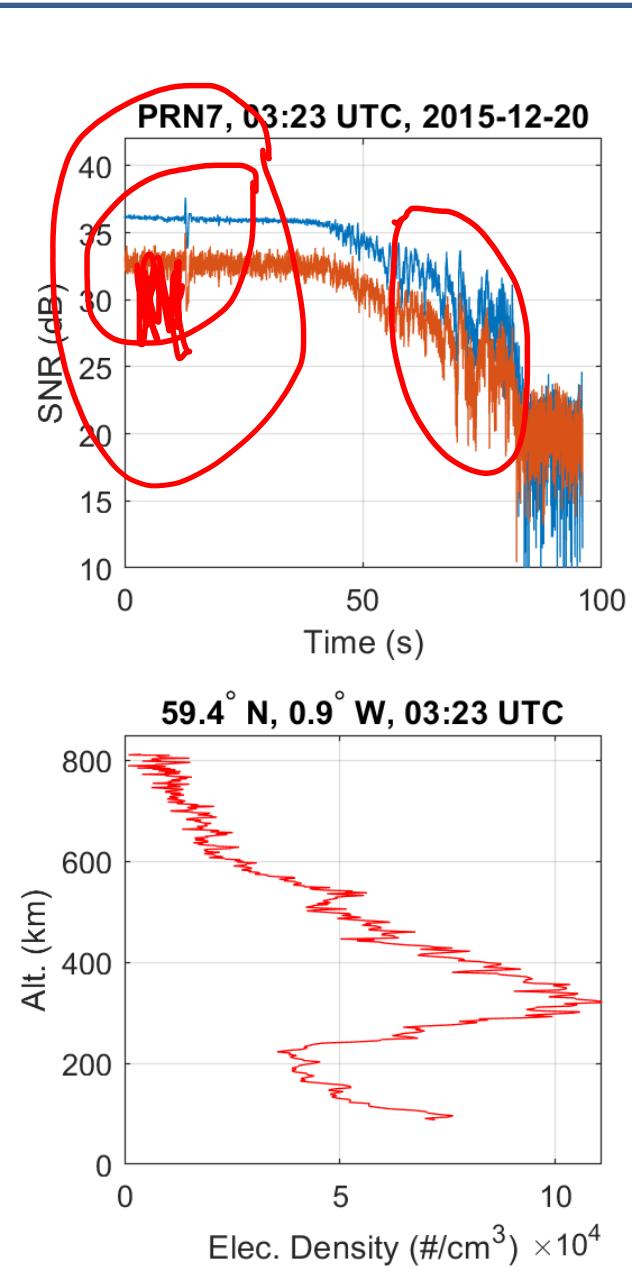
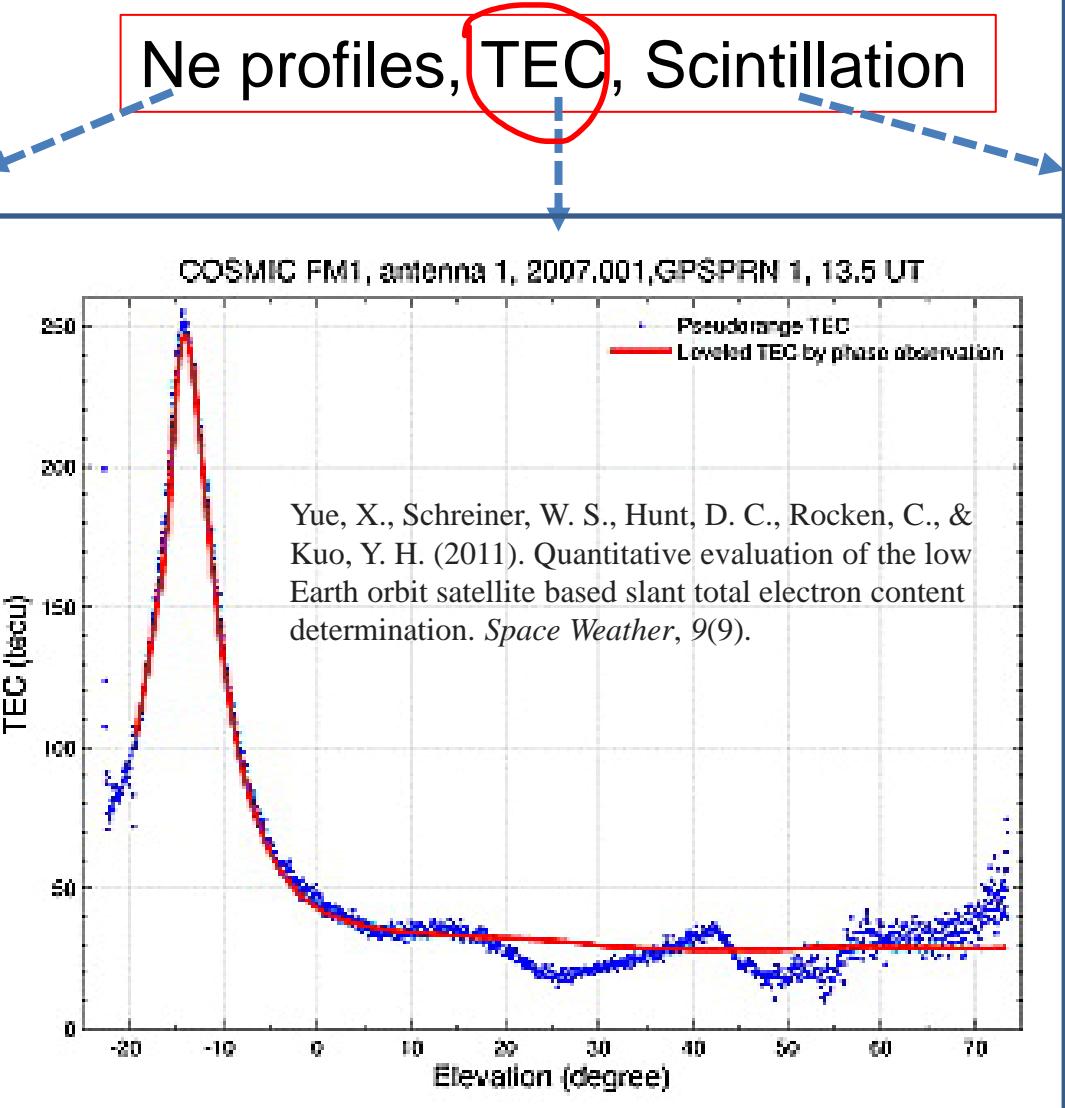
