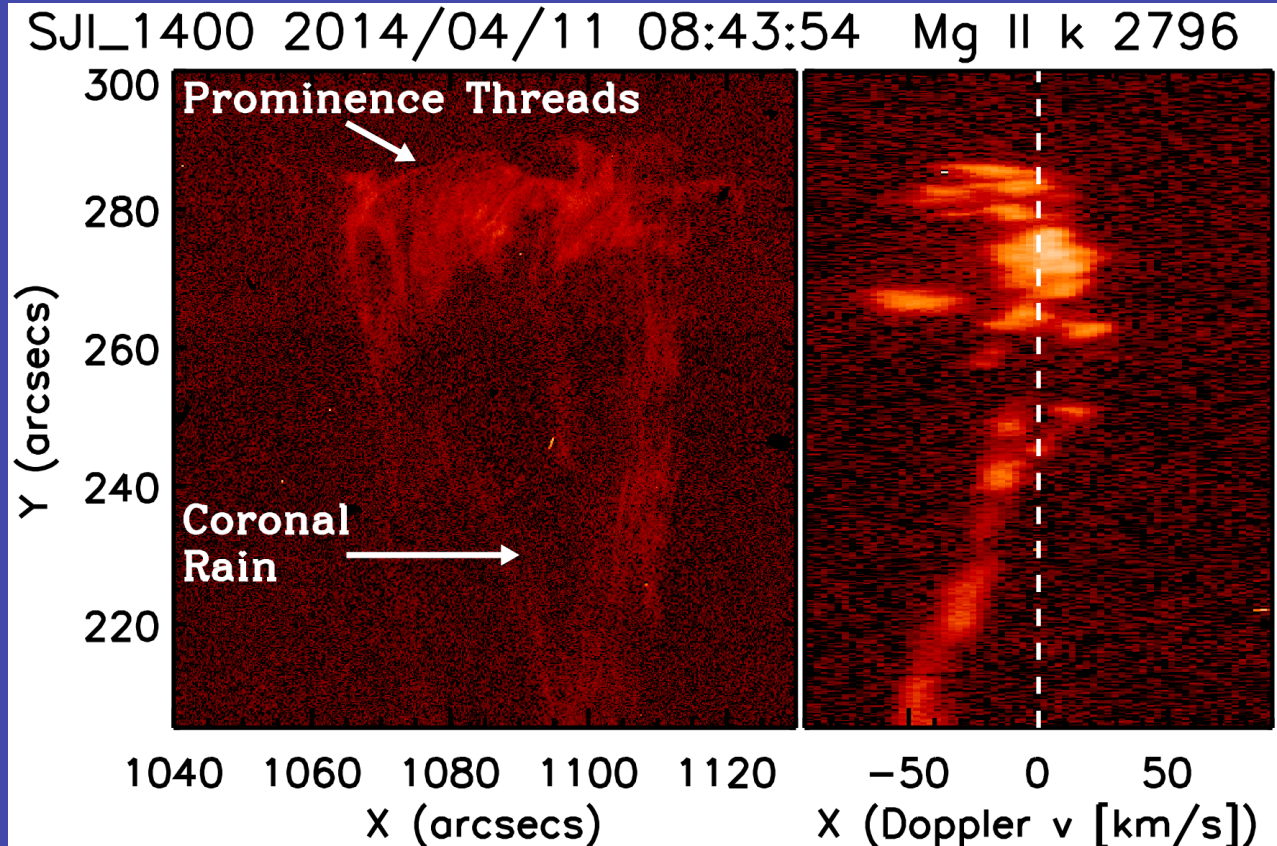
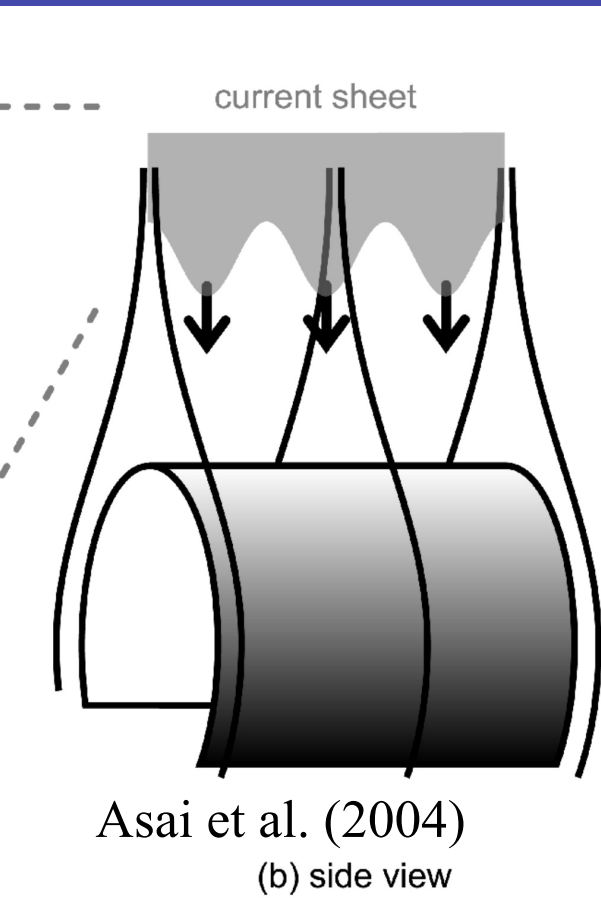


# A Novel, Hybrid Prominence-Coronal Rain Complex Near Magnetic Null Points

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# Outline

## 1. Introduction

Chromosphere–corona mass cycle, two forms of return flows: coronal rain and prominences (similarities and differences).

## 2. IRIS observations of hybrid prominence – coronal rain near null points, 2014/04/11

## 3. Why condensation near nulls?

More examples.

## 4. Summary

# 1. Introduction – Chromosphere–corona Mass Cycle (e.g., McIntosh+ 2012): prominence and coronal rain as return flows

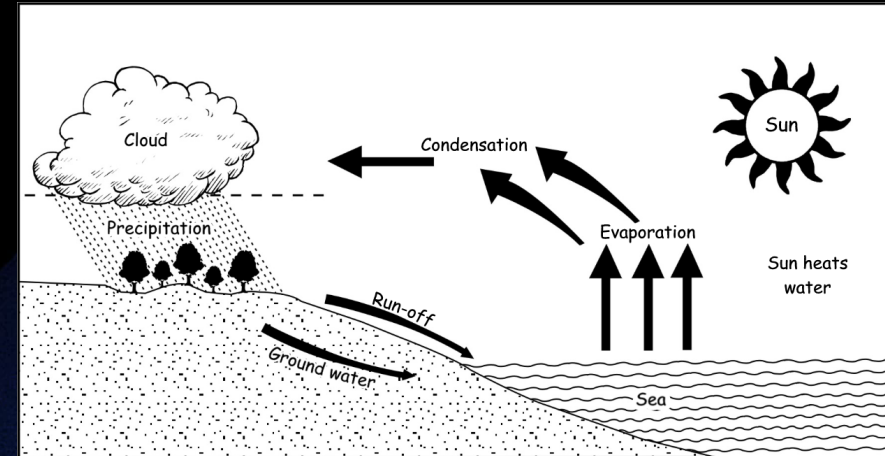
(1) Hot mass/Magnetic flux Up:

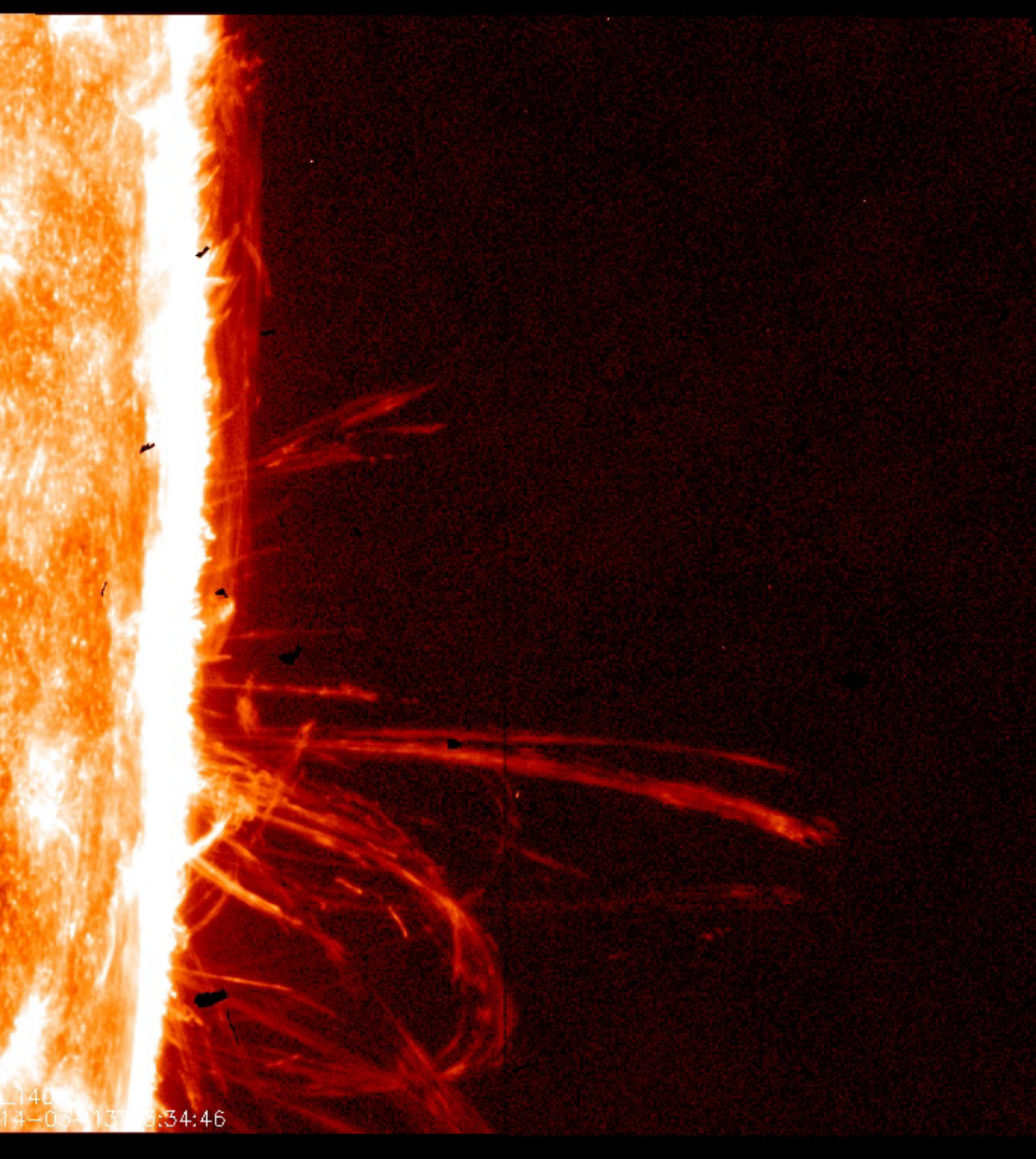
- Spicules, footpoint upflows, flux emergence (e.g., bubbles/plumes)

(2) Cool mass down:

- **Prominences** – filamentary, meandering downflow threads (reviews by, e.g., Labrosse+2010, Mackay+2010)

- **Coronal rain**





coronal rain –  
intermittent in active  
region loops, more  
substantial in flares

Always follows well-  
defined, curved paths

– quite different from  
meandering prominence  
threads

Why?

