

# FIDO Tool, Data Handling And Instrument Performance Calculators

# Science use case design

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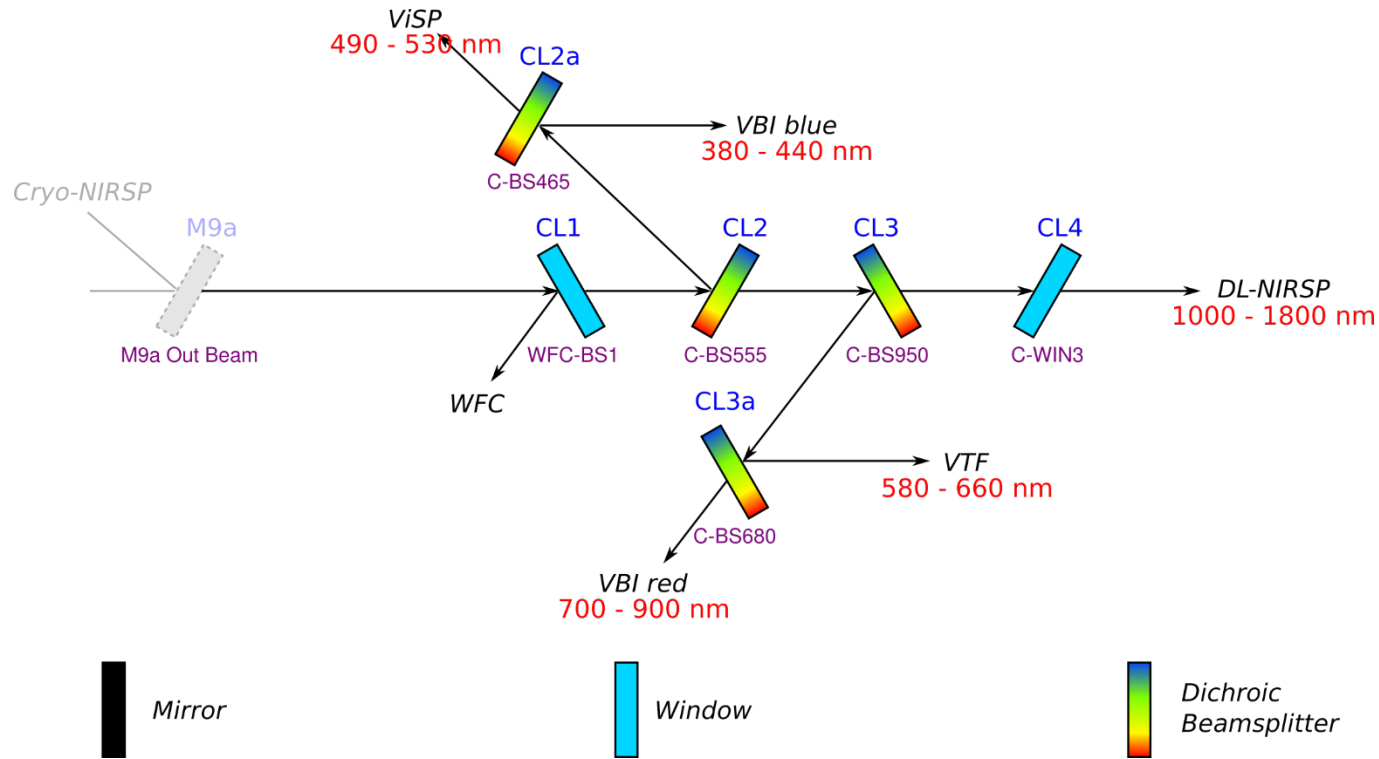
- DKIST instruments are complex, diverse, and flexible to support a very broad science portfolio.
- Users must understand and/or make decisions about:
  1. Telescope field-of-view, mosaicking, field sampling, coronagraphy, etc.
  2. Coudé table orientation (orientation of solar image on detector).
  3. **Spectral distribution of light to instruments.**
  4. Spectral/imaging/polarimetric capabilities of facility instruments.
  5. **Instrument parameters.**
  6. **Cadences and frame rates (limited by DHS capabilities).**

# FIDO Tool – What is it for?

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- Verifies whether the proposed wavelength distribution to the instruments is compliant with what FIDO can do.
  - Directive: “Either all light or no light” to an instrument.
  - There is no sharing of the same wavelength range between instruments; there are no grey beamsplitters (e.g. 50/50).
- Verifies whether the DHS can handle the proposed data rates.
  - DKIST is built for very high data rates, but still has limits on rate and volume.
  - First-order estimates of rate/volume calculated by FIDO tool.
  - Detailed rates/volumes are calculated by Instrument Performance Calculators.