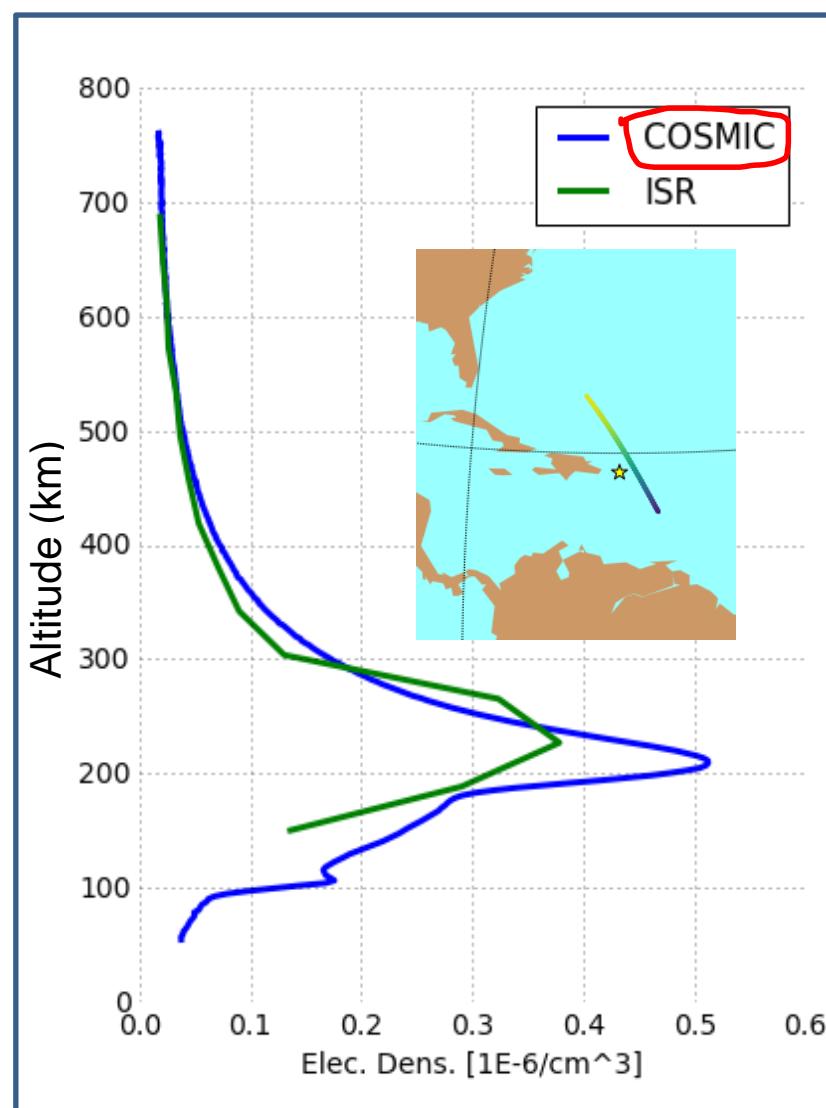
$$s(r) = \sqrt{r_0^2 - r_t^2} - \sqrt{r^2 - r_t^2}$$