L146  
14-03-13T09:34:46

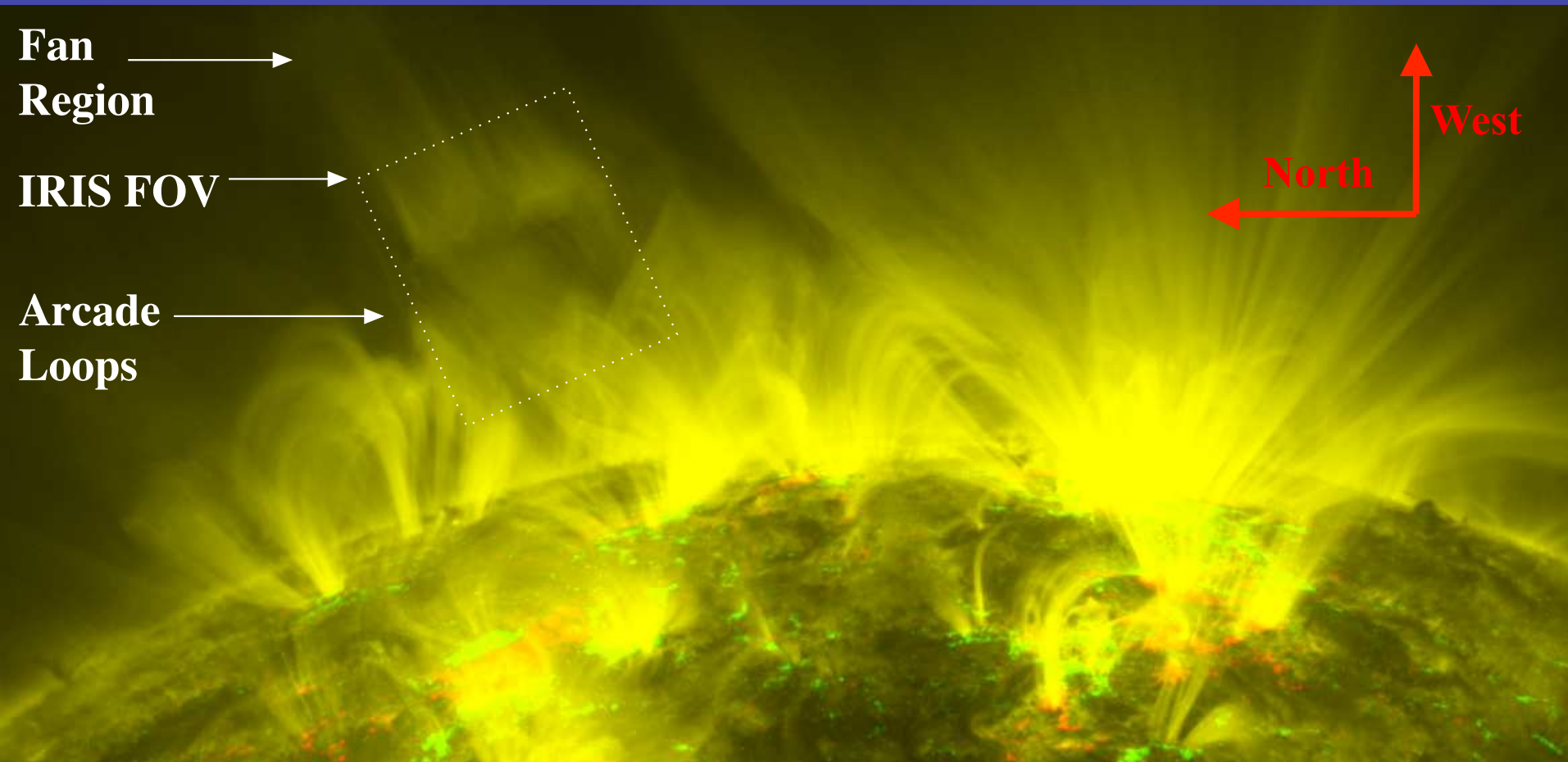
IRIS 1400 Å (Courtesy of Hui Tian)