# 3. Spectral distribution of light to instruments

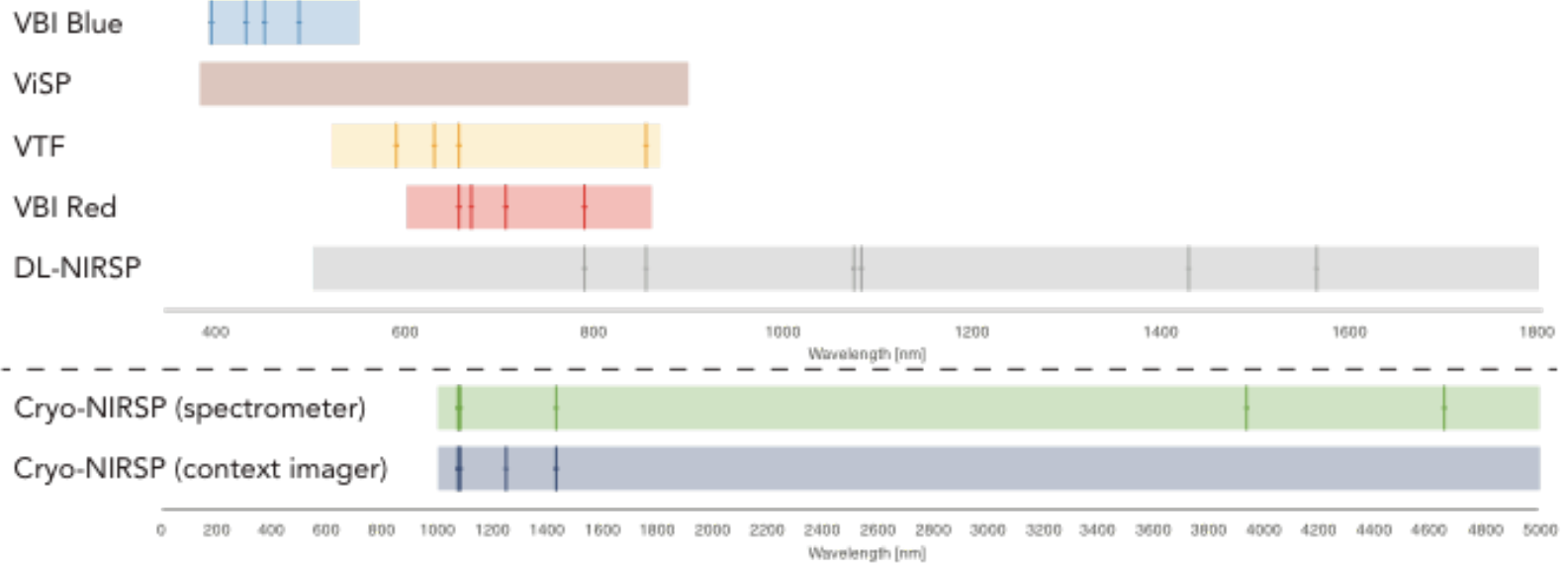


Dichroics are long-pass filters (transmit longer wavelengths)

FIDO: Many different ways to combine instruments → flexibility (diagnostic power) = data (calibration) challenges



# DKIST First Light Instrument Filters



VBI Blue	ViSP	VTF	VBI Red	DL-NIRSP	Cryo-NIRSP	Cryo Context
Ca II K 393.327nm	Access to entire spectral range between 380-900 nm	Na D 589.6nm	H-alpha 656.282nm	Fe XI 789nm	Fe XIII 1074.7nm	Fe XIII 1074.7nm
G-band 430.52nm		Fe I 630.25nm	Continuum 668.423nm	Ca II 854.2nm	Fe XIII 1079.7nm	He I 1083nm
Continuum 450.287nm		H-alpha 656.3nm	Ti O 705.839nm	Fe XIII 1074.7nm	He I 1083 nm	J Band 1250nm
H-beta 486.1nm		Ca II 854.2nm	Fe XI 789.186nm	He I 1083nm	Si X 1430nm	Si IX 1430nm
				Si X 1430nm	Si IX 3935 nm	
				Fe I 1565nm	CO 4651nm	

This table is meant to give an idea of the capabilities of the DKIST first light instrument suite. It cannot capture the large trade space that is provided by the flexibility of the instruments. For more information, visit <http://dkist.nso.edu/CSP/instruments>

Visible light cameras for instruments are provided by a UK consortium.