$$\frac{ds(r)}{dr} = \frac{r}{\sqrt{r^2 - r_t^2}}$$

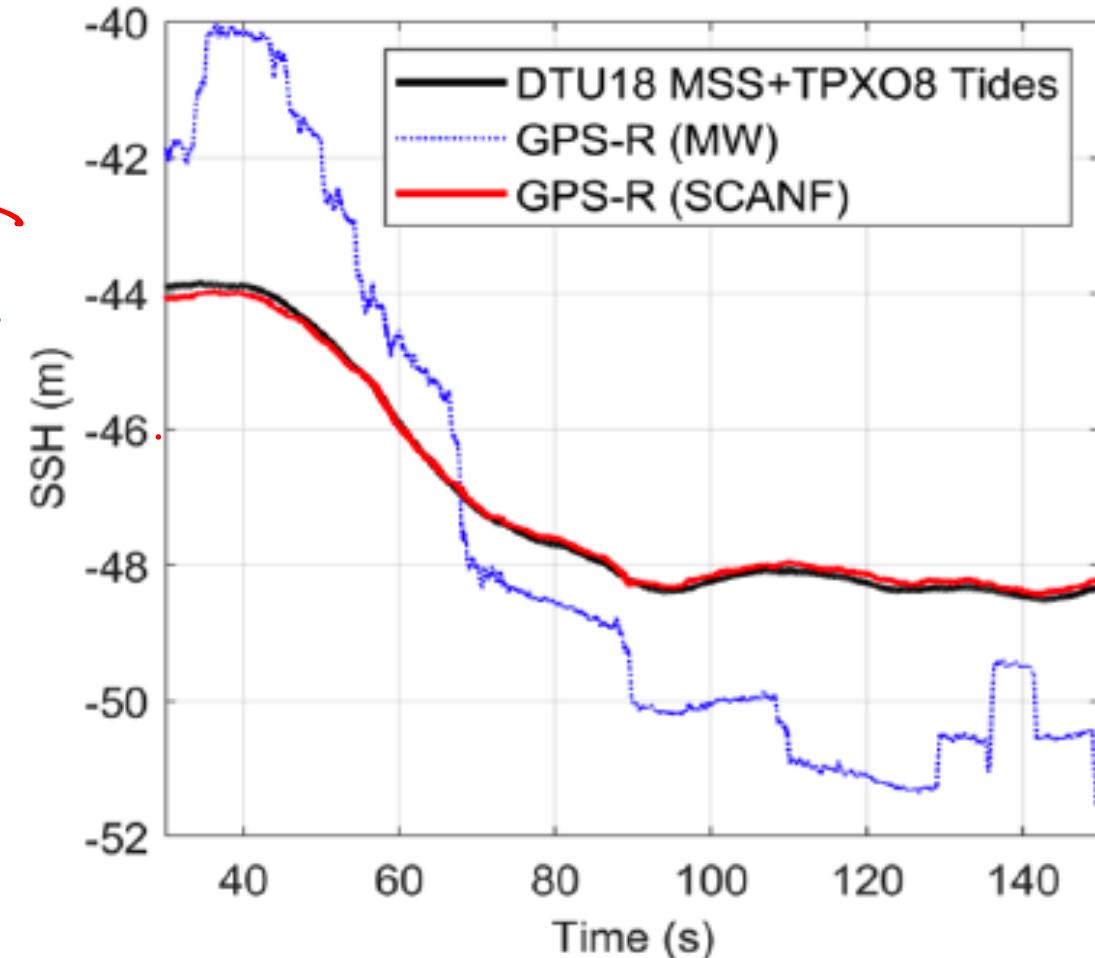
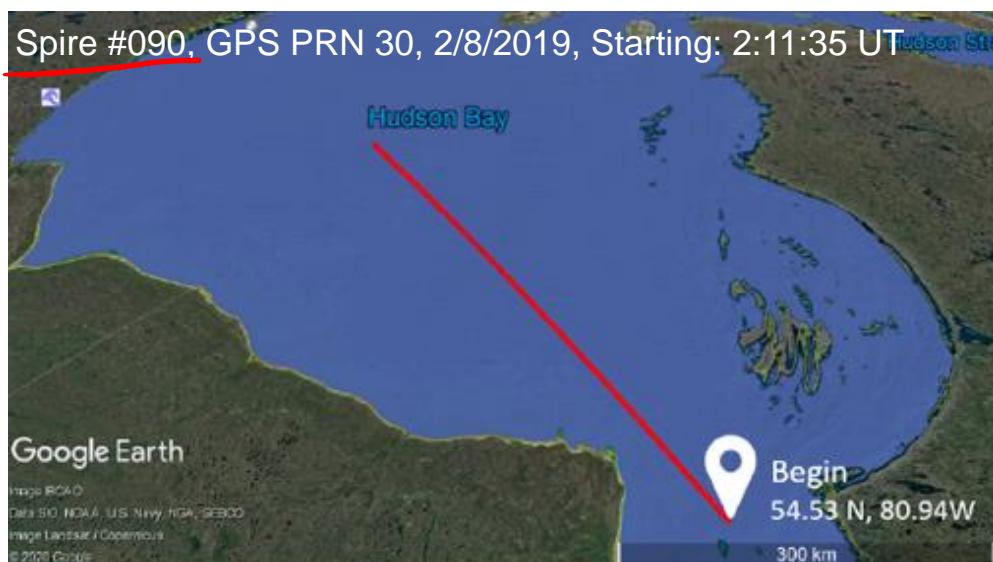