## 2. Hybrid prominence – coronal rain

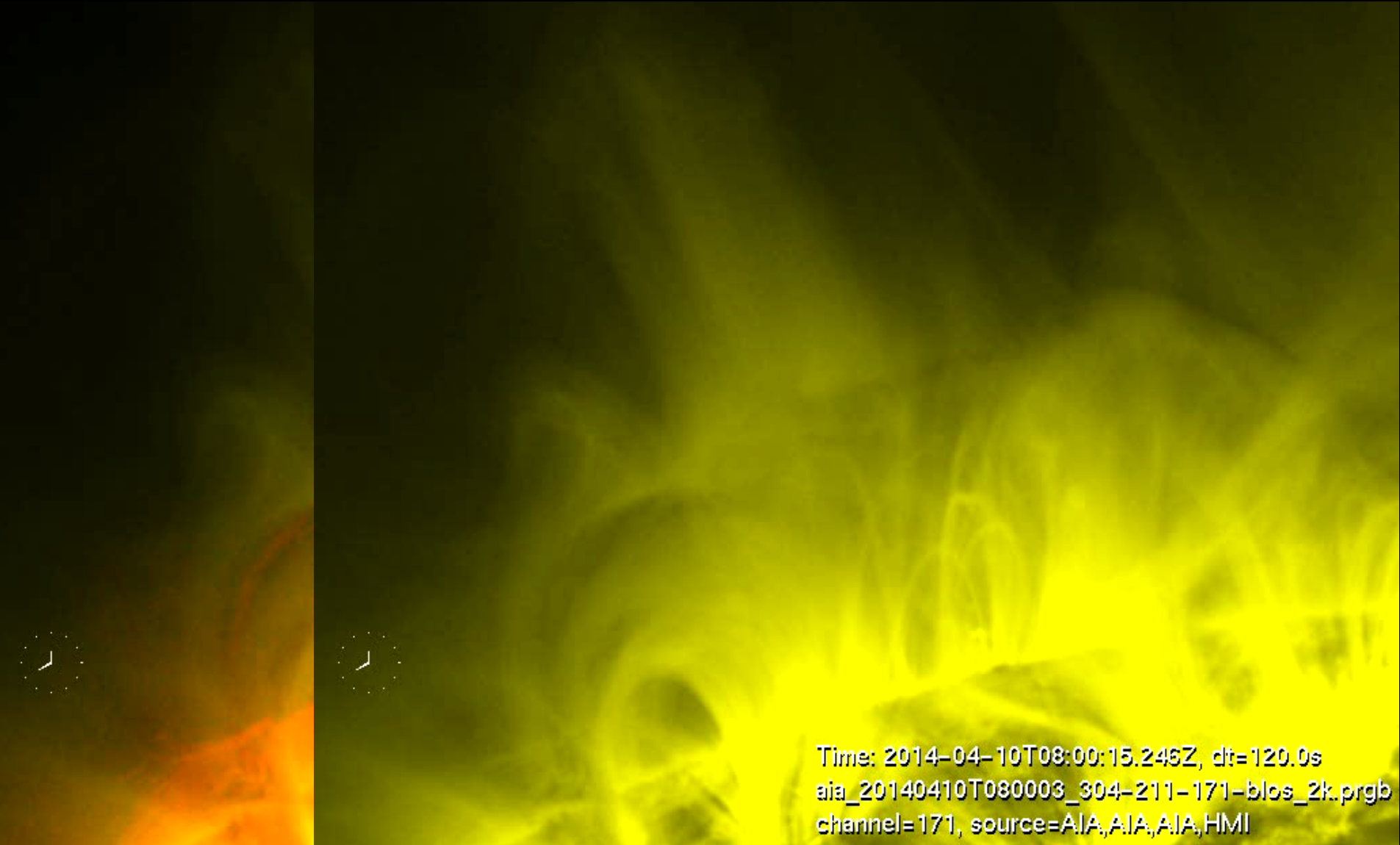
### 2.1) SDO/AIA overview

Top (“fan”-like): “prominence threads”, irregular flows

Bottom (closed loops): coronal rain



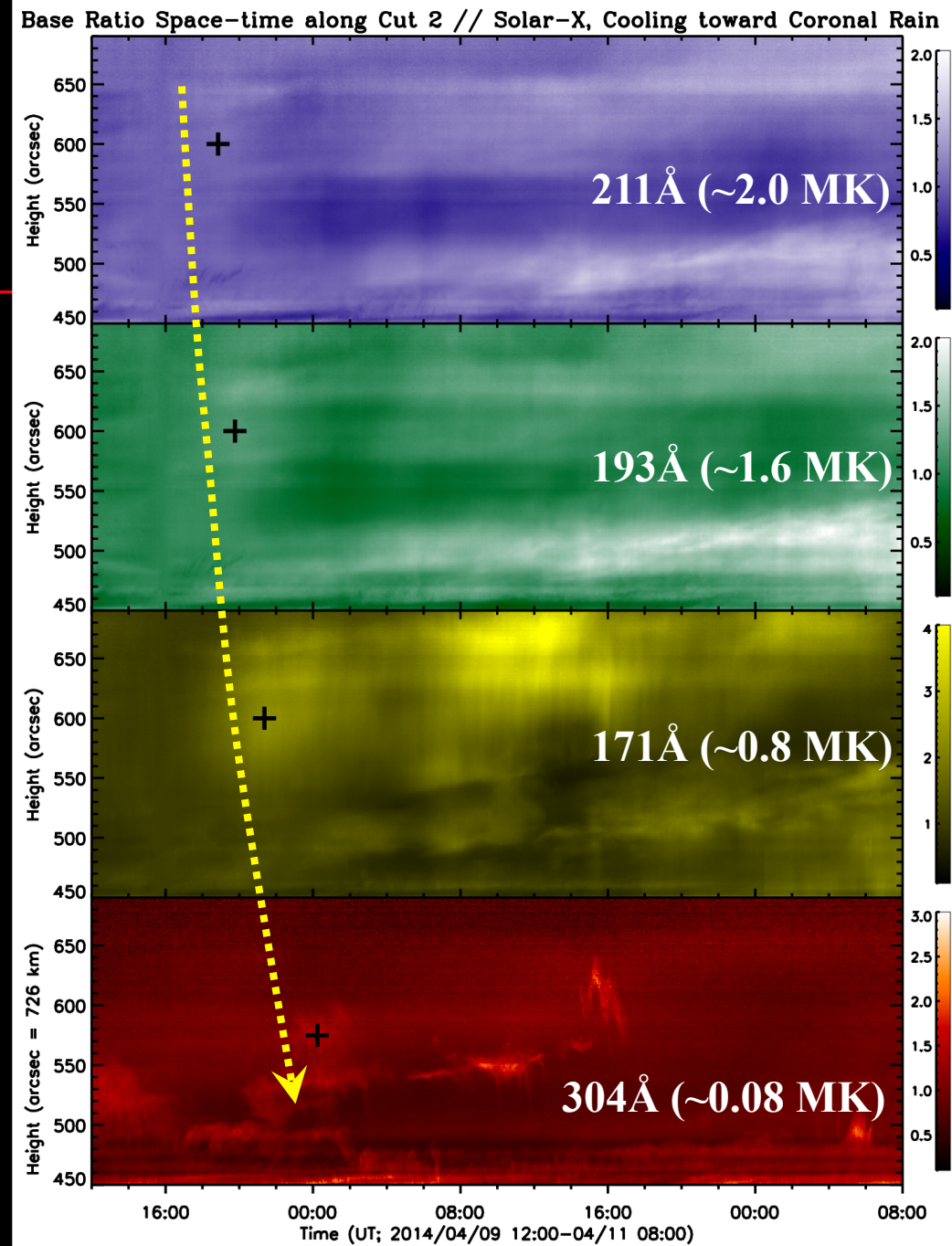
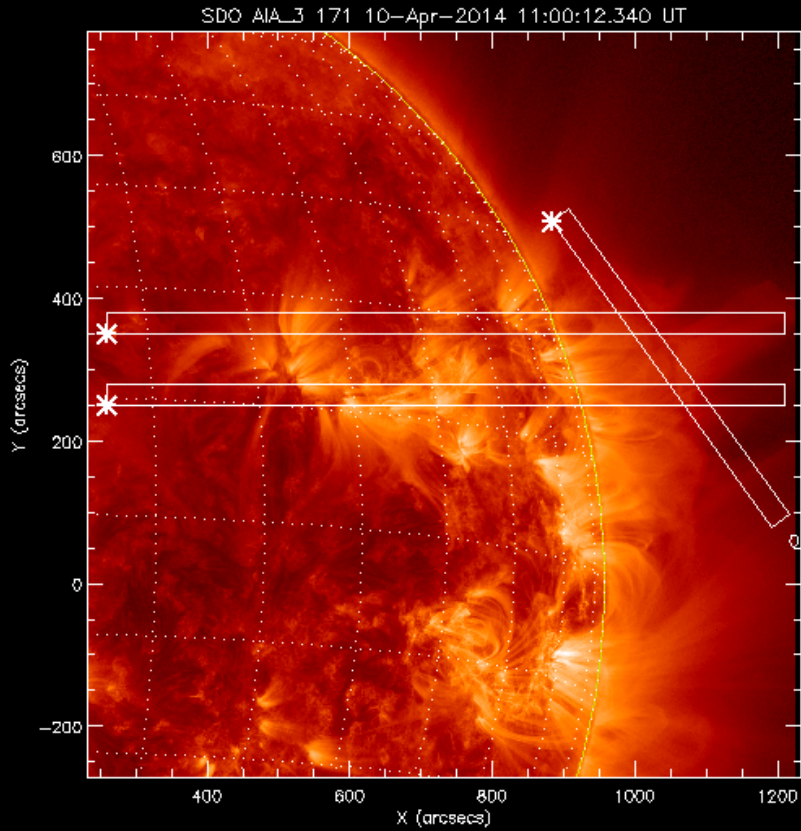
Hybrid prominence + coronal rain (cf., cloud prominences, coronal spiders;  
Lin, Martin, Engolvd 2006; Kleczek 1972);  
Coronal rain originates from prominence-like material AIA 171 Å (Y)/304 Å (R)



Time: 2014-04-10T08:00:15.246Z, dt=120.0s  
aia\_20140410T080003\_304-211-171-blos\_2k.prgb  
channel=171, source=AIA,AIA,AIA,HMI

# Cooling sequence toward condensation

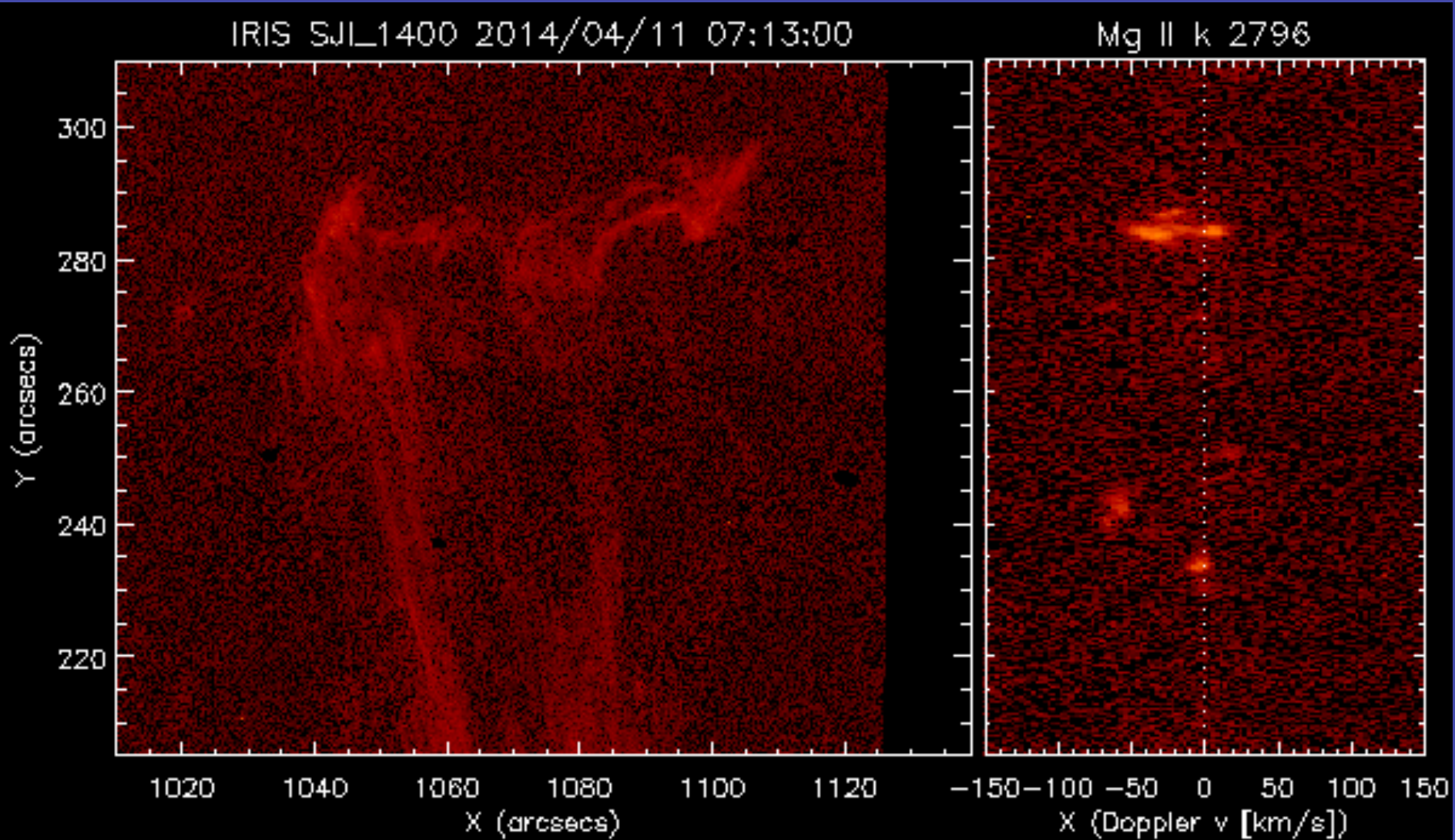
Delayed emission in progressively cooler AIA channels



## 2.2) IRIS: SJI + Mg II k spectra

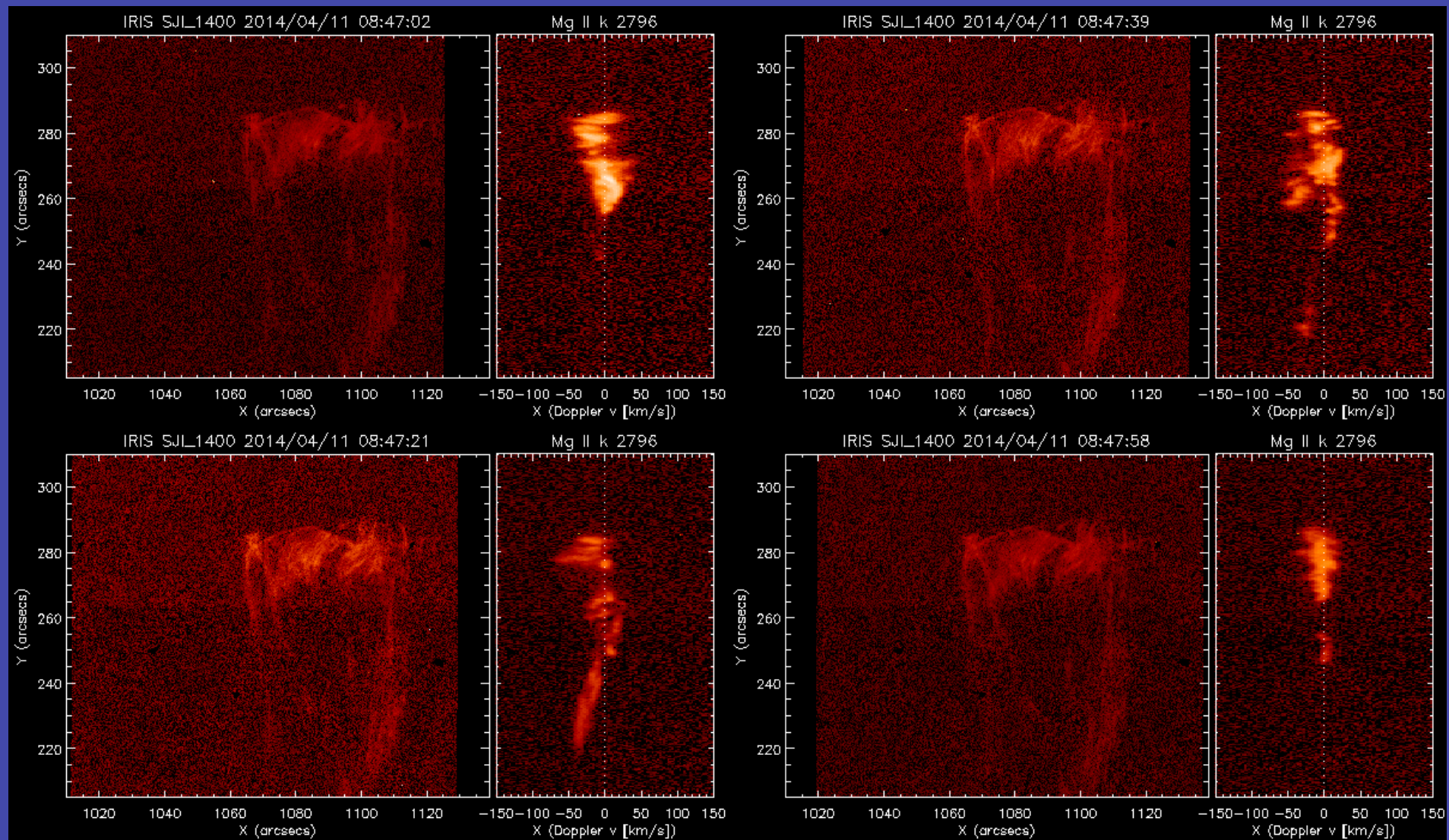
Top: prominence threads, irregular flows, large line width