# FIDO Tool – How does it work?

INPUTS:  
Wavelengths  
and modes for  
each desired  
instrument

Priorities can  
also be used for  
optimization  
(instrument  
selection)

The screenshot shows the FIDO Tool configuration window for DKIST. The window is titled "DKIST" and contains several sections for instrument configuration:

- VBI:** Configuration for Camera 1 and Camera 2. Camera 1 has checkboxes for 393 nm, 430 nm, 450 nm, and 486 nm. Camera 2 has checkboxes for 656 nm, 668 nm, 705 nm, and 789 nm. Each camera has a "ReconstructedImage" dropdown menu and a "Priority" input field set to 1.
- VTF:** Configuration for Cameras. Checkboxes for 525 nm, 630 nm, 656 nm, and 854 nm. A dropdown menu is set to "UnbinnedPolarimetric Mode". Each camera has a "Priority" input field set to 1.
- VISP:** Configuration for Camera 1, Camera 2, and Camera 3. Each camera has a "wavelength [nm]" input field (Camera 1 is set to 700) and a "Priority" input field set to 1.
- DL-NIRSP:** Configuration for Camera 1, Camera 2, and Camera 3. Each camera has radio buttons for different wavelengths (e.g., 789 nm, 854 nm for Camera 1) and a "Priority" input field set to 1.

At the bottom of the window, there is a "FastCadence (low pol. precision)" dropdown menu and an "Analyze Configuration" button.

# FIDO aka Beamsplitter Configuration Tool

## Configuration Outputs

```
=====
The following Coude Optics configurations [CL2, CL2a, CL3, CL3a]
deliver the highest ranking:
=====
```

```
[BS_465,BS_950,BS_680,BS_555]; [BS_465,MI_001,BS_680,BS_555];
[BS_465,MI_002,BS_680,BS_555];
```

```
::VBI1:: waverange [nm]: [380,440]
```

```
-----> Max. Data Rate (successful diagnostics): 960 MB/s
success: 393; 430;
```

```
fail:
```

```
::VBI2:: waverange [nm]: [580,660]
```

```
-----> Max. Data Rate (successful diagnostics): 960 MB/s
success: 656;
```

```
fail:
```

```
::VTF:: waverange [nm]: [490,530]
```

```
-----> Max. Data Rate (successful diagnostics): 2880 MB/s
success: 525;
```

```
fail:
```

```
::ViSP1:: Channel receives no light!
```

```
success:
```

```
fail: 700;
```

```
::DLN1:: waverange [nm]: [700,1800]
```

```
-----> Max. Data Rate (successful diagnostics): 67 MB/s
success: 854;
```

```
fail:
```

```
::DLN2:: waverange [nm]: [700,1800]
```

```
-----> Max. Data Rate (successful diagnostics): 67 MB/s
success: 1083;
```

```
fail:
```

```
::DLN3:: waverange [nm]: [700,1800]
```

```
-----> Max. Data Rate (successful diagnostics): 67 MB/s
success: 1565;
```

```
fail:
```

```
Aggregate Bandwidth (max. 3500-4000 MB/s): 5001 MB/s
```

```
=====
```

## 5. Instrument parameters including domain coverage and SNR

- In most cases, must be calculated using the Instrument Performance Calculators. Some examples given in summary docs.
- IPCs available online ([dkist.nso.edu/CSP](http://dkist.nso.edu/CSP)) and will be discussed in a few minutes.

## 6. Data rate/volume limits

- DKIST is built for very high data rates, but still has limits on rate and volume.
- Rates/volumes are calculated by Instrument Performance Calculators for your detailed use case.
- First-order estimates of rate/volume provided the FIDO beam splitter calculator.



# Capabilities of the Data Handling System (DHS)

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- The DHS provides **five** dedicated “camera lines” of 960 MiB/s bandwidth each.
- Camera lines can be **shared** among multiple cameras within 960 MiB/s limit (e.g. when binned, or not running at the highest frame rate)
- Max cumulative bandwidth is 4800 MiB/s; however, depending on sharing, practical max is 3500 to 4000 MiB/s
- Data acquisition at the max possible rate might have to be limited in duration due to data volume issues – typically daily data volumes of 5-30 TB

# Maximum Instrument Detector Rates

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**VBI:** 2x (4096 x 4096) [Andor Balor].  
Max rates: 2 x 960 MiB/s