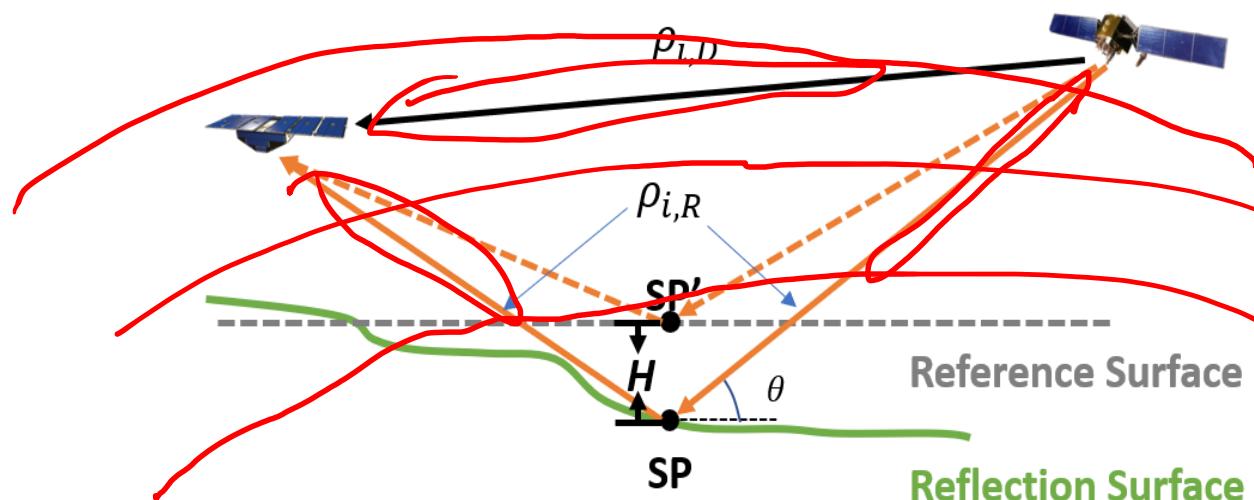
$$TEC(r) = 2 \int_{r_0}^{r_t} \frac{r N_e(r)}{\sqrt{r^2 - r_t^2}} dr$$