Bottom (closed loops): coronal rain, narrow line width

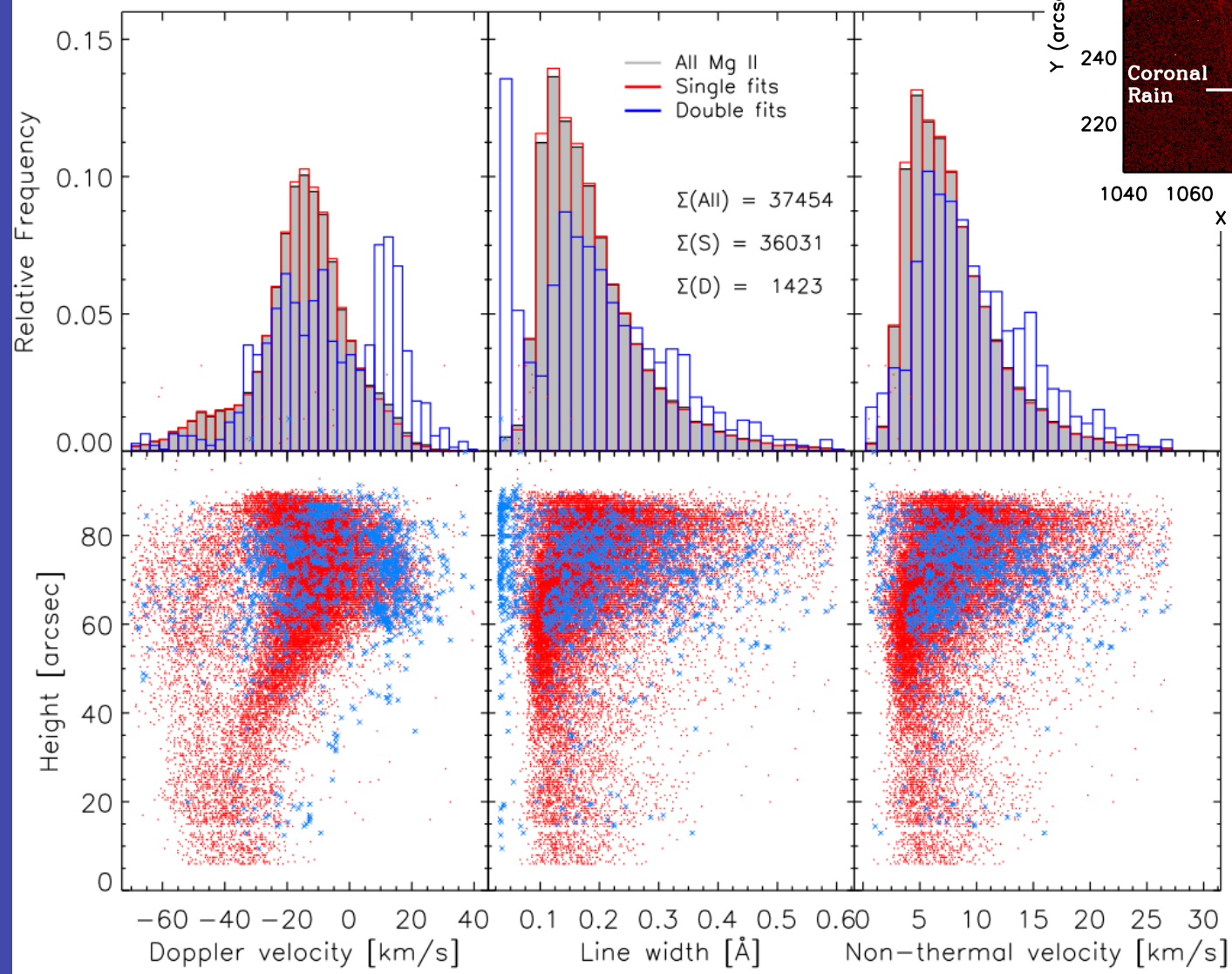
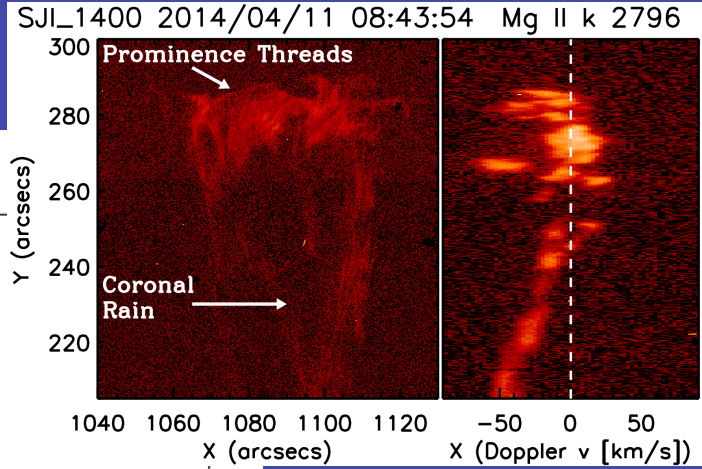




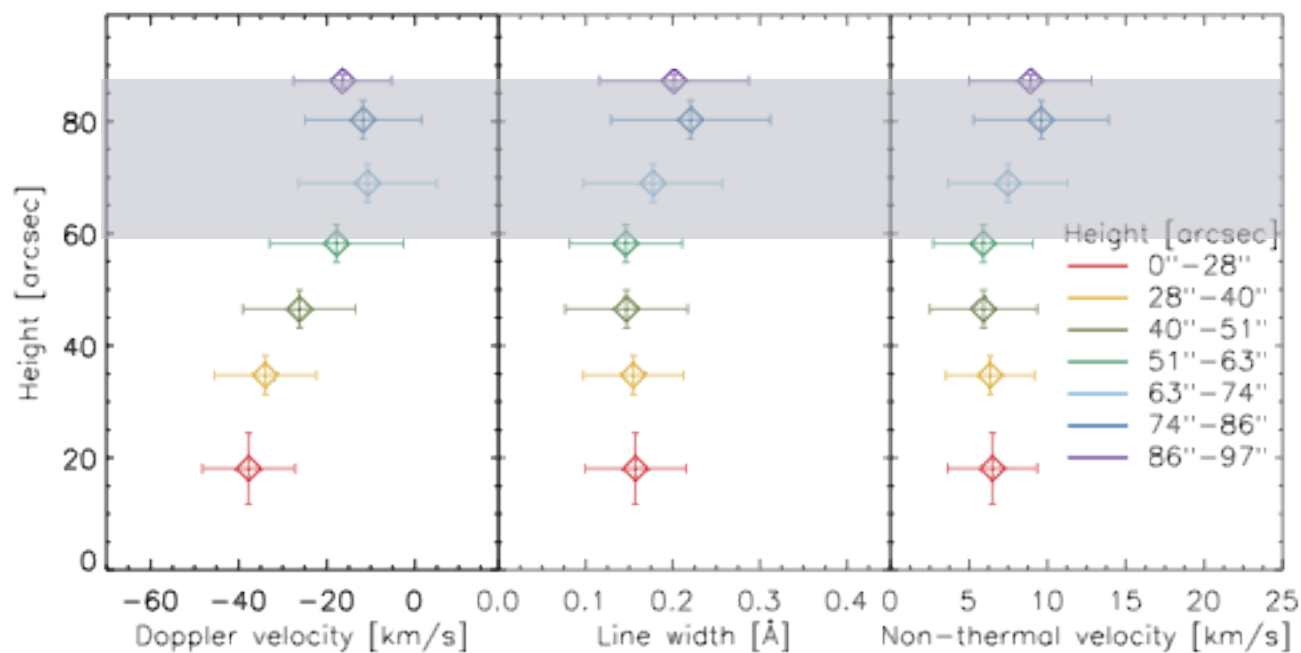
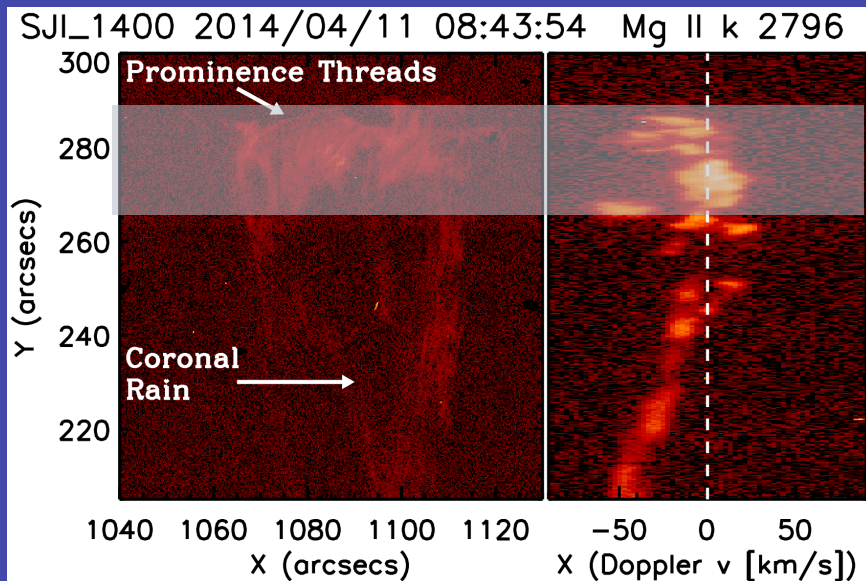
Top: prominence threads, large line width  
Bottom (closed loops): coronal rain, narrow line width