**VTF:** 3x (4096 x 4096) [Andor Balor].  
Max rates: 3 x 960 MiB/s

**ViSP:** 3x (2560 x 2160) [Andor Zyla 5.5].  
Max rates: 3 x 433 MiB/s

**DL-NIRSP:** 1 x (4096 x 4096) [Andor Balor] + 2 x (2048 x 2048) [H2RG]  
Max rates: 1 x 960 MiB/s; 2 x 240 MiB/s

**Cryo-NISRP:** 2 x (2048x2048) [H2RG]  
Max rates: 2 x 80 MiB/s

# FIDO aka Beamsplitter Configuration Tool

## Data Rate Outputs

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The following Coude Optics configurations [CL2, CL2a, CL3, CL3a] deliver the high
ranking:
=====
```

```
[BS_465,BS_950,BS_680,BS_555]; [BS_465,MI_001,BS_680,BS_555];
[BS_465,MI_002,BS_680,BS_555];
```

```
::VBI1:: waverange [nm]: [380,440]
-----> Max. Data Rate (successful diagnostics): 960 MB/s
      success: 393; 430;
      fail:
::VBI2:: waverange [nm]: [580,660]
-----> Max. Data Rate (successful diagnostics): 960 MB/s
      success: 656;
      fail:
::VTF:: waverange [nm]: [490,530]
-----> Max. Data Rate (successful diagnostics): 2880 MB/s
      success: 525;
      fail:
::ViSP1:: Channel receives no light!
      success:
      fail: 700;
::DLN1:: waverange [nm]: [700,1800]
-----> Max. Data Rate (successful diagnostics): 67 MB/s
      success: 854;
      fail:
::DLN2:: waverange [nm]: [700,1800]
-----> Max. Data Rate (successful diagnostics): 67 MB/s
      success: 1083;
      fail:
::DLN3:: waverange [nm]: [700,1800]
-----> Max. Data Rate (successful diagnostics): 67 MB/s
      success: 1565;
      fail:
```

```
Aggregate Bandwidth (max. 3500-4000 MB/s): 5001 MB/s
=====
```

# Instrument Performance Calculators

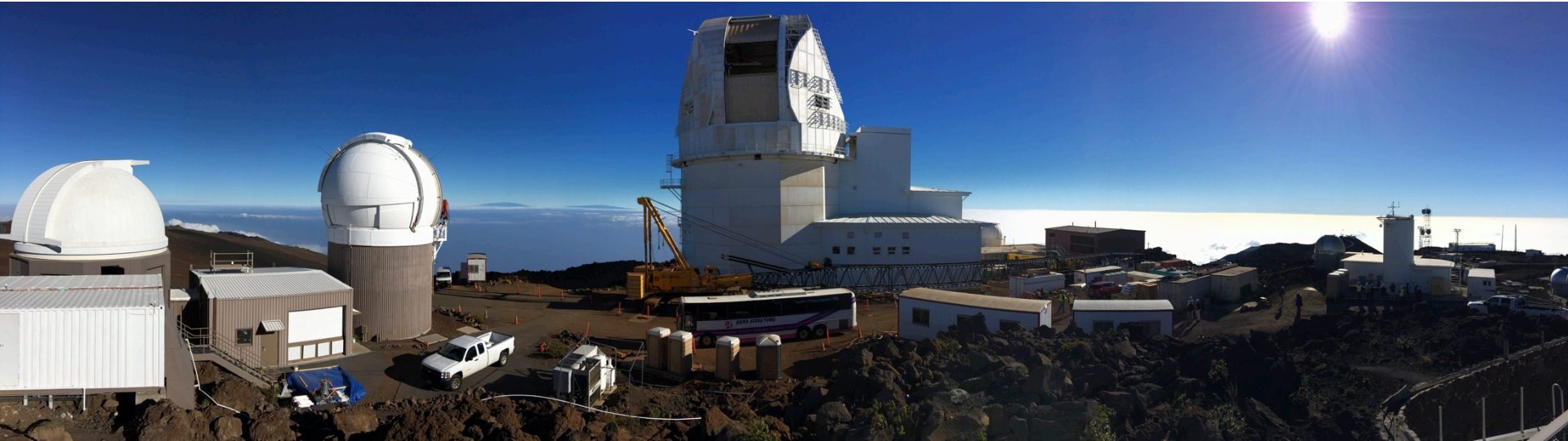
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- Set of tools (i.e. software programs/applications) to explore instrument capabilities – e.g. line selection, exposure times, SNR etc.
- A \*separate\* IPC needs to be run for each instrument. VBI and VTF are Java applications (1.9); ViSP and DL-NIRSP run in IDL (8+). (The Cryo-NIRSP IPC unfortunately is not yet ready for distribution)
- It is useful to run FIDO ahead of the IPCs, to check if the intended spectral distribution is feasible.

# Some examples of IPCs use

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# Thanks !



<http://dkist.nso.edu/>