$$N_e(r) = \frac{1}{\pi} \int_r^{r_0} \frac{dTEC}{dr_t} \frac{1}{\sqrt{r_t^2 - r^2}} dr_t$$

Ionospheric Observations from GNSS-RO



GNSS-R Phase-Delay Altimetry

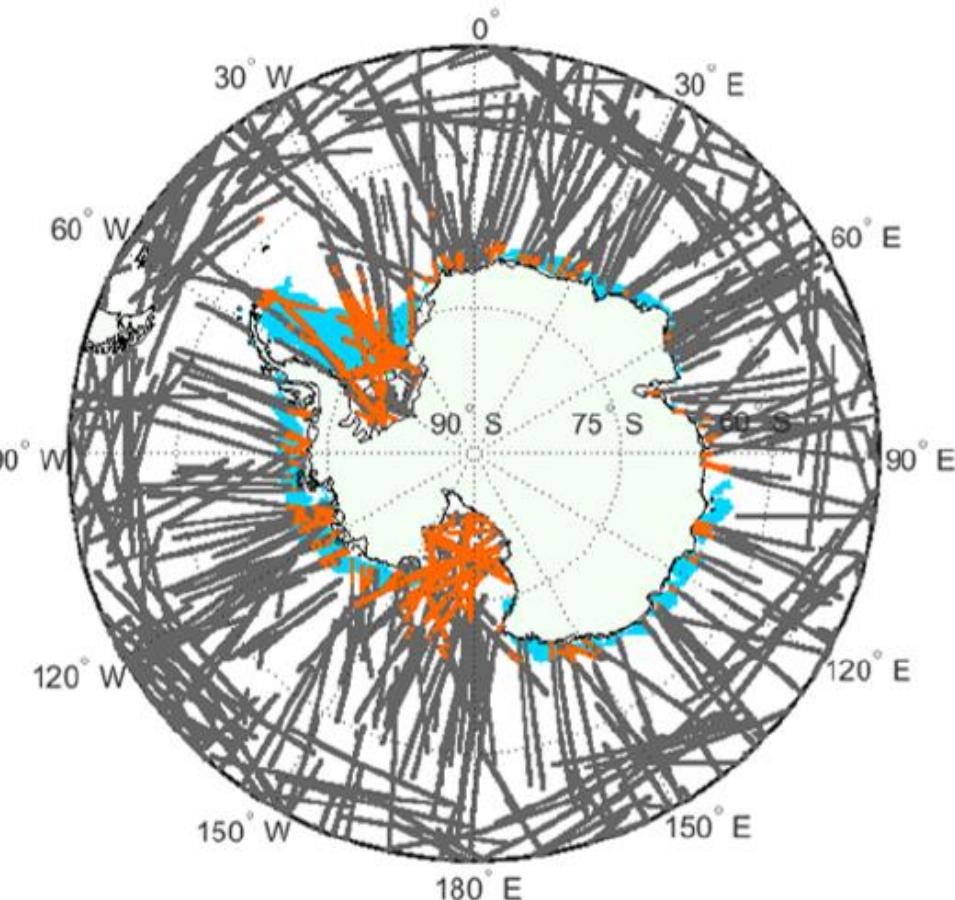
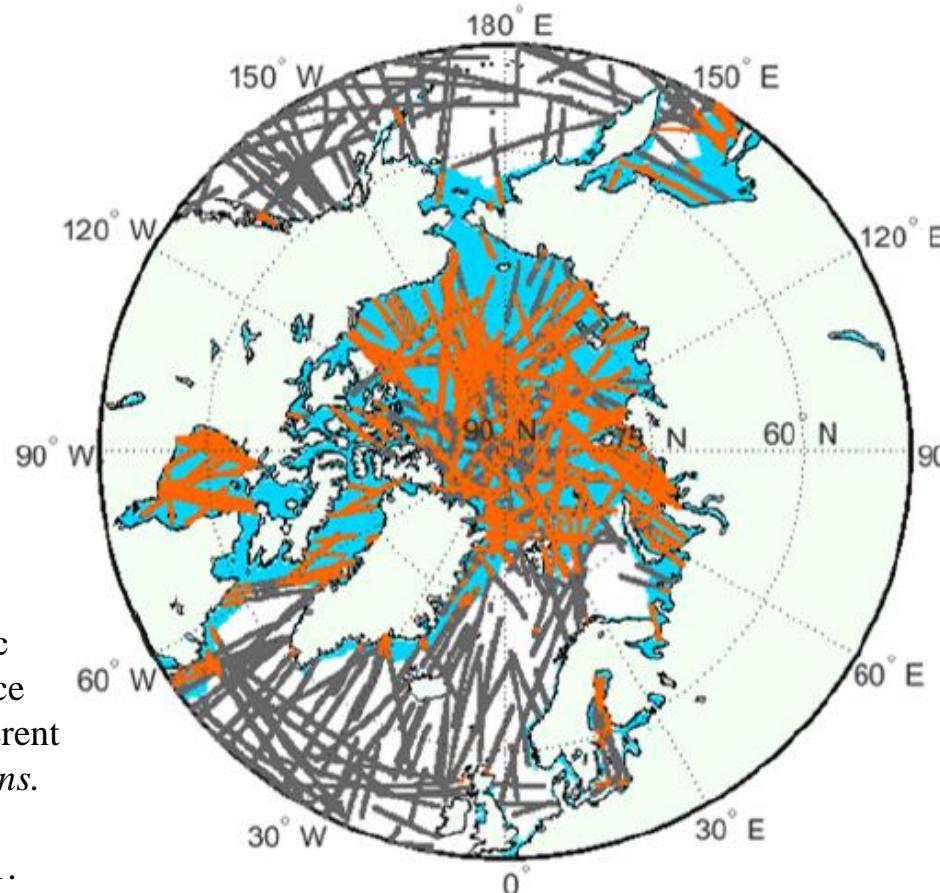


Wang, Y., Y. Morton, "Coherent GNSS reflection signal processing for high-precision and high-resolution spaceborne applications," *IEEE Trans. Geosci. Remote Sensing*, doi:10.1109/TGRS.2020.2993804, 2021.

Arctic and Antarctic: High Rate Coherent Reflections

42% over sea ice. 75% over 1st year ice

- Land
- Ice
- Coherent Tracks
- Non-Coherent Tracks



Wang, Y., Y. J. Morton, "Ionospheric total electron content and disturbance observations from space borne coherent GNSS-R measurements," *IEEE Trans. Geosci. Remote Sensing*, doi: 10.1109/TGRS.2021.3093328, 2021.