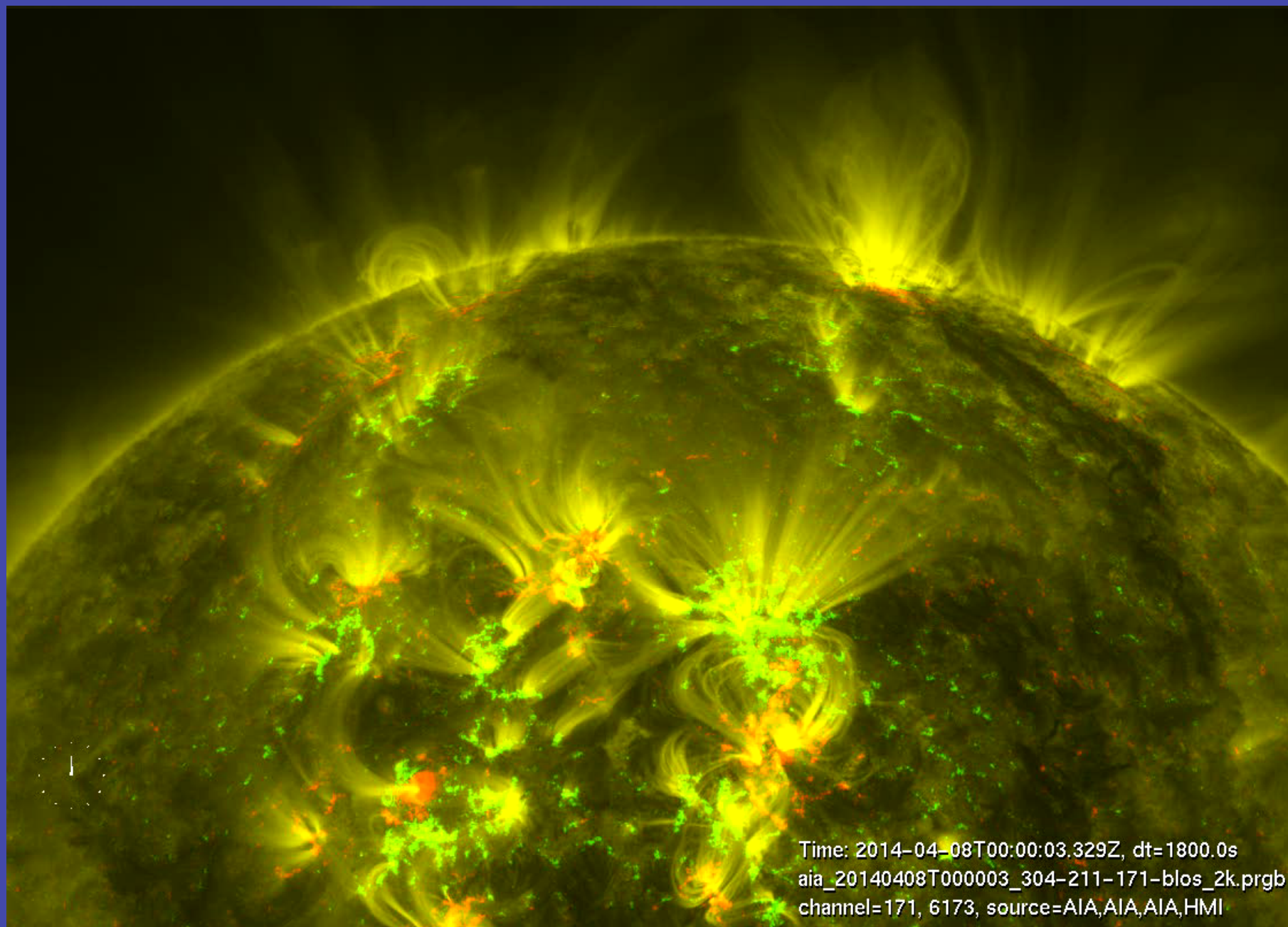
# Spectral fits and height dependence (Patrick Antolin)



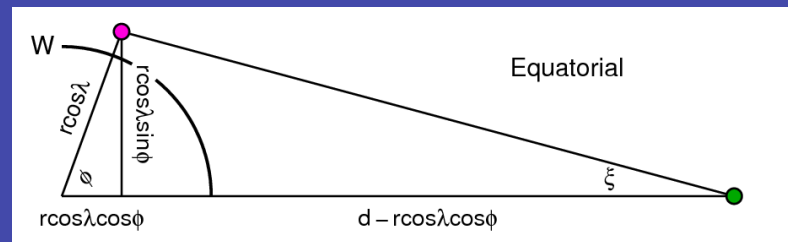
# Line width height dependence (Patrick Antolin)



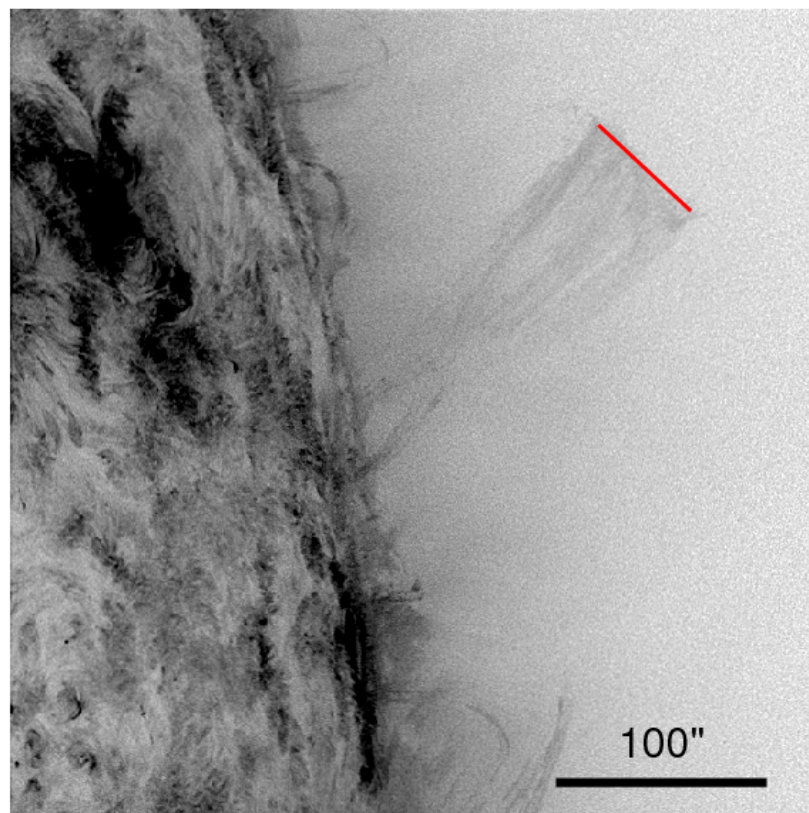
## 2.3) Magnetic geometry: AIA 171 Å (yellow) on HMI magnetogram



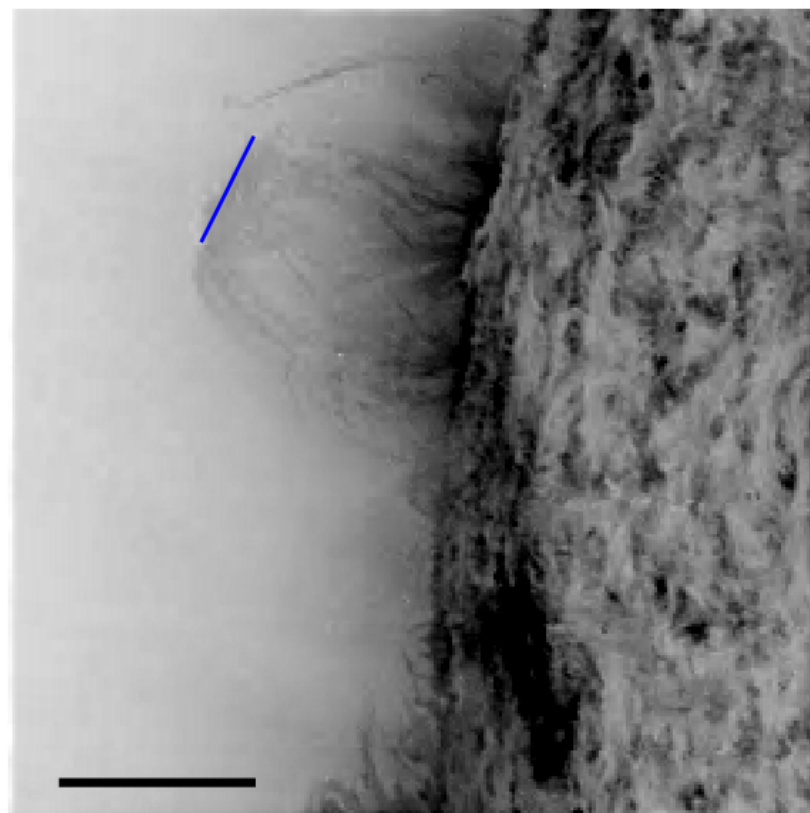
# Localization by triangulation of 304 Å material (Xudong Sun)