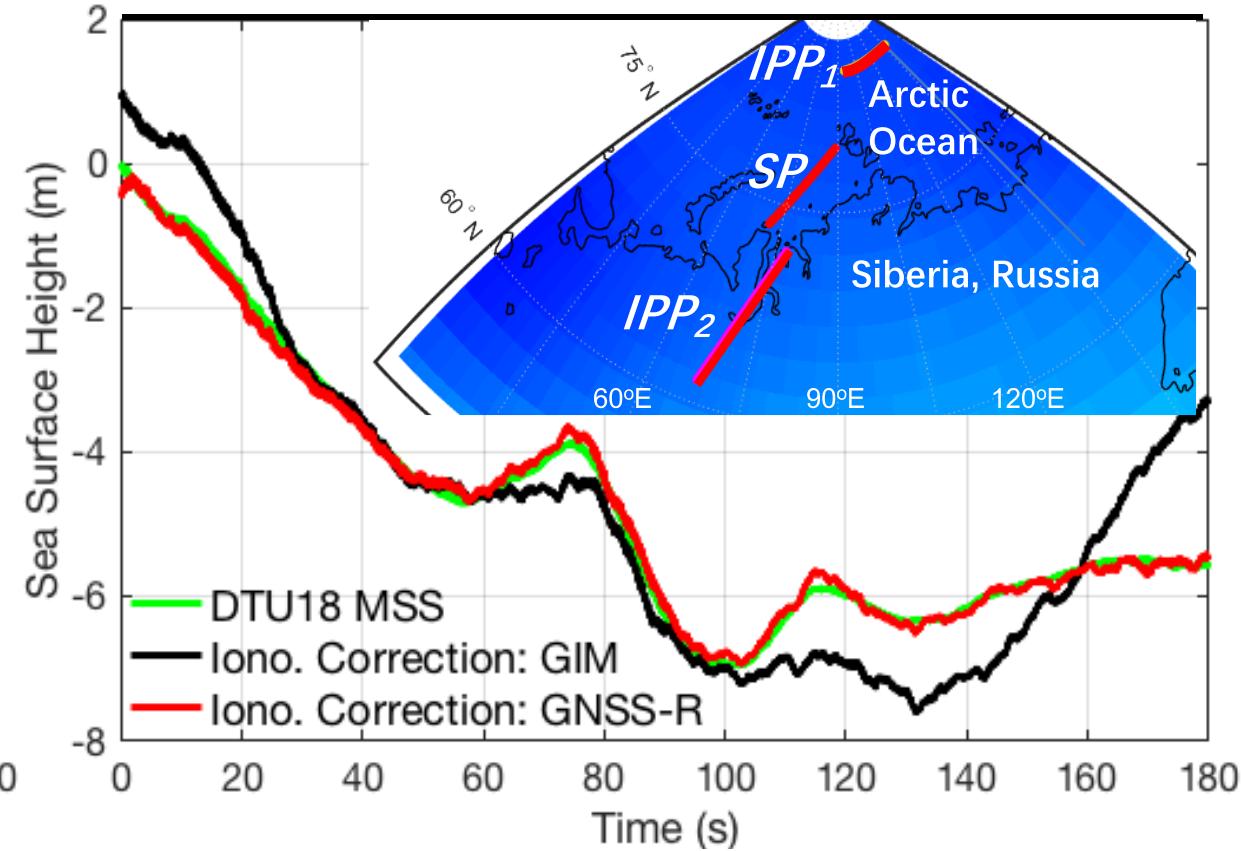
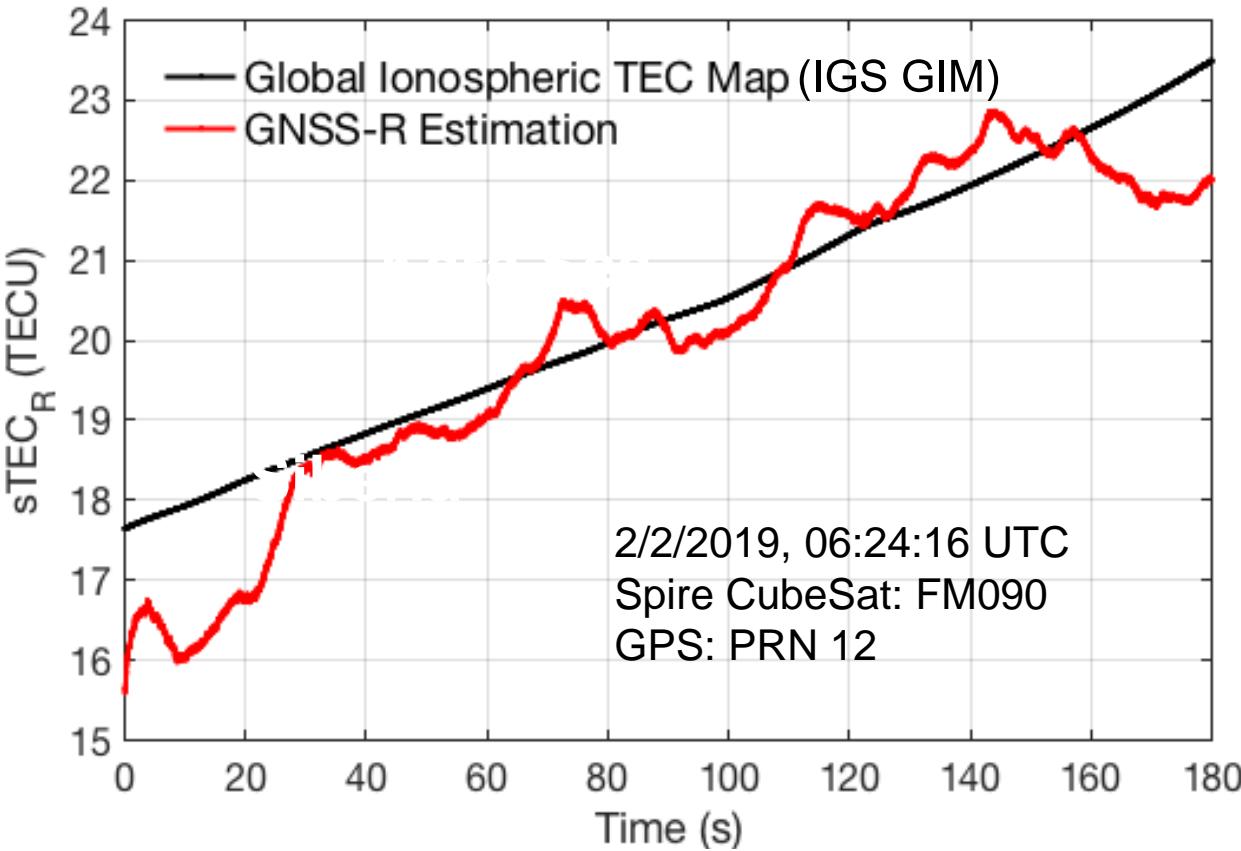


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2021 ION GNSS+

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Example TEC Retrieval from Spire Data: Kara Sea

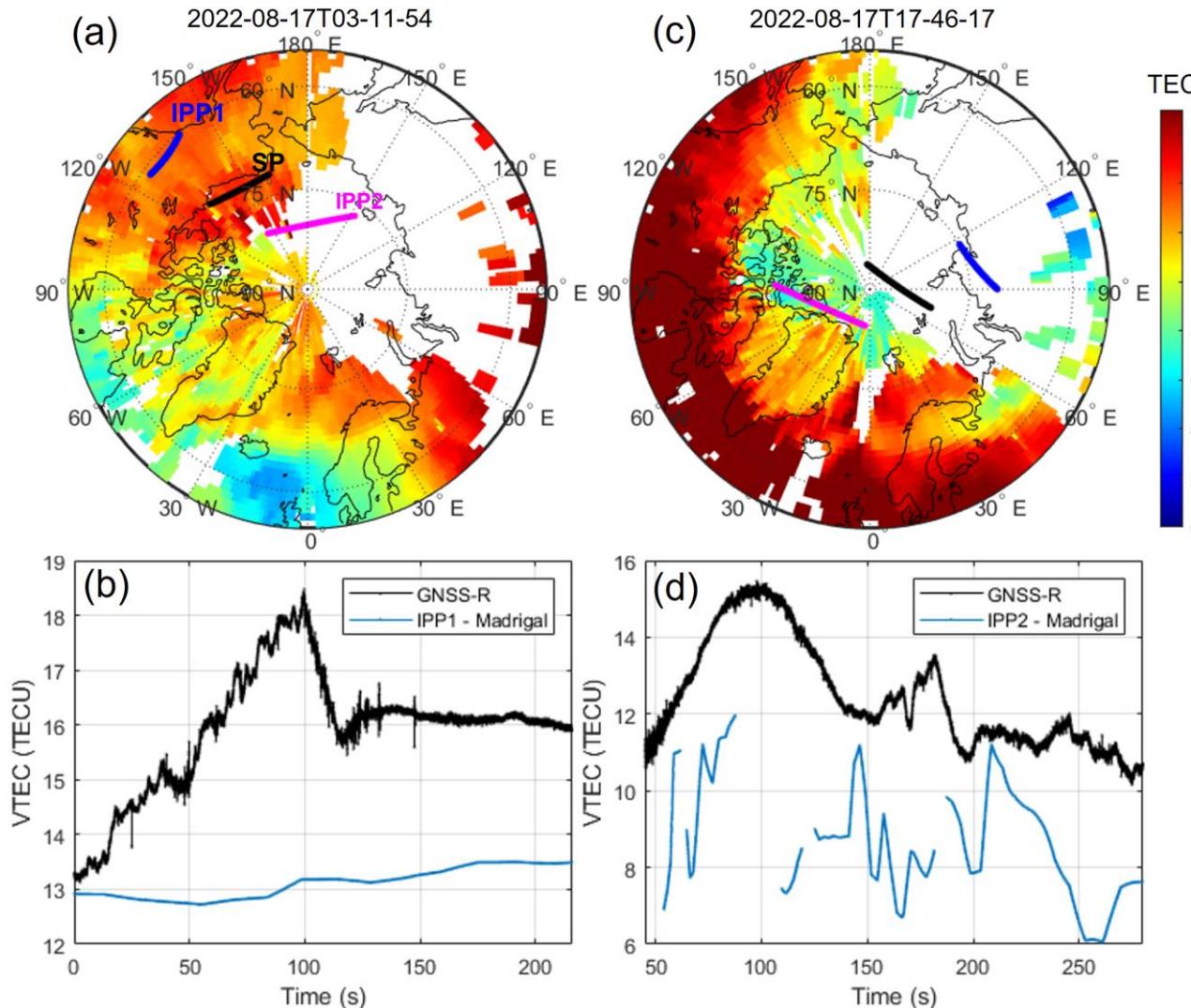


Wang, Y., Y. J. Morton, "Ionospheric total electron content and disturbance observations from space borne coherent GNSS-R measurements," *IEEE