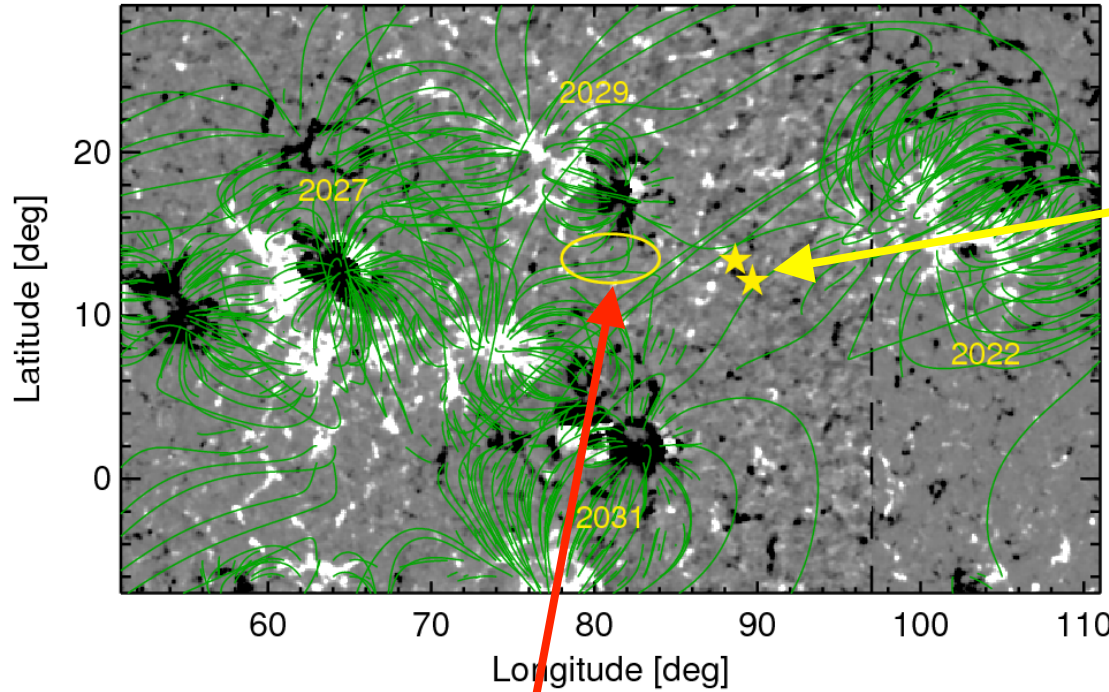
SDO/AIA



STEREO-A/EUVI

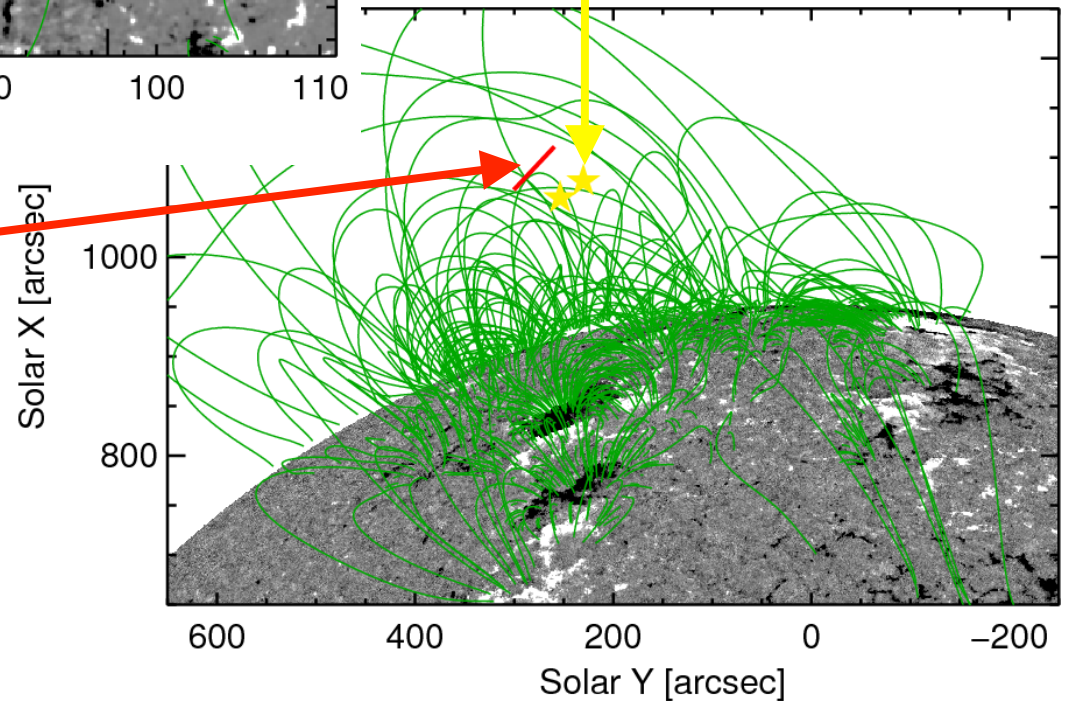


# Magnetic geometry: Potential field (PFSS) model from HMI magnetogram (Xudong Sun)

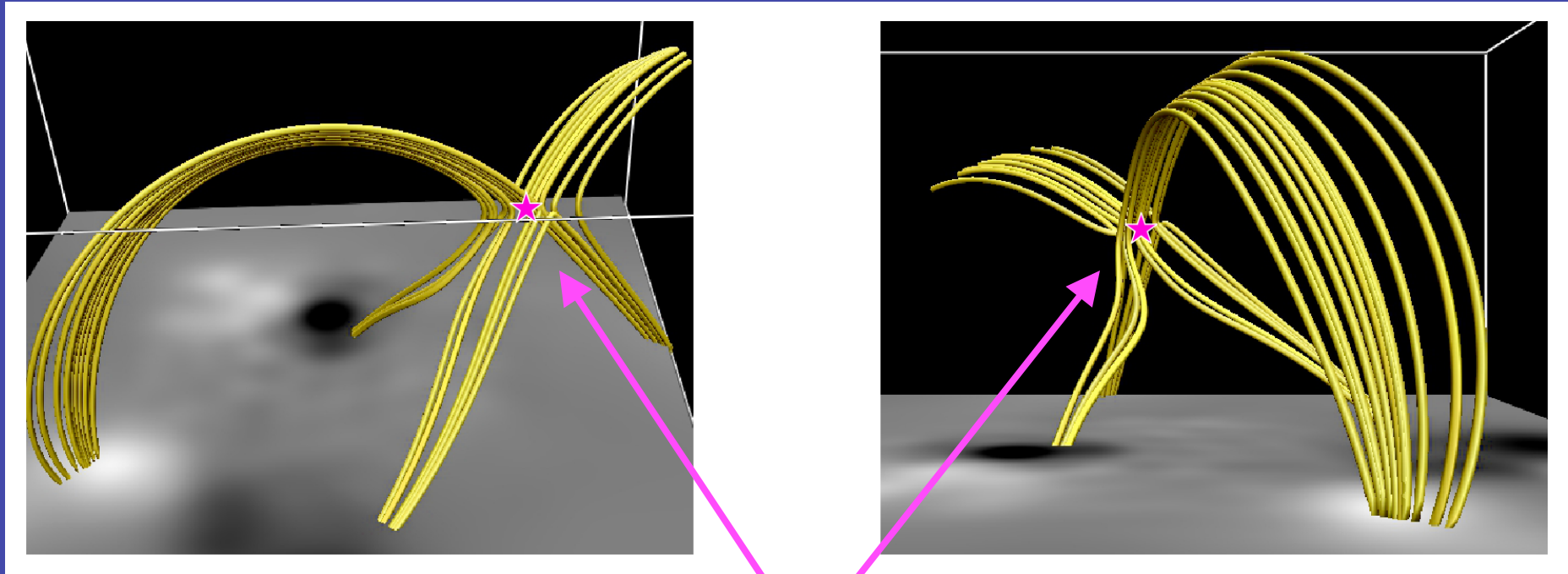


Null points

Prominence –  
coronal rain



# Magnetic geometry: detailed view near null (Xudong Sun)

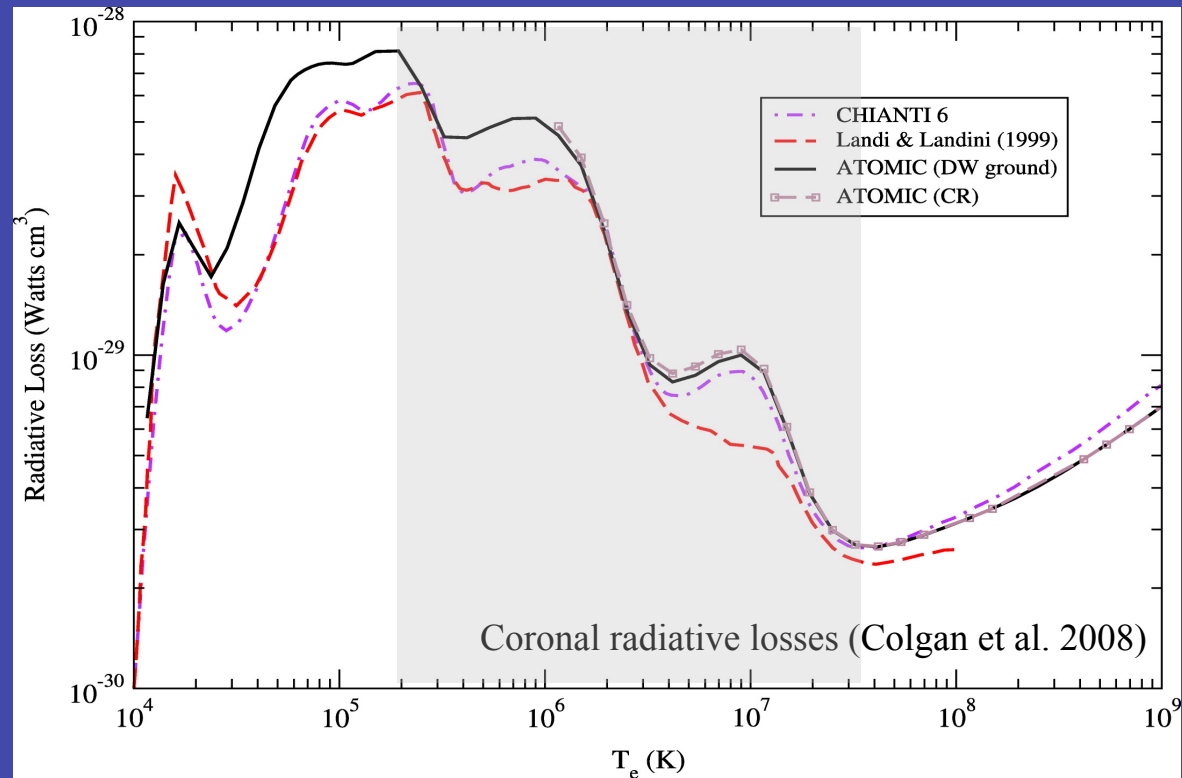
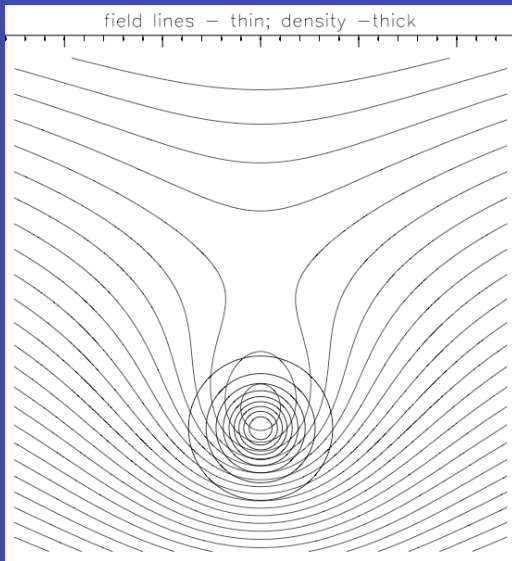


Null point

### 3. Why null points – (1) Coronal Condensation (thermal non-equilibrium, e.g., Karpen, Xia, Keppens, et al.; magnetic fields are crucial, e.g., Low+2012a,b)

#### (1) Radiative loss instability

$L = n_e^2 \Phi(T)$  increases with  $n^2$ , decrease with  $T$  ( $10^5 - 10^7$  K)



#### (2) Magnetic fields (insulators): reduced cross-field thermal conduction

$$\kappa_{\parallel}(T) = 2 \times 10^{-6} T^{5/2} \text{ erg cm}^{-1} \text{ s}^{-1} \text{ K}^{-1}$$

$$\kappa_{\perp} = \kappa_{\parallel} / (\omega_e \tau_e)^2$$

$$\omega_e \tau_e \gg 1$$

$\omega_e$ : electron gyro frequency;

$\tau_e$ : Coulomb collision time;

$(\omega_e \tau_e)^2 = 10^{10}$  for solar corona

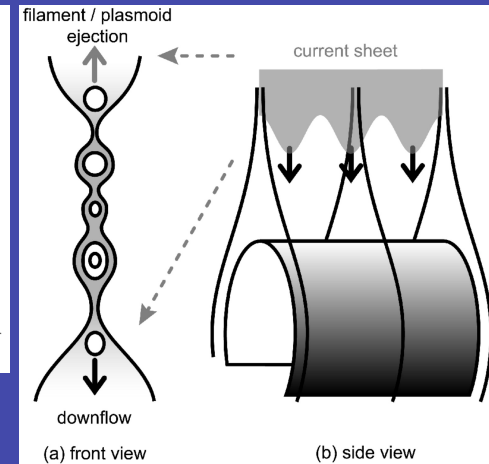
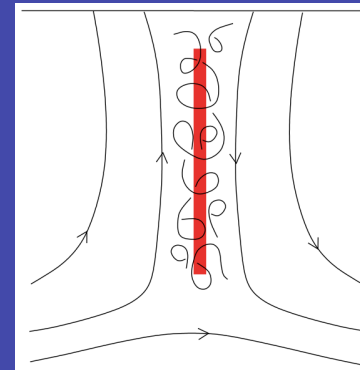


# Why null points – (2) plasma dynamics

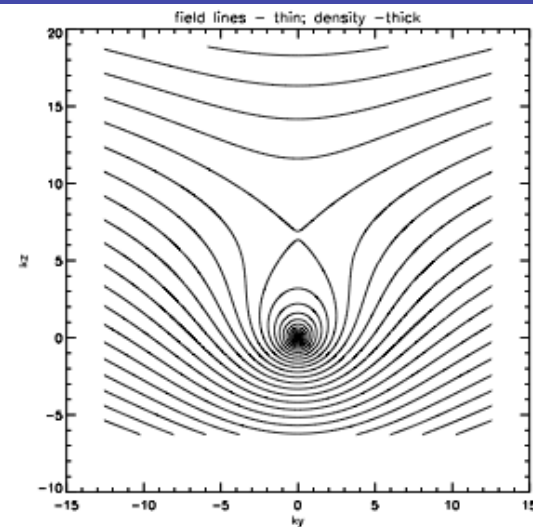
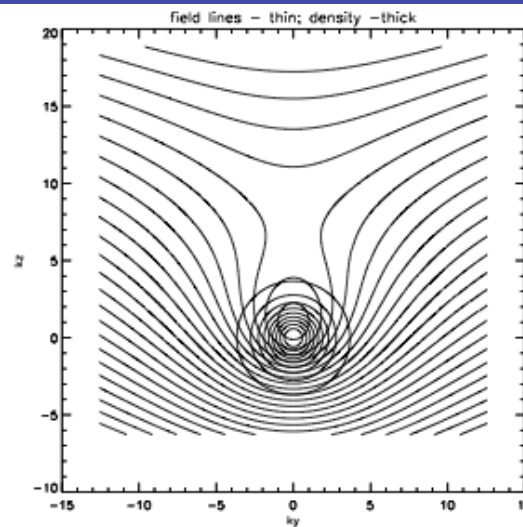
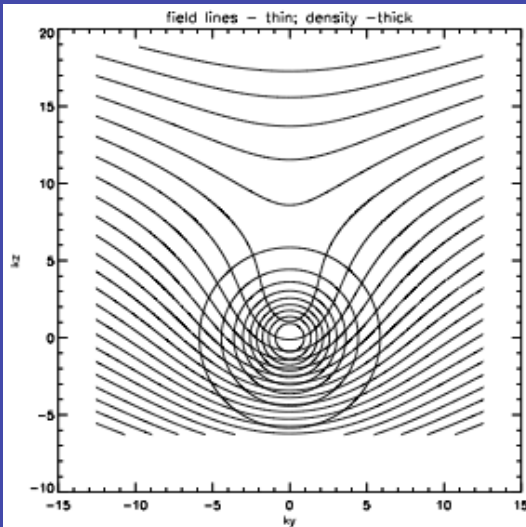
1) Tangled magnetic fields in current sheets (van Ballegooijen & Cranmer 2010) – suppress thermal conduction (Chandran & Cowley 1998), meandering flows

2) Current sheets: weak B-field, high-beta? Thus fluid/turbulent-like flow pattern (cf. SADs, McKenzie, Savage et al.).

3) Null points result from condensation/reconnection (Low, B.C.+2012a,b); neutrals; plasmoids (Guo, Bhattacharjee, Huang 2013) by reconnection as seed condensation; pileup at “Y” (Guo+2014); entropy mode (Murawski+2011)...



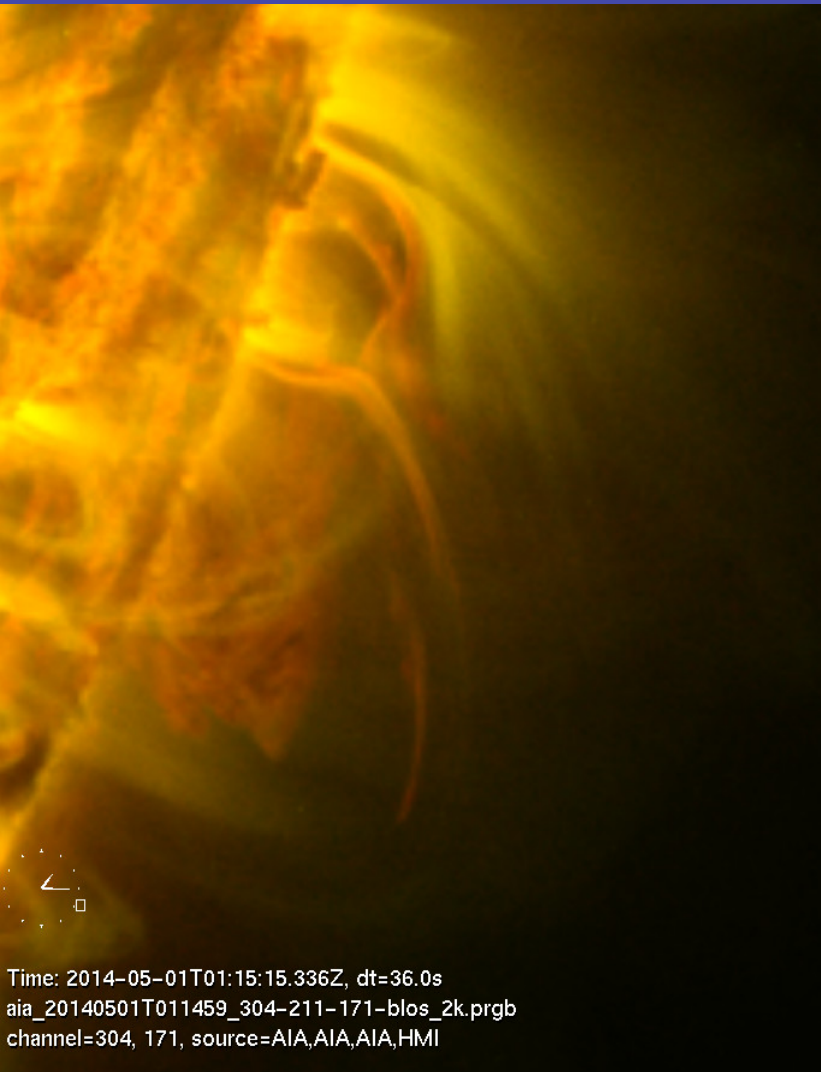
Asai et al. 2003



# Another example: Fan-spine geometry and coronal rain (eruption analyzed by Kathy Reeves et al. 2015 ApJ)

AIA171 Å (Yellow)/304 Å (Red)