Trans. Geosci. Remote Sensing*, doi: 10.1109/TGRS.2021.3093328, 2021.



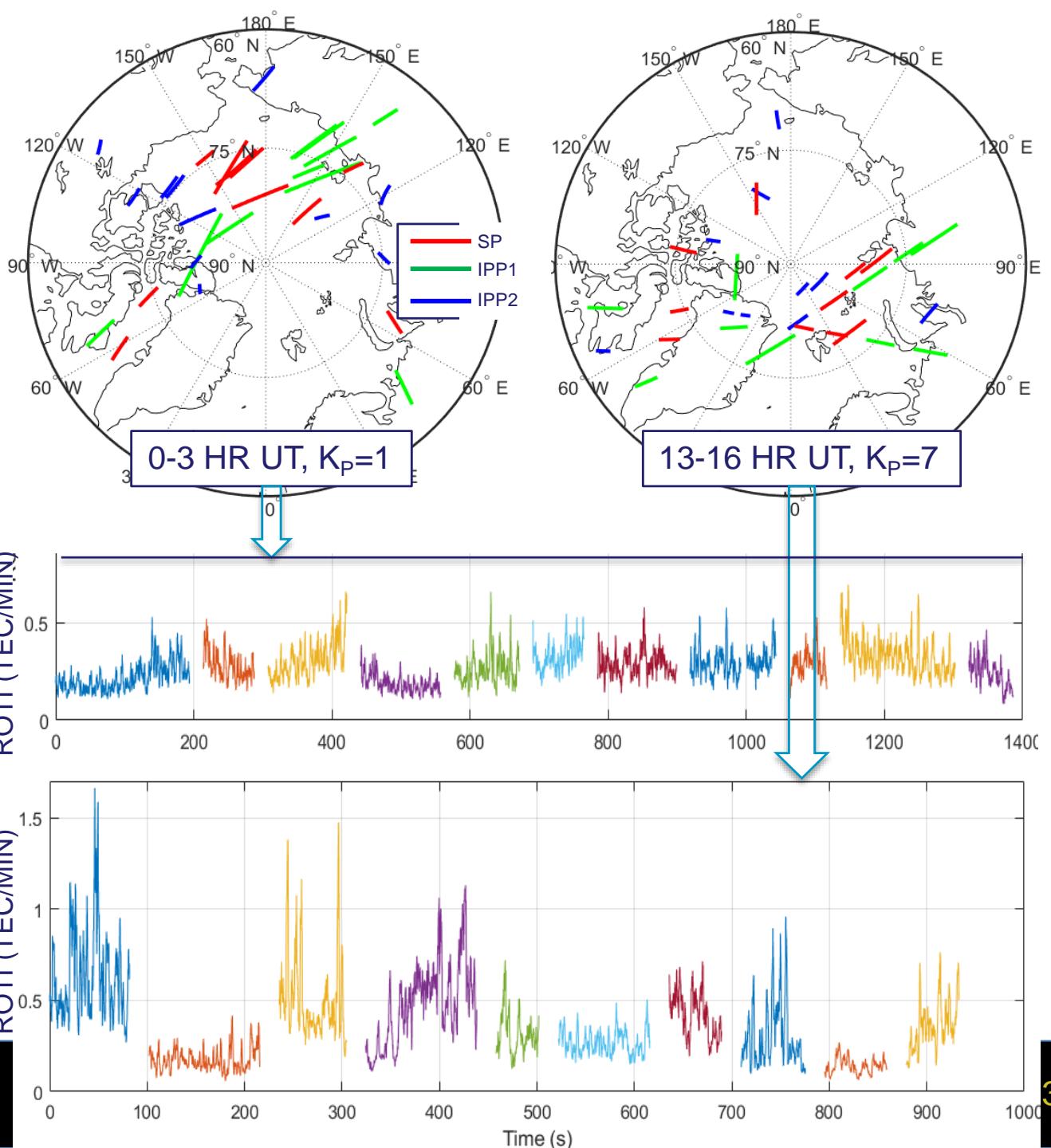
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Ionosphere Structure Observation GNSS-R



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GNSS-R Monitoring Ionospheric Disturbances



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