IRIS 1330 Å SJI

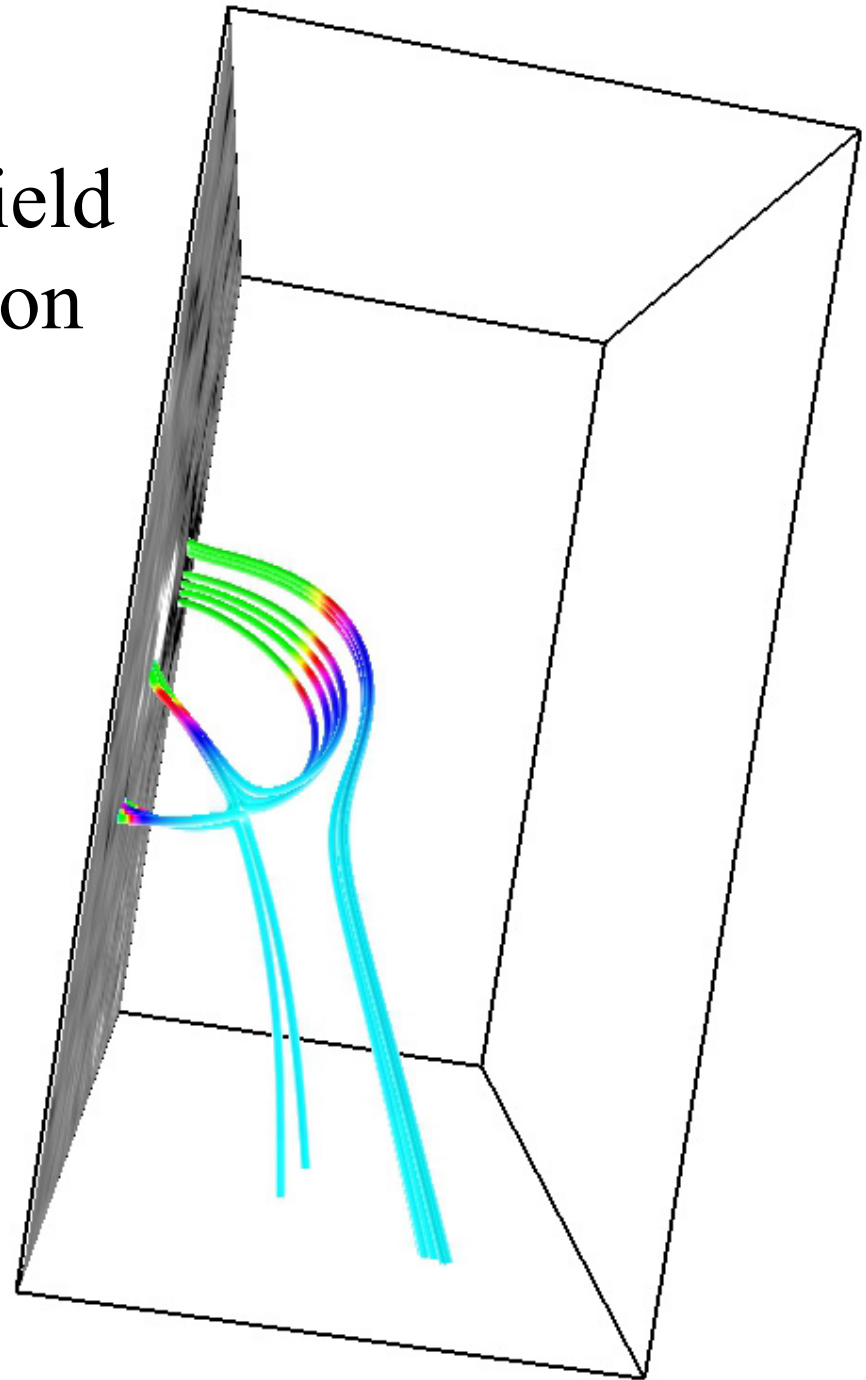




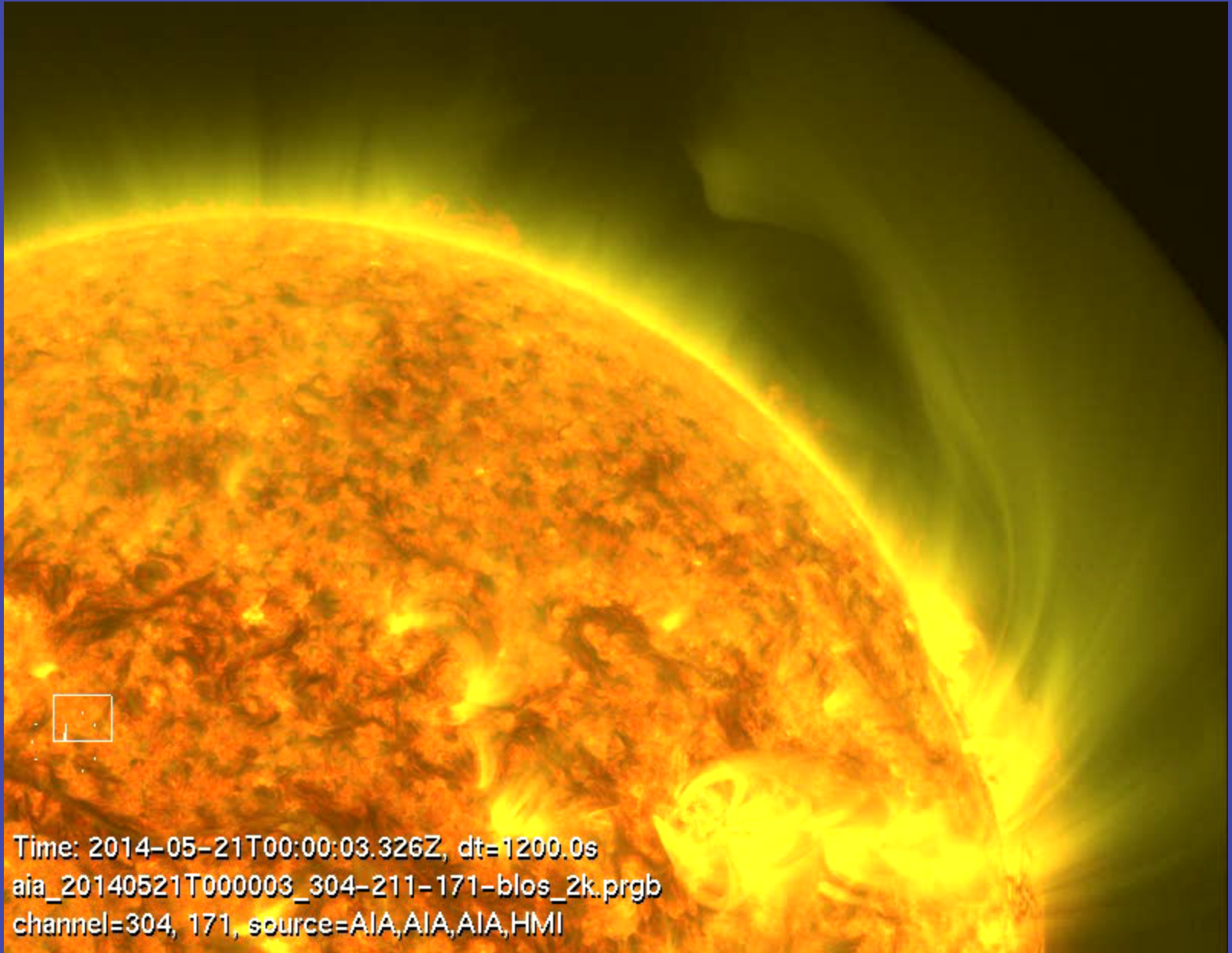
2014/05/01 01:15:00.920



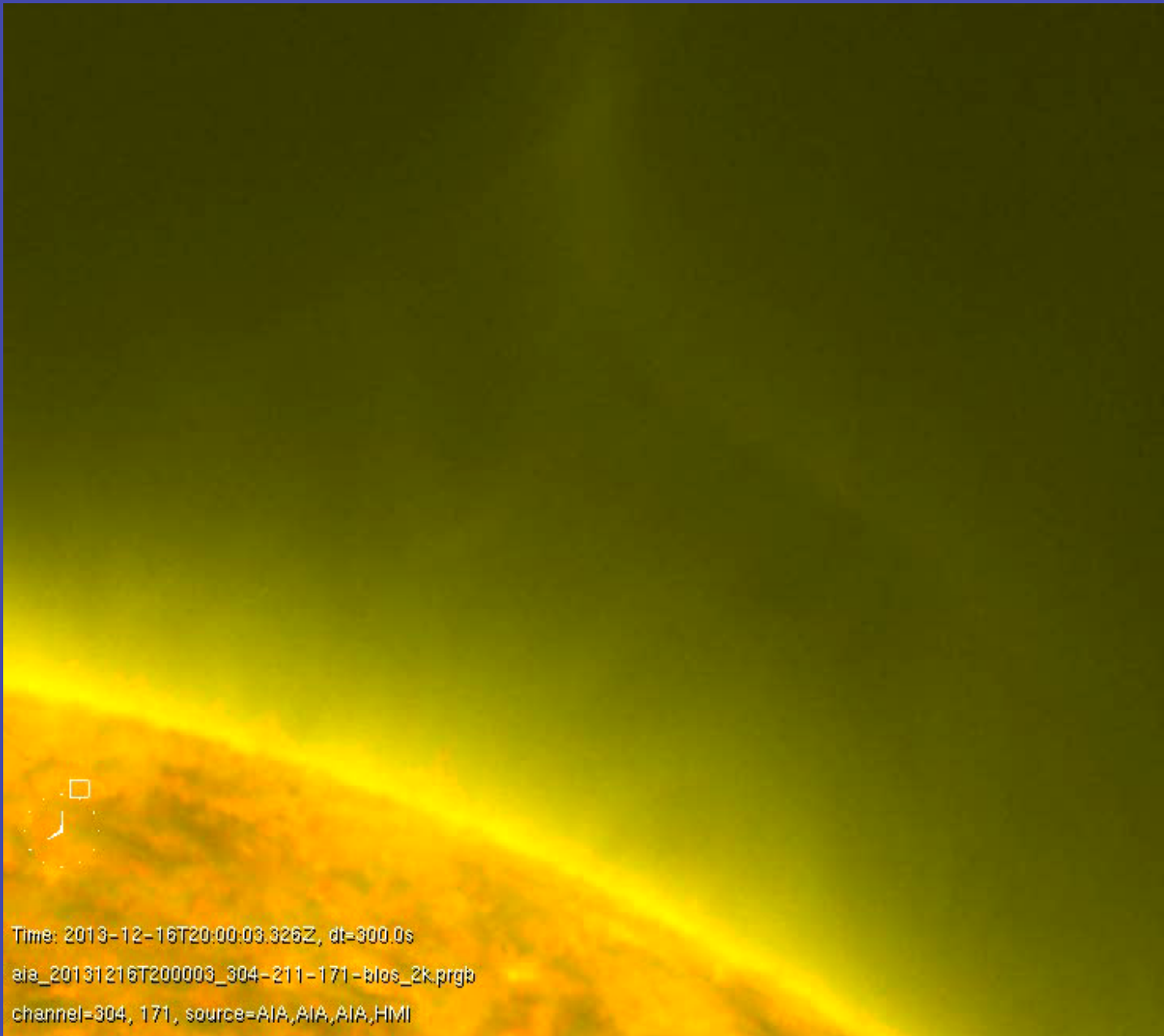
# Potential field extrapolation (Xudong Sun)



# Another example: Preferential occurrence of coronal rain at dips/cusps



# More examples: prominence threads turning into underlying coronal rain



# 4. Summary

A new paradigm: Hybrid Prominence – Coronal Rain complex (coronal spider?)

- 1) Prominence, irregular flow patterns near null points/current sheets: tangled field, enhanced cooling condensation, high-beta?
- 2) Underlying loops in arcade: coronal rain sliding down loops, stronger magnetic field, low-beta.
- 3) Future outlook:  
DKIST/VBI, VTF, and ViSP for POS/Doppler velocity, DL-/Cryo-NIRSP for magnetic field.  
ALMA for prominence (chromospheric) material – (see Nicolas Labrosse's talk this afternoon)

